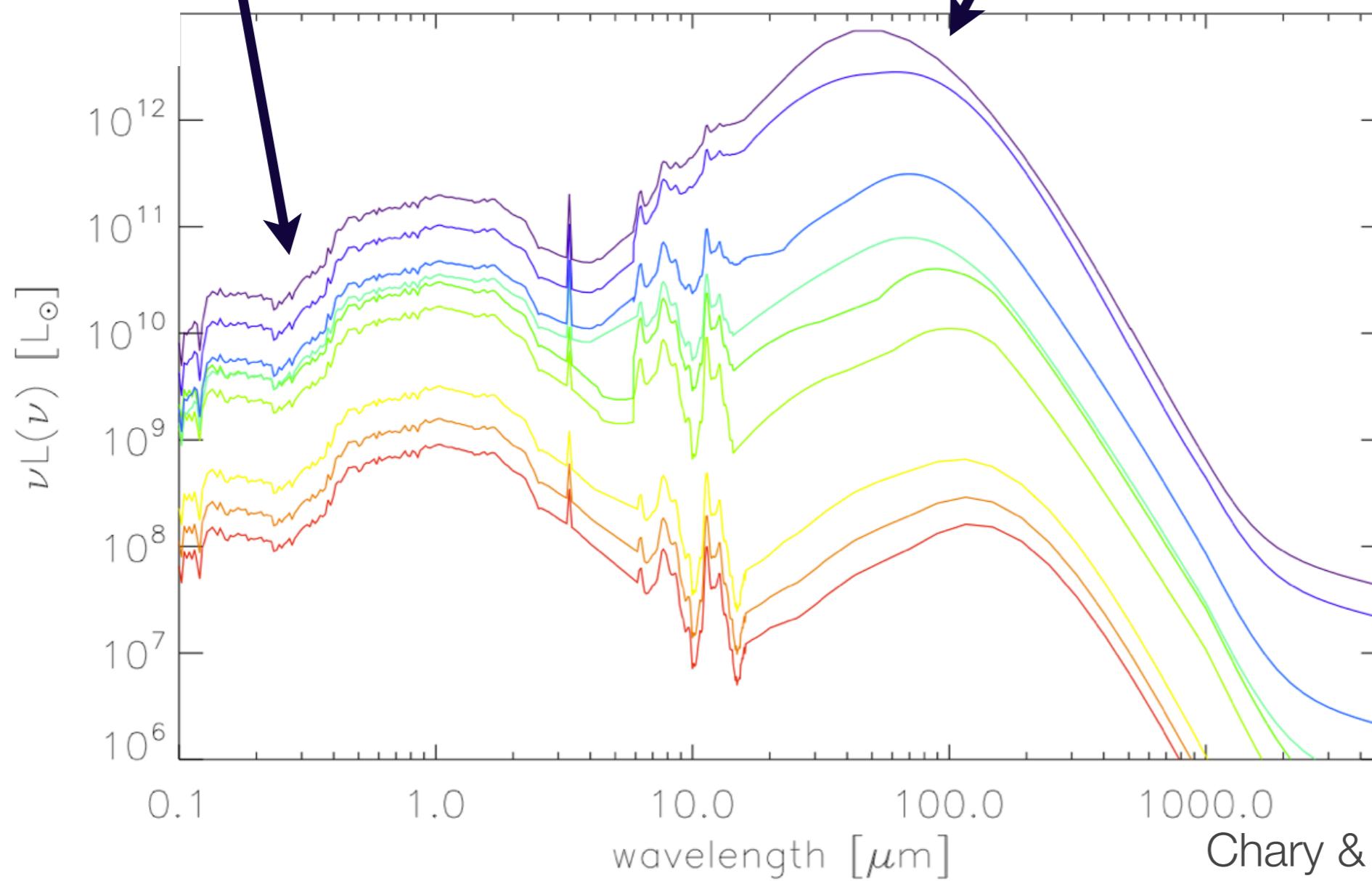


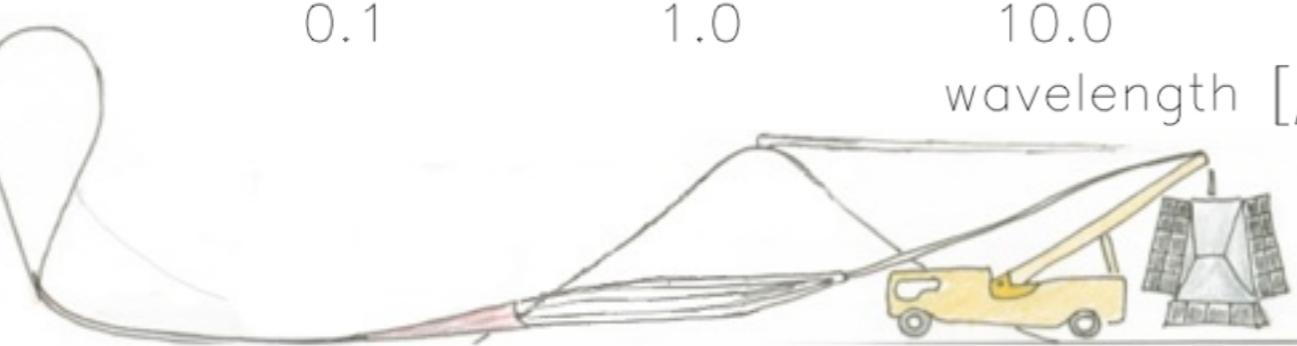
