

Phil Appleton

NASA Herschel Science Center

Project Scientist and Task Lead

3.5-m IR/Sub-mm Herschel Space Observatory

- Fourth "Cornerstone" mission of European Space Agency's "Horizon 2000" Program and launched in May 2009. It finished observing in April 2013
- Many countries, including the USA, have participated in its design and implementation
- ESA's European Space and Astronomy Center (ESAC) in Madrid, Spain, hosts the Herschel Science Center (HSC).
 Caltech hosts the NASA Herschel Science Center (NHSC)
- Both HSC and NHSC will continue to support observers and archive users for the next few years during a Postops phase.

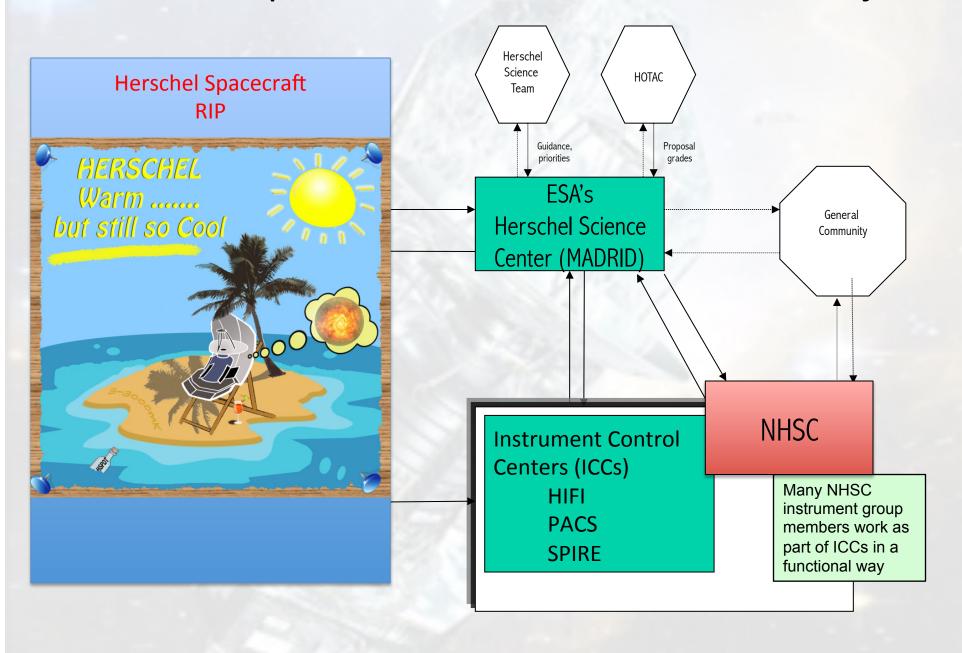
Objective and Scope of the NASA Herschel Science Center

The NHSC has been established at the Infrared Processing and Analysis Center (IPAC) at the California Institute of Technology under support from NASA.

NHSC began operations in 2001 and will continue to operate throughout the lifetime of the Herschel Mission, including the post-cryogen archival phase.

The primary mission of the NHSC is to support the US-based astronomical community in exploiting the Herschel science opportunity, at a level equivalent to that provided to the European Community by ESA's Herschel Science Centre.

Relationship between NHSC and Herschel Project





How to find us on the web



What the NHSC Does for US Astronomers?

- Provide resources and tools to help US Community exploit Herschel science (Inc. FREE ACCESS TO REMOTE COMPUTERS)
- Provide support and documentation for Herschel-specific software (HIPE, HSPOT + others)
- Advocate US user needs within Herschel Project
- Engage US community in Herschel Science
- Help develop the best possible Herschel Science Archive
- Administration of data analysis grants for US-based observers

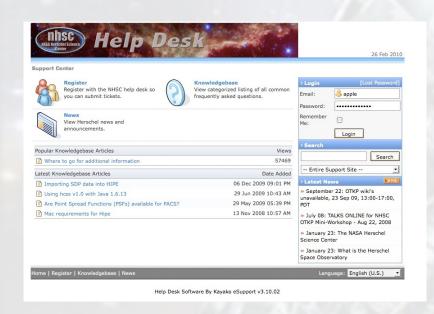
RESOURCES AND TOOLS of the NHSC

NHSC Resources	New User	Expert User	Comments	
NHSC Helpdesk NHSC Wiki Pages NHSC Newsletter	Y Y Y	Y Y Y	nhsc.ipac.caltech.edu/helpdesk nhscsci.ipac.caltech.edu nhscsci.ipac.caltech.edu/sc/index.php/ NewsLetters	
Intro Workshops	Y	N	Best for beginner users or new archival users	
Data School/Advanced DP workshop	N	Y	Users bring their own data to reduce using big machines	
Webinars	Y	Y	On-line topical discussion groups	
On-line Video and other web- tutorials	Y	Y	Step by Step guides	
NHSC In-person visits	N	Υ	Staff Availability	
Free Remote Access Computer Resources	N	Y	Requested through HELPDESK	

NHSC Helpdesk

FIRST STOP FOR US OBSERVERS IN NEED OF HERSCHEL HELP:

http://nhsc.ipac.caltech.edu/helpdesk



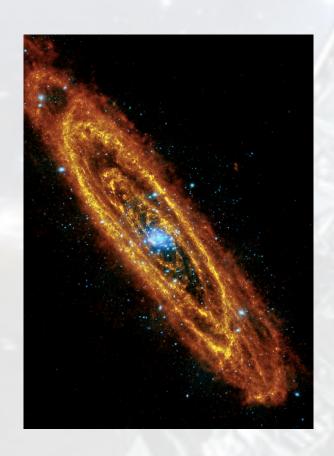
- Has been fully operational since Jan 2007
- NHSC Response:
 - Provide initial OST acknowledgement
 - Ticket forwarded immediately to expert who responds directly within 2 business days
 - In rare case that we are unable to assist,
 we open and follow-up with European help desk ticket.
- Average response time < 1 business day (often ~ 2-5 hrs)



Our first NHSC Data Processing Workshop in 2009

Visitors to NHSC were common during early operations and still provide a channel for help and advice with DP

NHSC Remote Computing Resources for USbased Observers



- FREE large-memory (64-128 GB) fast machines by remote access
- LATEST HIPE VERSION FULLY INSTALLED
- Large disk space to store and transfer to home institution
- Access to variety of different Herschel mappers for comparison of different mapping methods (currently available MADmap, Scanamorphos and Unimap, SANIPIC + more coming

REQUEST MACHINE VIA NHSC HELPDESK

Outreach Activities

AAS and Other Meeting Support

Help with Publicizing Exciting Results

Image or Press Release—coordinated with ESA if appropriate

Organizing Conferences and Science Workshops

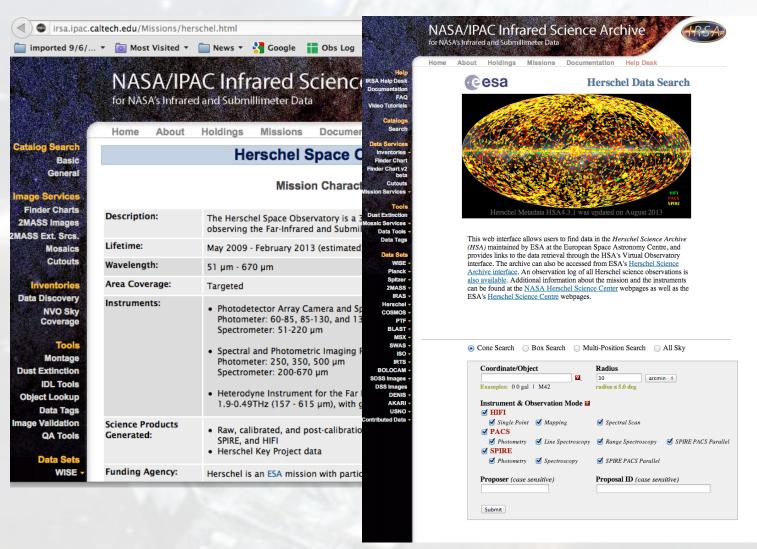
Team Wiki Pages

The NASA HERSCHEL SCIENCE CENTER ALSO SUPPORTS AAS MEETINGS



Herschel images were presented on the giant electronic "Hyper-board" by Robert Hurt (IPAC) as part of a demo on astronomical images in the IR

In Post-Operations NHSC has a strong emphasis on multiplying the value of the Herschel Science Archive —we have developed interfaces with the NASA InfraRed Science Archive (IRSA) at IPAC



HERSCHEL THE OBSERVATORY

What is special about Herschel Data?