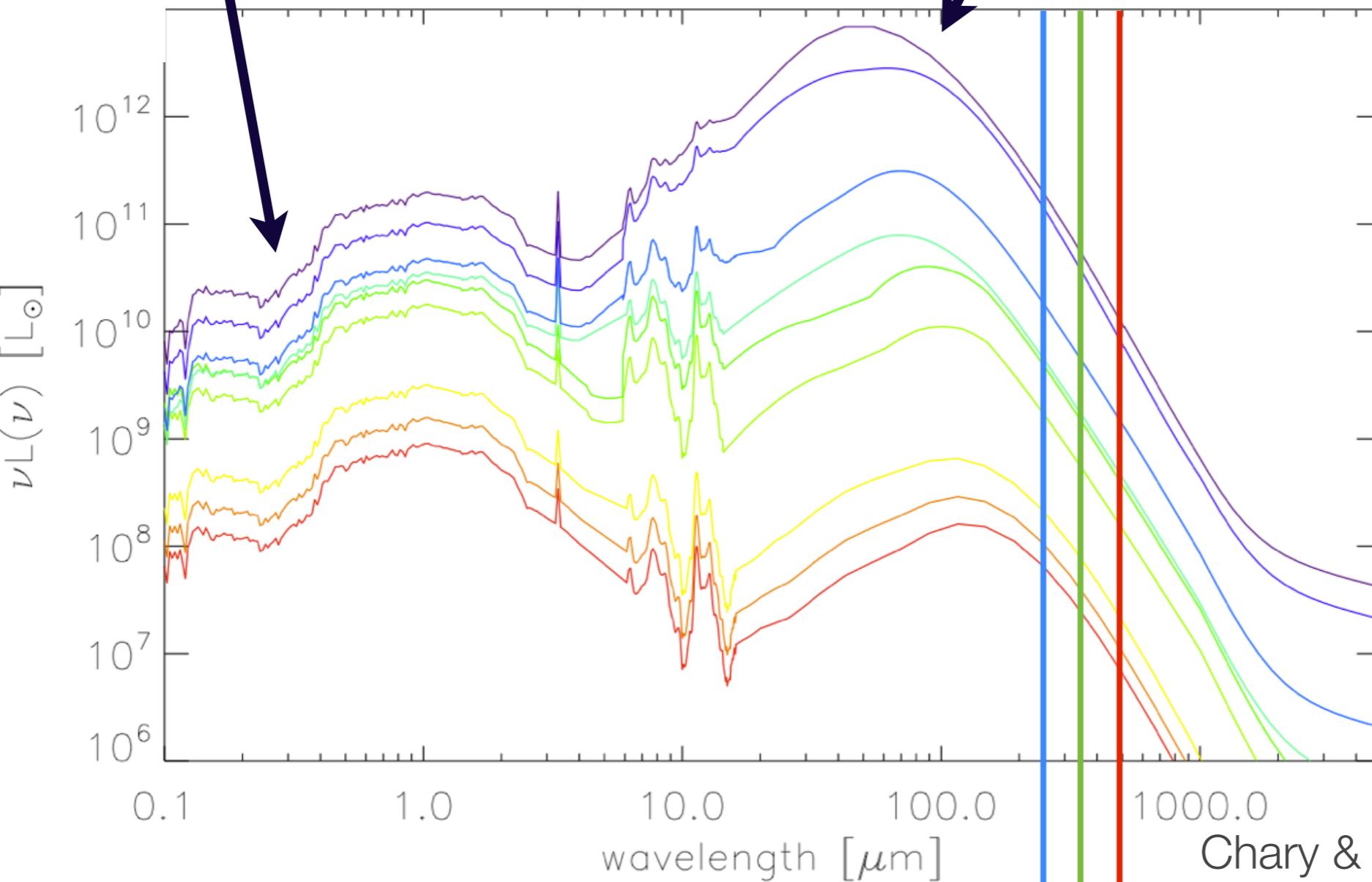
M31 in the optical and IR

Dust re-emits in the FIR  
Optical/UV