- Herschel is a large telescope
 - 3.5 m diameter
 - collecting area and resolution
- Exploring new spectral window
 - 55-672 μm bridging the far infrared & submillimetre – the 'cool' universe
- Novel instruments
 - wide area mapping in 6 'colours'
 - imaging spectroscopy
 - very high resolution heterodyne spectroscopy
- Herschel archive is:
 - designed to capture the most interesting far-IR/sub-mm astrophysics:
 - galaxy evolution over cosmic time
 - star formation near and far
 - ISM physics/chemistry
 - birth of solar systems
 - many things yet to be discovered in archive!

Launched to L2 on 14 May 2009

End of Helium 29 April 2013

Made over 35,000 scientific observations (25,000 hrs)

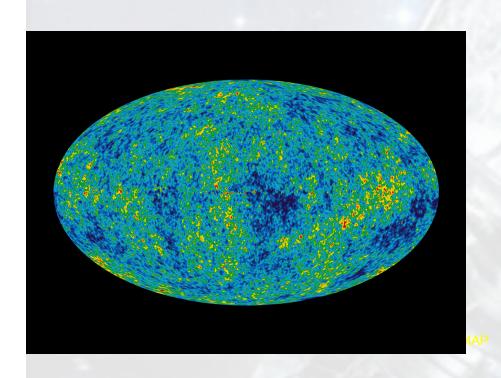
Archive is the Herschel Legacy

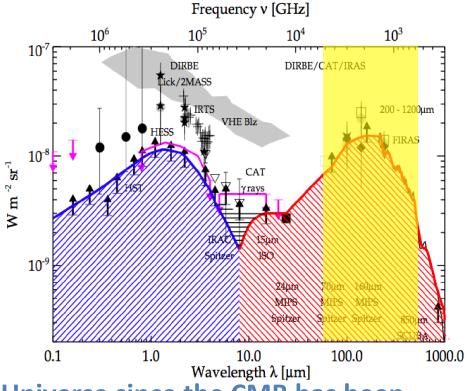
-This is why you are here!



Science Background-I

(Galaxies SEDs across cosmic time)

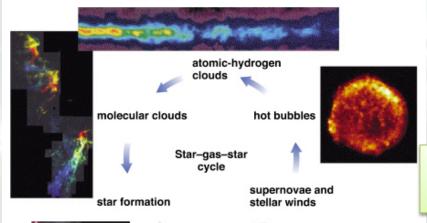




- Half of the energy created in the Universe since the CMB has been reprocessed into the IR
- Herschel covers the IR peak and pushes into the submillimetre
- Herschel is able to resolve a large fraction of galaxies contributing to IR peak

Science Background -II The Molecular Trail to the formation of Life

Herschel is helping to uncover the inter-relationship between the life-cycle of grains and chemistry that created life



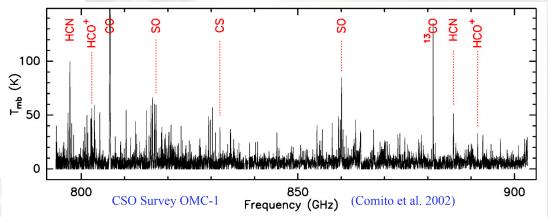
Hydrogen and water chemistry and the relationship to dust and dust recycling in the life of stars.

It is exploring the properties of multi-phase gas and reprocessing in star formation regions

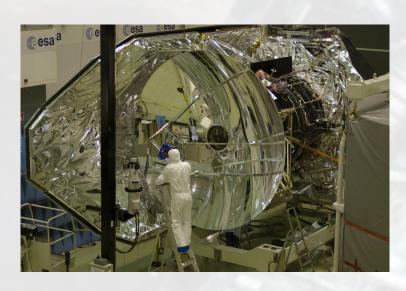


stellar burning/ heavy-element formation

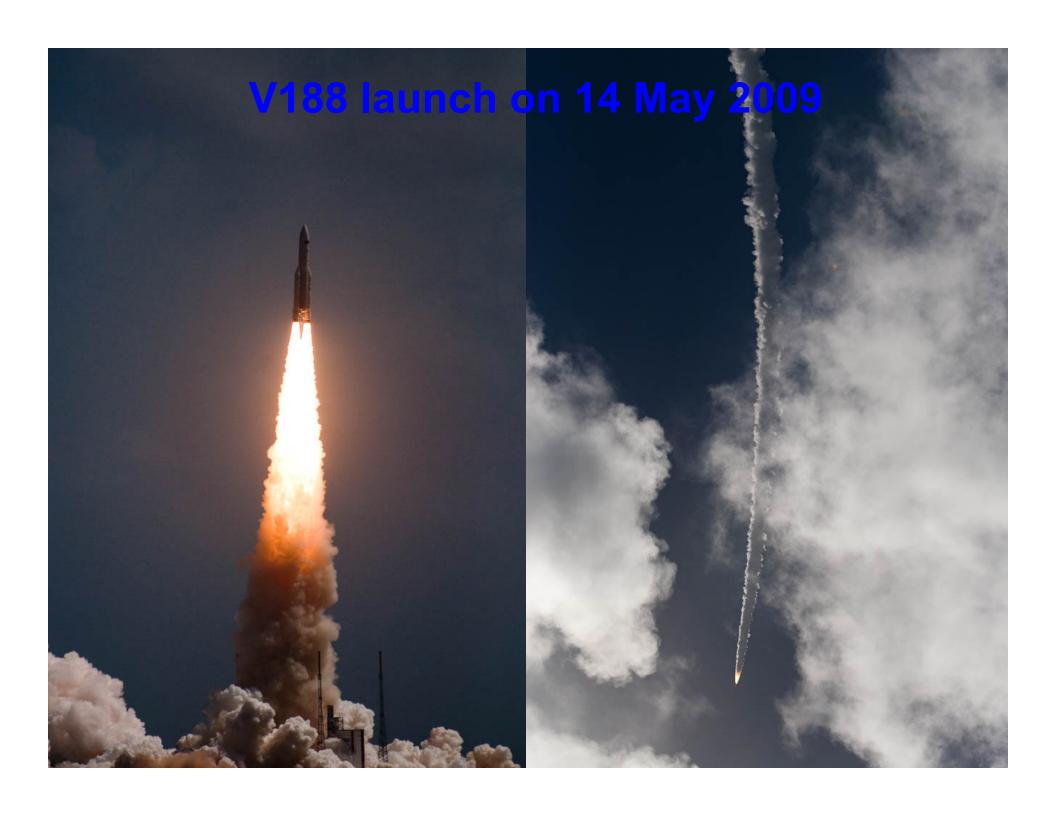
Copyright @ Addison Wesley



Herschel and Planck on the launch-pad! Last view of Herschel on ground before launch!

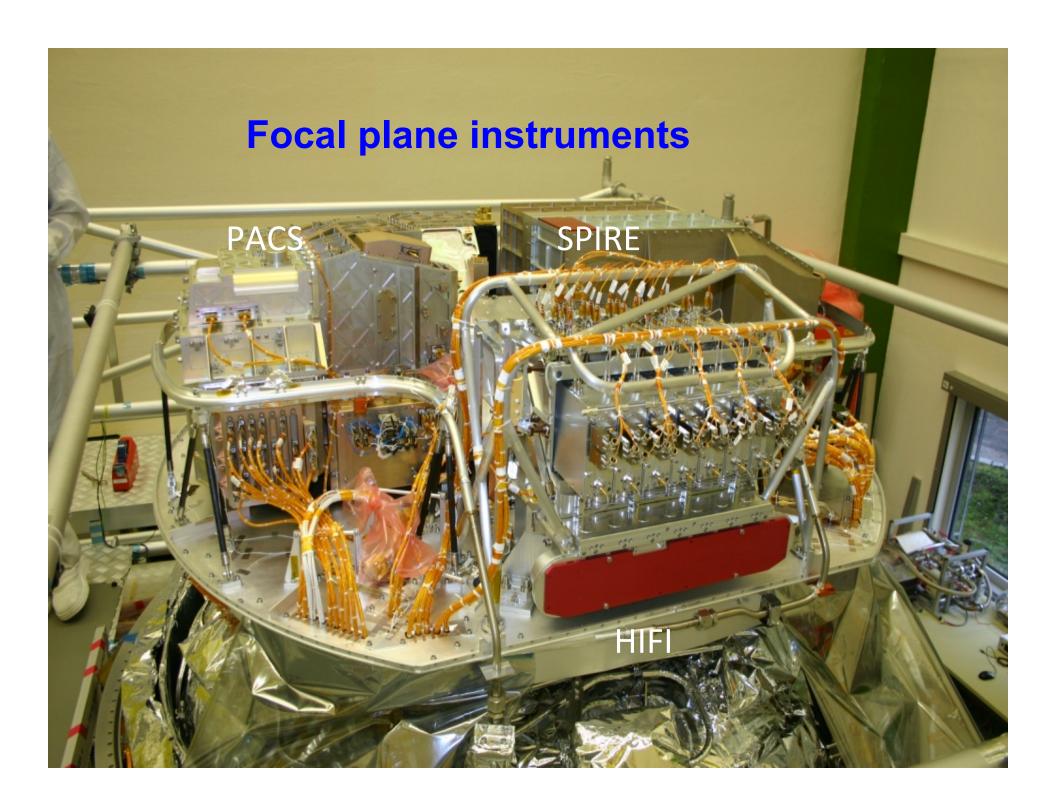






Herschel's view of Planck-Sylda above Africa en route for L2



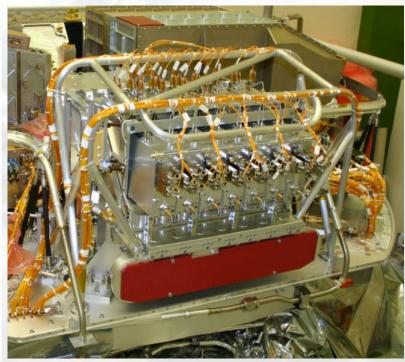


HIFL

HIFI – Heterodyne Instrument for the Far Infrared

- PI: Frank Helmich, SRON, Groningen, The Netherlands
- Co-Pls: Tom Phillips, Caltech, USA; Jürgen Stutzki, U Köln, Germany;
 Emmanuel Caux, CESR, France; and Thijs de Graauw, ALMA
- Very high resolution spectroscopy over 480-1250 and 1410-1910 GHz
- SIS and HEB mixers, auto-correlator and AOS spectrometers





SPIRE

SPIRE – Spectral and Photometric Imaging Receiver

- PI: Matt Griffin, U Cardiff, Cardiff, United Kingdom
- Co-PI: Laurent Vigroux, CEA, Saclay, France
- Imaging photometry and spectro-photometry/-scopy over 194-672 μm
- 3 bolometer arrays for photometry, 2 bolometer arrays for spectroscopy













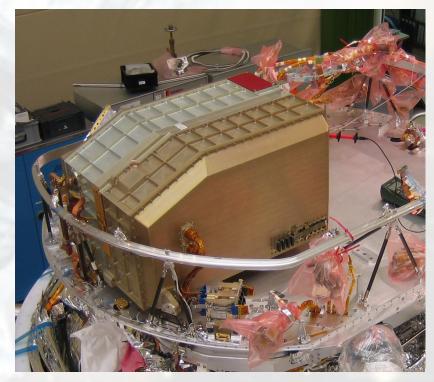


3-band imaging photometer

- 250, 350, 500 μm (simultaneous)
- $-\lambda/\Delta\lambda \sim 3$
- 4 x 8 arcminute field of view
- Diffraction limited beams (17, 24, 35")

Imaging Fourier Transf Spectrom

- 194 324, 316 672 μm
- 2.6 arcminute field of view
- $\Delta \sigma$ = 0.04 2 cm⁻¹ ($\lambda/\Delta\lambda \sim$ 20 1000 at 250 μ m)



PACS

PACS – Photodetector Array Camera and Spectrometer

- PI: Albrecht Poglitsch, MPE, Garching, Germany
- Co-PI: Christoffel Waelkens, KU Leuven, Belgium
- Imaging photometry and spectroscopy over 55-210 μm
- 2 bolometer arrays for photometry, 2 Ge:Ga arrays for spectroscopy



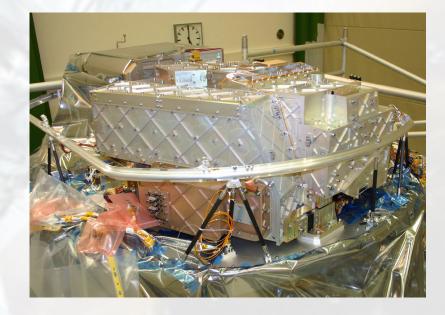
3-band imaging photometer

$\lambda (\mu \mathbf{m})$	70	100	160
FWHM (arcsec)	6	8	12
$\lambda/\Delta\lambda$	2.5	2.8	2.1

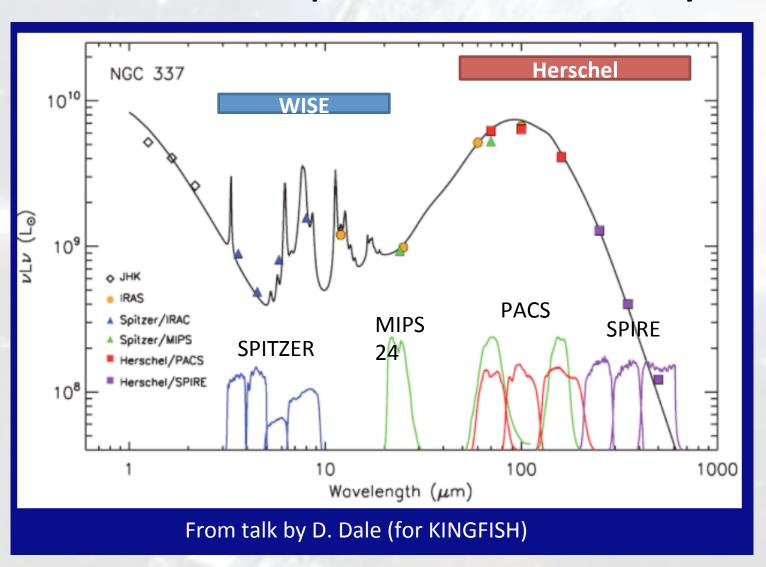
- Simultaneous obs at 70/100 & 160 μm
- 3.5 x 1.75 arcmin fully sampled FOV

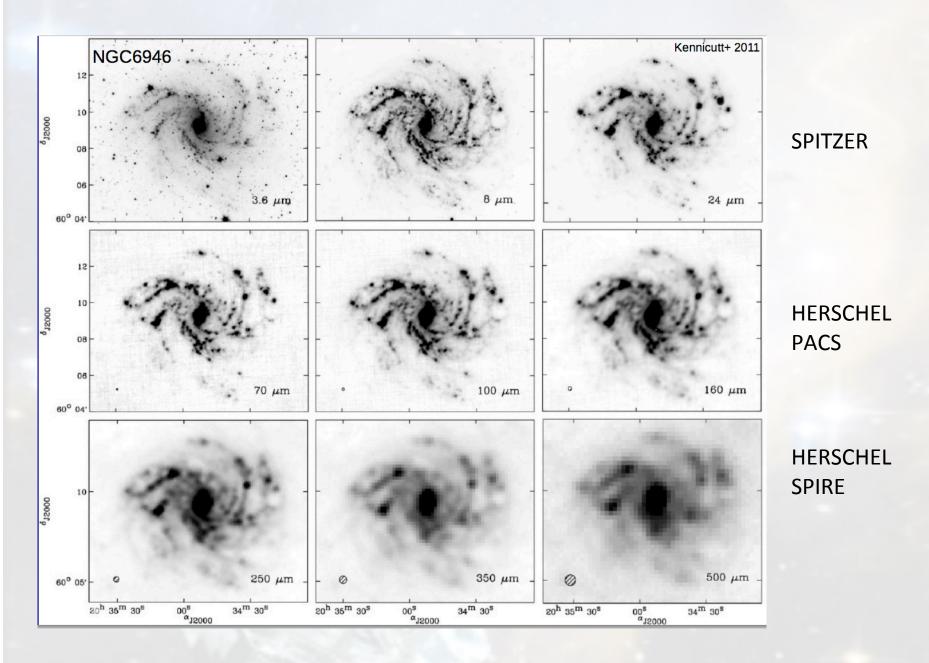
Imaging line grating spectrometer

- FOV (arcmin) fully sampled 0.8 x 0.8
- λ order 1,2,3 102-210, 71-98, 55-73 μm
- $\lambda/\Delta\lambda$ = 1500-4000



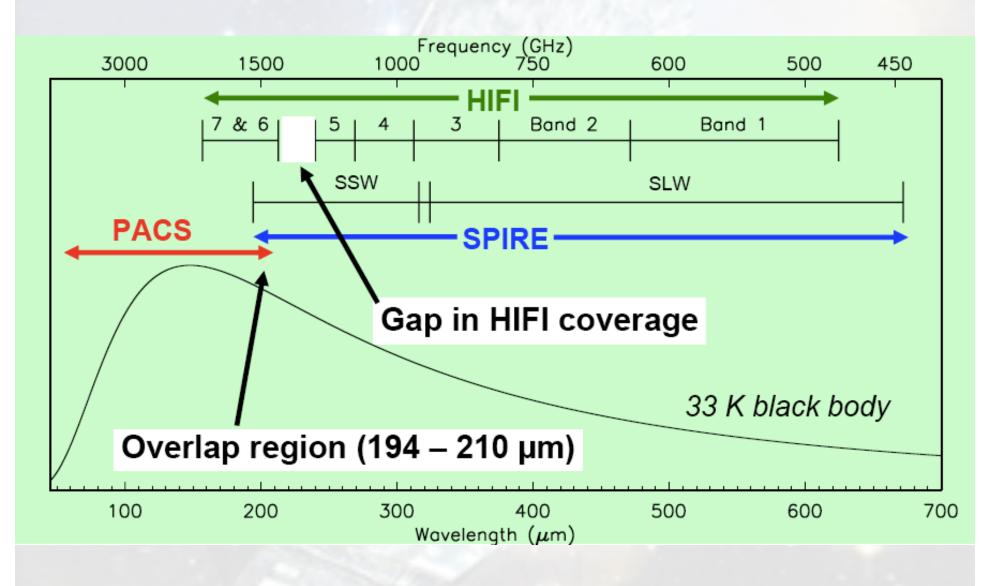
Complementarity of Herschel and WISE/Spitzer Photometry

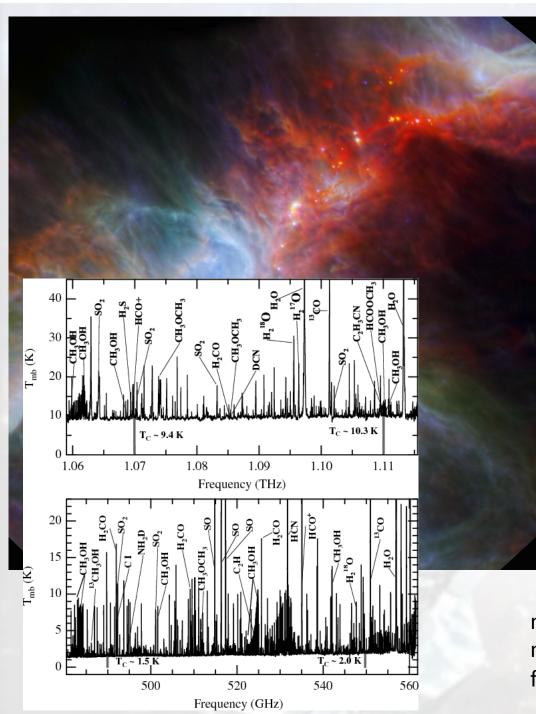




KINGFISH OPEN TIME KEY PROGRAM -From talk by Danny Dale (2012)

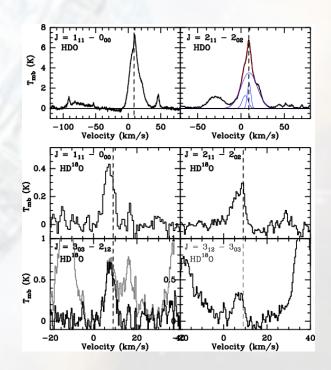
Herschel's Three Spectrometers





Example of Power of HIFI (from HEXOS Project; Bergin et al. 2010)

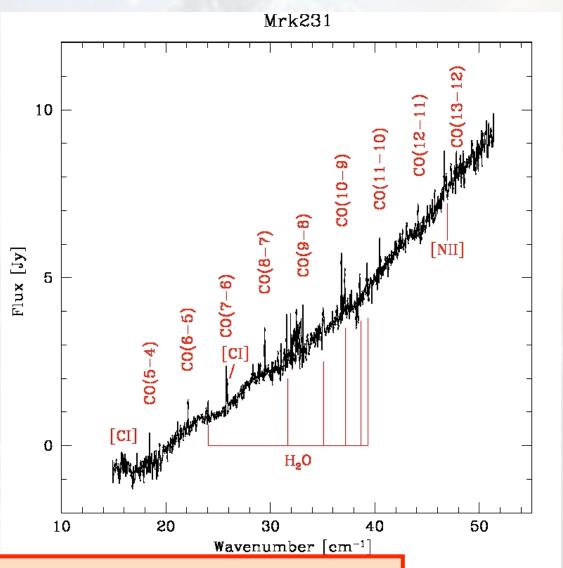
Spectra of Orion KL object



many species! many lines! fantastic spectral resolution!

SPIRE FT SPECTROMETER

Mk 231
P. van der Werf
HerCules
project

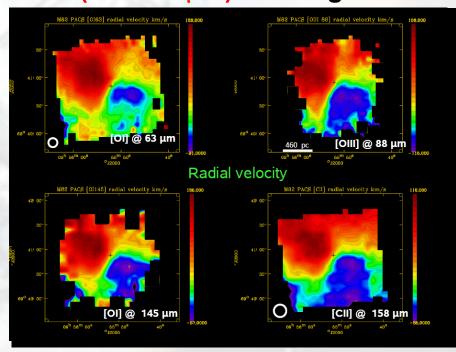


Entire spectral 194-672 μm coverage is observed in one go!

Integral Field Spectrometer 5x5 pixels 1 pixel = 9.7GeGa Two (25×16) arrays Stressed and unstressed

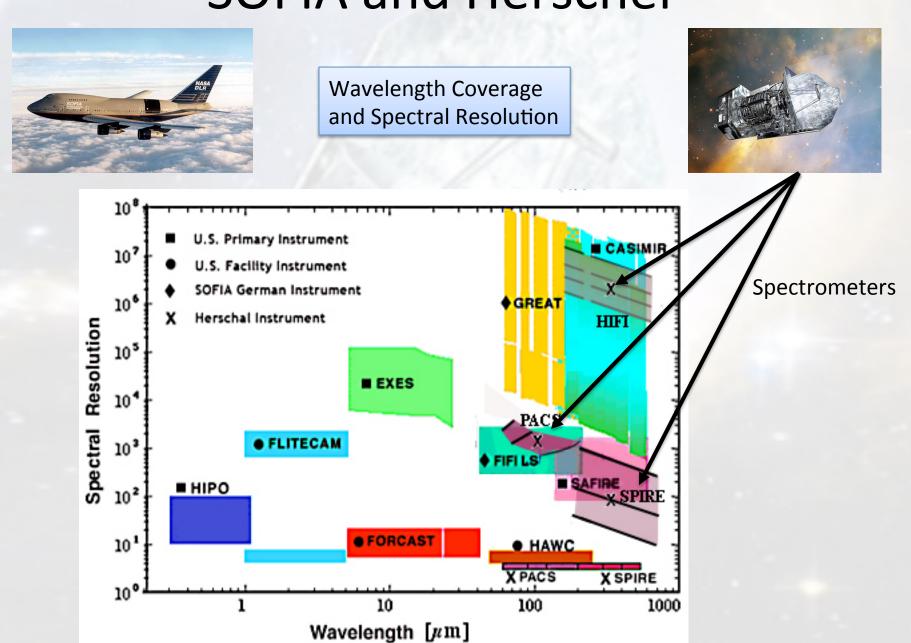
PACS Spectrometer

Simultaneous Blue (55-98 μm) Red (102-210μm) coverage

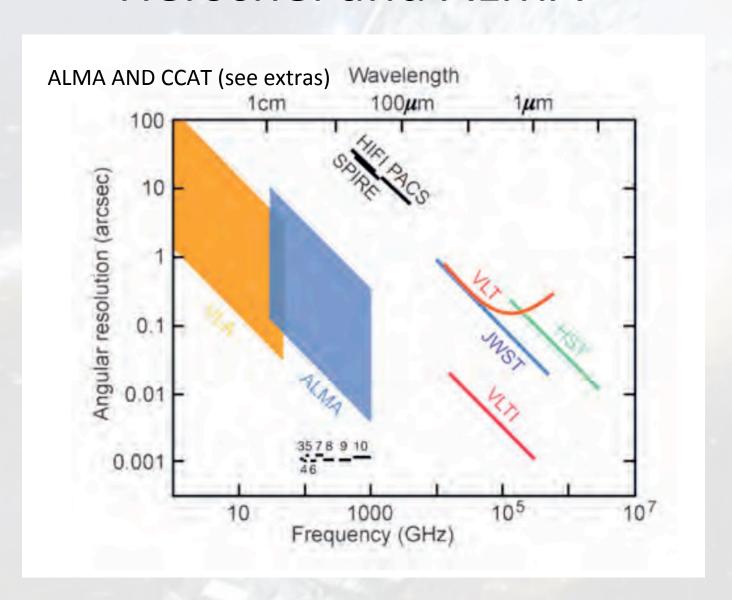


3-D Spectral Mapping of M82 (SHININGS KP)

SOFIA and Herschel



Herschel and ALMA



GT and Open Time Key Programs

First Year of Operations: Call for Large Projects

LARGE PROJECTS 21 chosen (7 US led)

Projects are providing "User Supplied Data Products"

http://herschel.esac.esa.int/UserReducedData.shtml

These are highly processed data created by the OTKP and GTKP projects and are held at the HSC and eventually will reside in the Herschel Science Archive

SOME KEY
HERSCHEL RESULTS

Imaging the COOLEST DUST in the MW:
Crowded Filaments of proto-stars

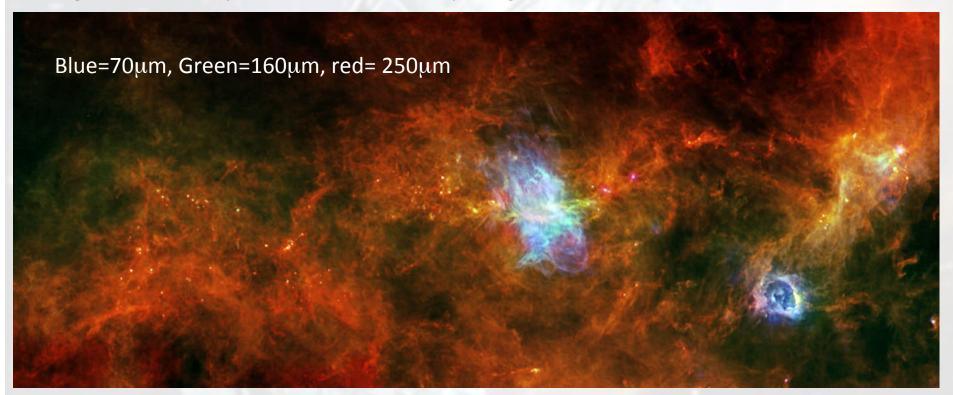
The Vela C region

Cool proto-stars and filaments

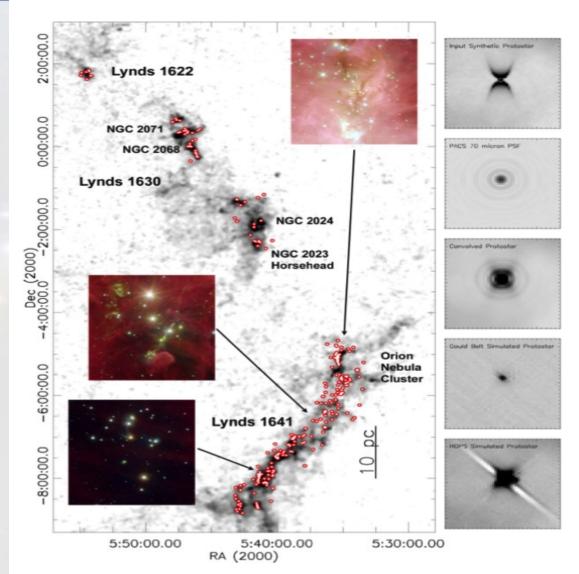
Part of the Vela complex, by ESA's Herschel space observatory.

The image demonstrates Herschel's ability to trace both high and low-mass stars at a range of evolutionary stages, from cool filaments, pre-stellar cores and protostars to more evolved regions in which young stars are heating the dust.

Image release: July 9th 2012 HOBYS Key Program/ESA PACS & SPIRE Consortia



HOPS team discovers extremely young protostars

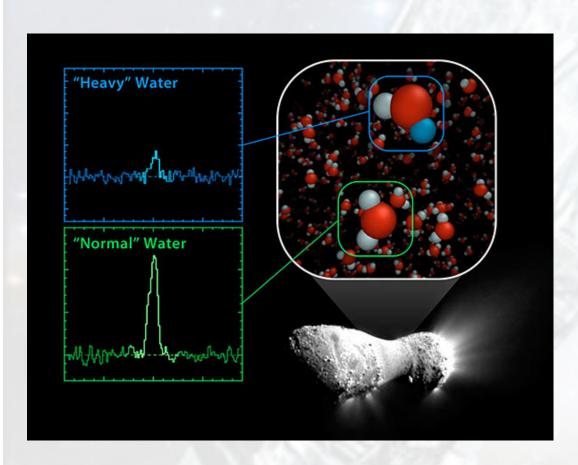


Observations made by the HOPS team (U. of Toledo) Have discovered four Protostars which are so cold they did not show up in previous Spitzer imaging of The Orion region (RHS). The Herschel image (LHS) shows 4 circles which contain very cold protostars. The same circles on the Spitzer Image show no emission.

March 19 2013

erschel 160 microns (green) Green (IRAC band 3,4) and SOME KEY
HERSCHEL RESULTS
II
Comets and dustdisks around stars

Heavy Water found in Comet Hartley 2 in same proportion to that on Earth: Kuiper Belt Origin of Earth's Oceans?



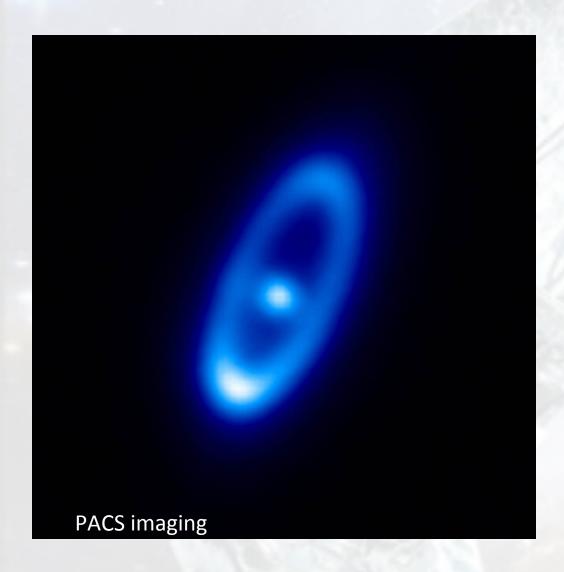
Space Observatory Provides Clues to Creation of Earth's Oceans

Press Release • nhsc2011-017 • Wednesday, October 5th, 2011

PASADENA, Calif. -- Astronomers have found a new cosmic source for the same kind of water that appeared on Earth billions of years ago and created the oceans. The findings may help explain how Earth's surface ended up covered in water.

New measurements from the Herschel Space Observatory show that comet Hartley 2, which comes from the distant Kuiper Belt, contains water with the same chemical signature as Earth's oceans. This remote region of the solar system, some 30 to 50 times as far away as the distance between Earth and the sun, is home to icy, rocky bodies including Pluto, other dwarf planets and innumerable comets.

"Our results with Herschel suggest that comets could have played a major role in bringing vast amounts of water to an early Earth," said Dariusz Lis, senior research associate in physics at the California Institute of Technology in Pasadena and co-author of a new paper in the journal Nature, published online today, Oct. 5. "This finding substantially expands the reservoir of Earth ocean-like water in the solar system to now include icy bodies originating in the Kuiper Belt.



Herschel Spots Comet Massacre Around Nearby Star FORMALHAUT

Feature • nhsc2012-006 • Thursday, April 12th, 2012

The Herschel Space Observatory has studied the dusty belt around the nearby star Fomalhaut. Scientists say the dust appears to be coming from collisions that destroy up to thousands of icy comets every day.

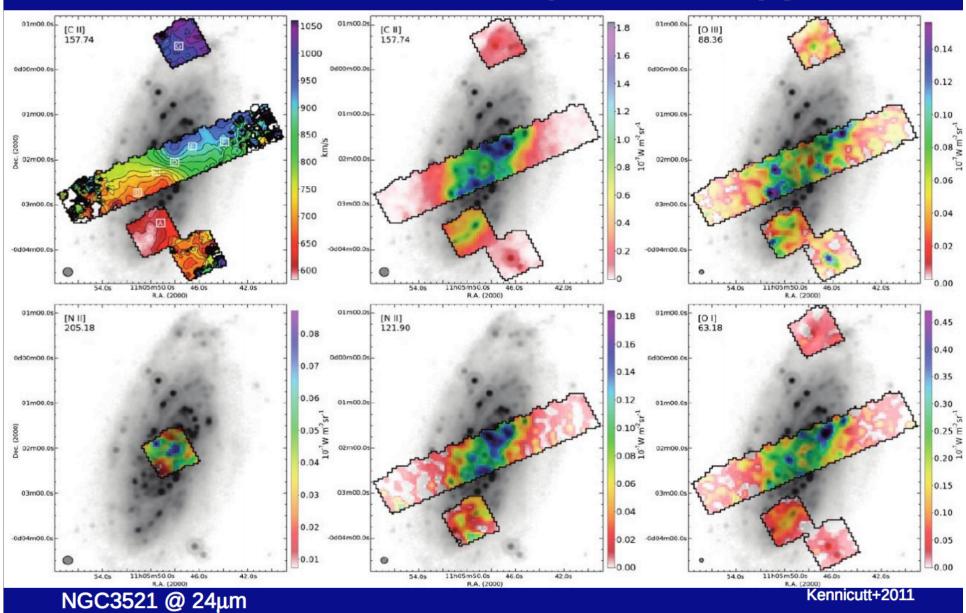
Fomalhaut is a young star, just a few hundred million years old, and twice as massive as the sun. Its dust belt was discovered in the 1980s by the Infrared Astronomical Satellite, in which NASA played a key role. Herschel's new images of the belt show it in much more detail at longer infrared wavelengths than ever before.

The results indicate the grains in the dust belt are fluffy and tiny, only a few millionths of a meter across (one meter is about 3 feet). They are similar to dust particles released from comets in our own solar system.

SOME KEY HERSCHEL RESULTS III

Measuring the Far-IR Cooling lines in the ISM of Galaxies

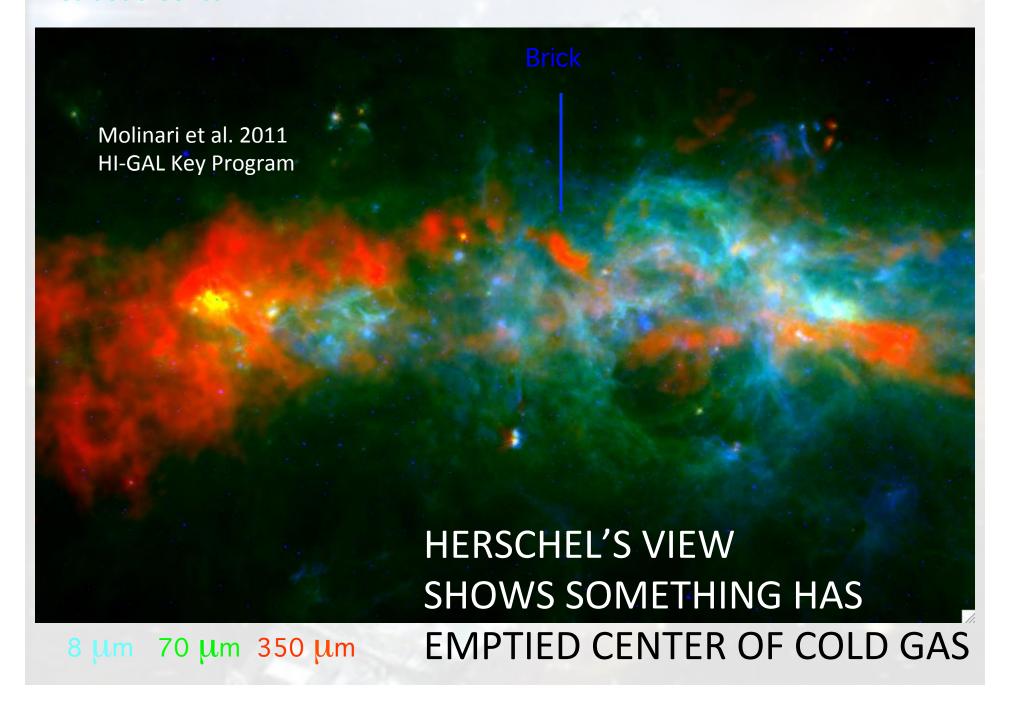
Herschel Far-IR Spectroscopy



Some Key Herschel Results IV

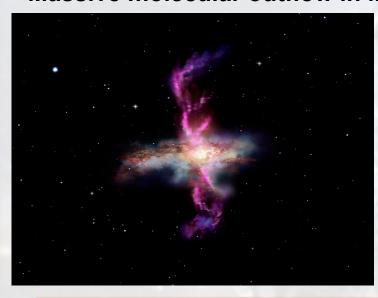
Centers of Galaxies and AGN Feedback

Galactic Center



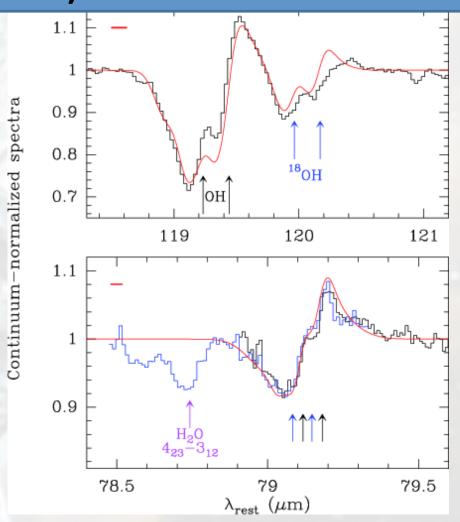
Herschel detects massive molecular outflows in AGN SHINING Key Program (E. Sturm PI)

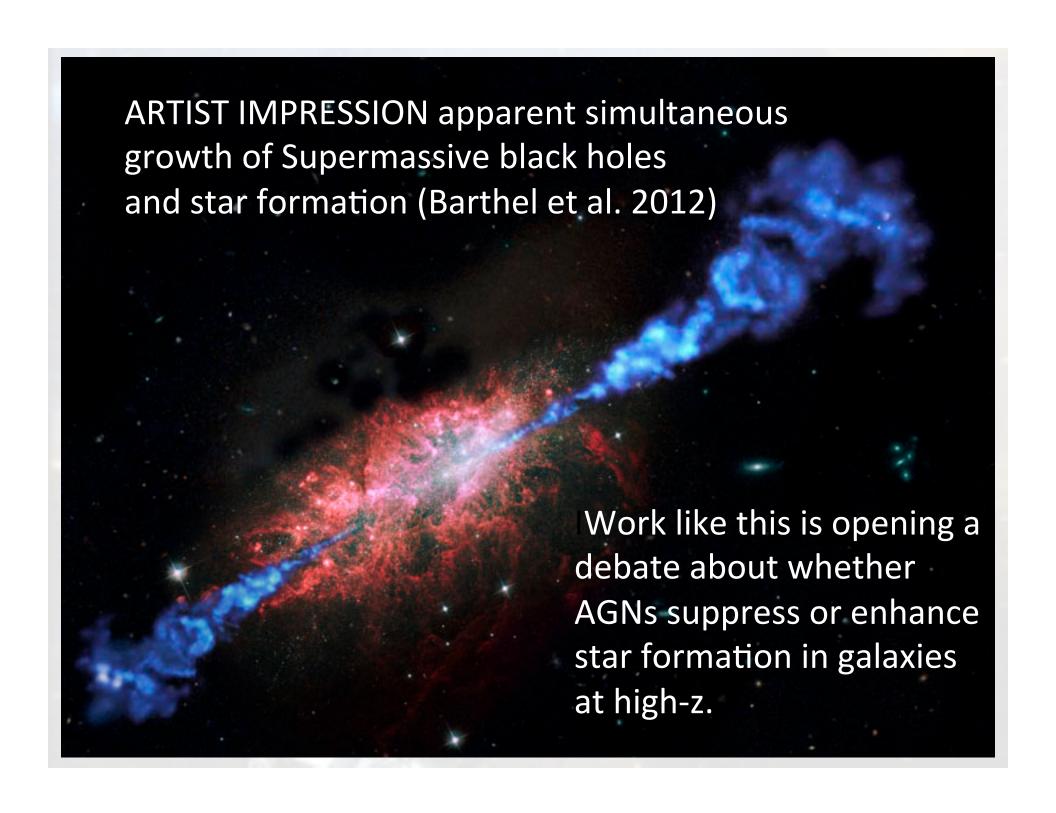
Massive molecular outflow in Mrk 231



Outflow mass of 7x10⁷ M_☉
Outflow velocities of -1400 km/s
Mechanical energy ≥ 10⁵⁶ ergs

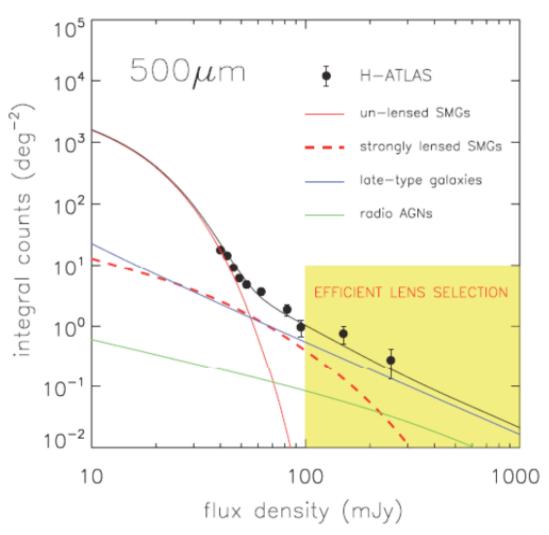
Fischer et al. 2010



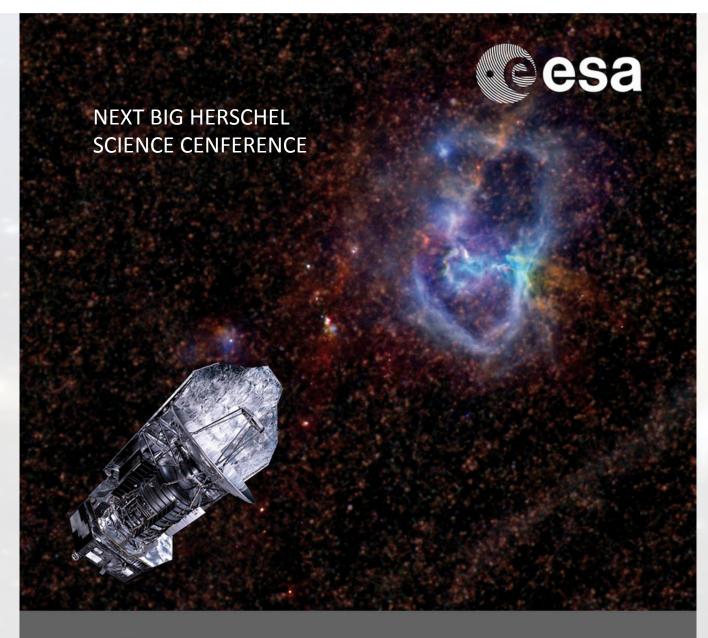


LENSED GALAXIES ARE BRIGHT IN SPIRE BANDS

Galaxy-galaxy gravitational lenses select high-redshift galaxies



Herschel-ATLAS identifies high-z galaxies **ID81** SPIRE image from the Herschel Science **Demonstration Phase ID11 ID17** ID130 ID9



→ THE UNIVERSE EXPLORED BY HERSCHEL

15-18 October 2013 ESA/ESTEC, Noordwijk, The Netherlands

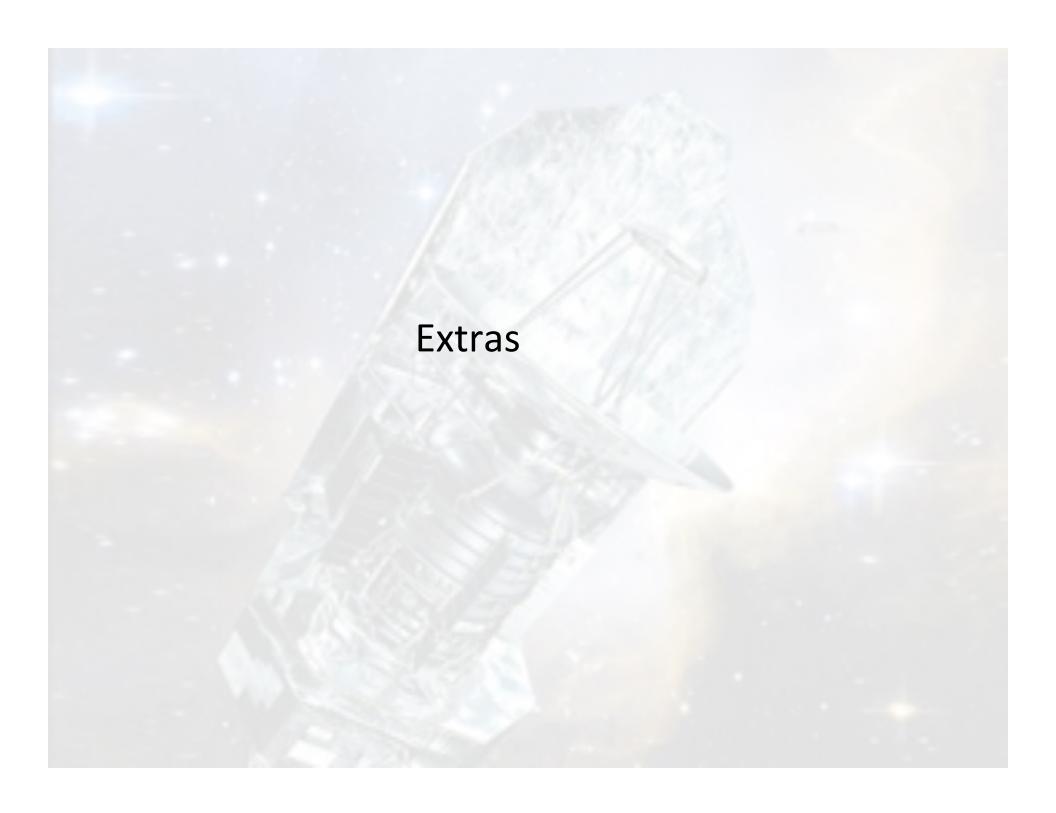
HOW TO GET INVOLVED IN HERSCHEL ISSUES?

NHSC and HSC Helpdesks (US users should work with the NHSC Helpdesk)

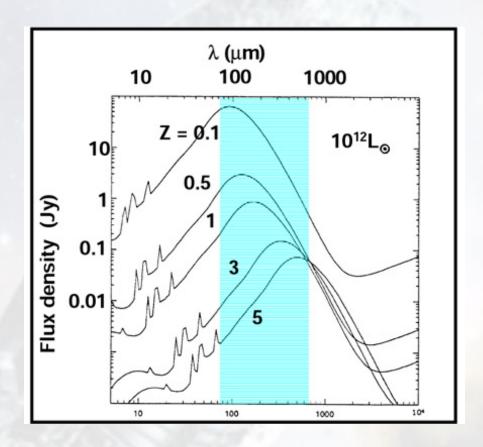
NHSC User Panel (NUP): Current Chair: Margaret Meixner STScl Committee make-up and recent reports—see https://nhscsci.ipac.caltech.edu/sc/index.php/Main/NUP

Herschel User Group (HUG): European User Group led by http://herschel.esac.esa.int/HUG.shtml

DPUG DATA Processing User Group: Deals with DP Issues for whole Herschel project. Although historically has contained mainly Herschel experts and developers, the membership is expanding to include users.

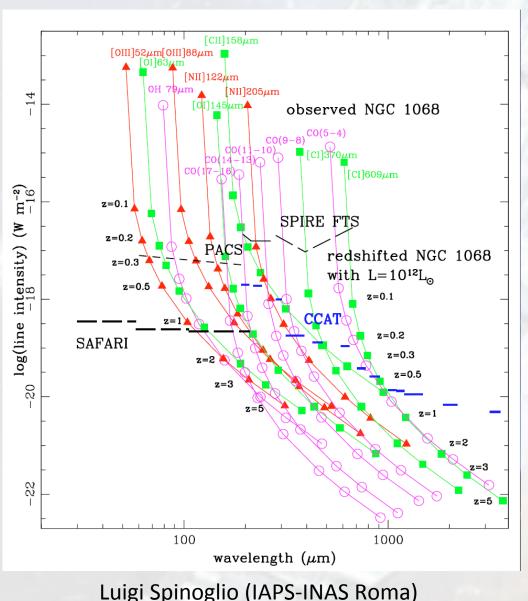


Importance of the FIR & submm



 Herschel covers the IR peak and pushes into the submillimetre: IRbright galaxies

Herschel: CCAT and SPICA/SAFARI

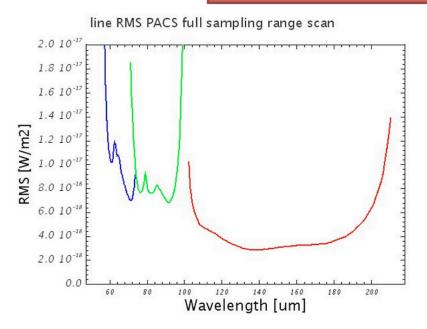


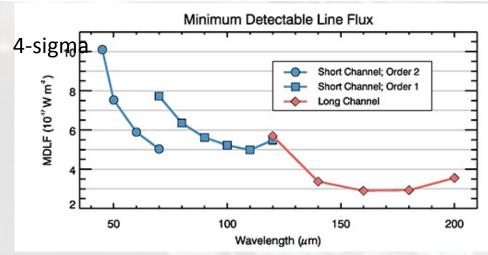
CCAT and SPICA/SAFARI will do for the higher-z what Herschel can do for low-z

Far-IR Spectral Line Diagnostics

5-sigma 1hr sensitivity in W/m^2

SOFIA FIFI-LS versus HERSCHEL PACS





PACS Spectrometer 450s Integration rms line flux

Herschel line sensitivity is ~3-5 better than SOFIA over similar range of wavelengths (integrate for 9-16x longer). 6hr SOFIA = 20mins PACS

SOFIA: FIFI-LS 900s projected Minimum detectable line flux (4x rms)

HERSCHEL



You—as users can help us build synergy with other projects



WISE IR Surveys

Spitzer Warm Mission
Deep IR Imaging from Space



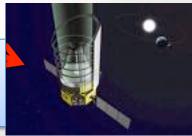
JWST Mid-IR from Space



SOFIA mid/Far IR Trom aircraft



SPICA Far-IR from Space



ALMA AND CAT sub-mm astronomy from ground



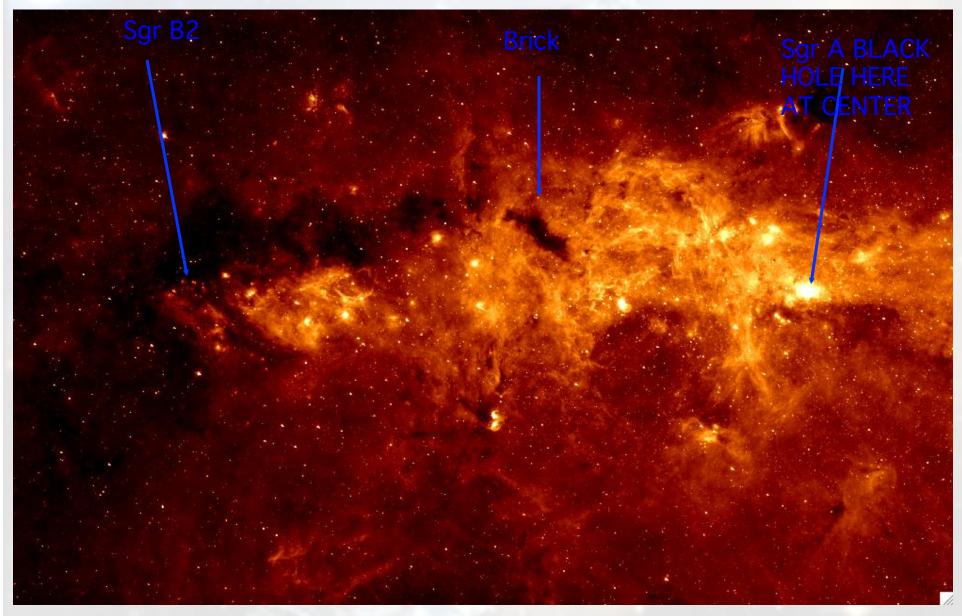


GBT and eVLA Deep Radio Astronomy





Galactic Center



8 μm NEAR INFRARED FROM SPITZER SPACE TELESCOPE