Starlight absorbed by dust

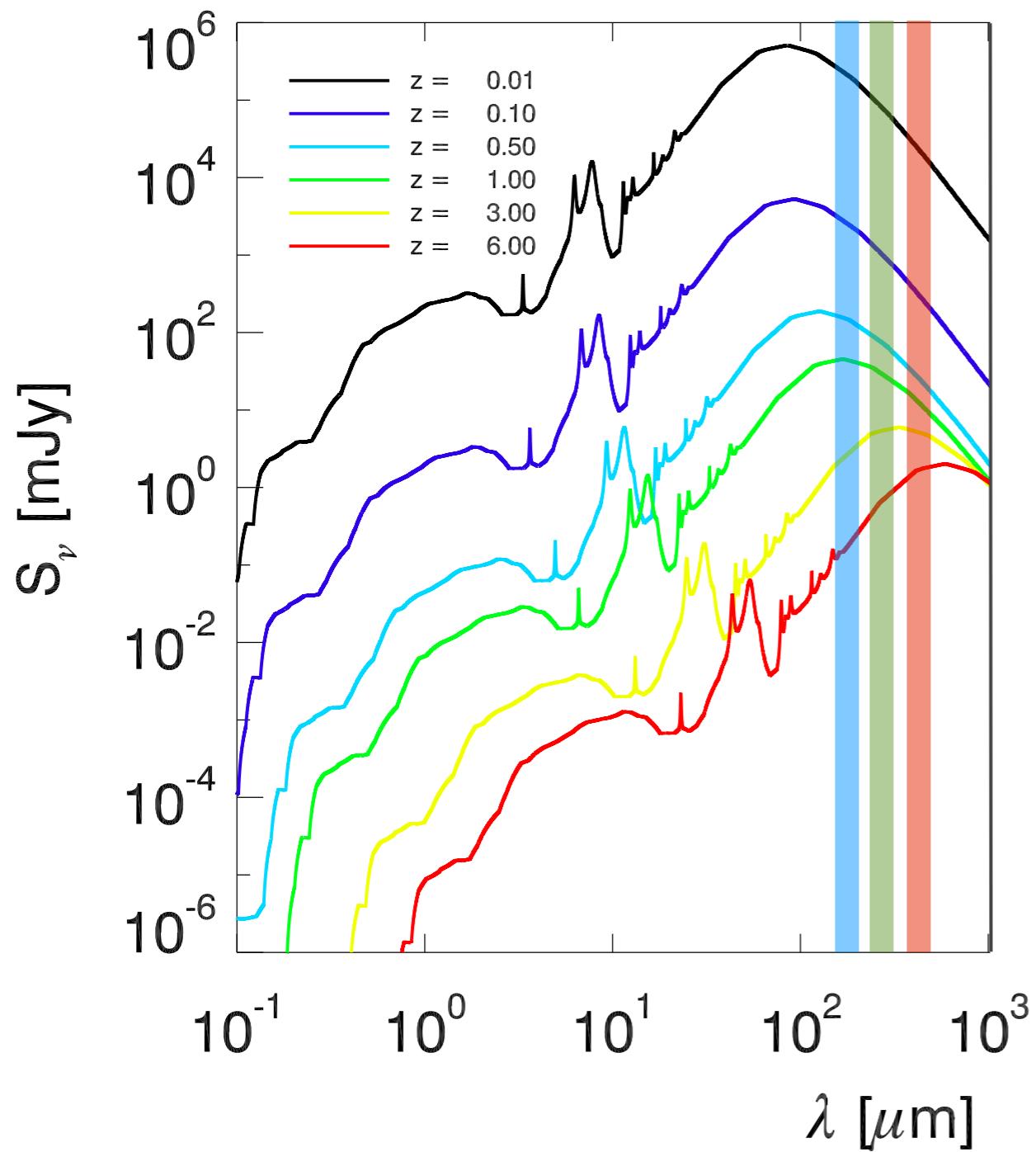


UV/optical to IR relationship

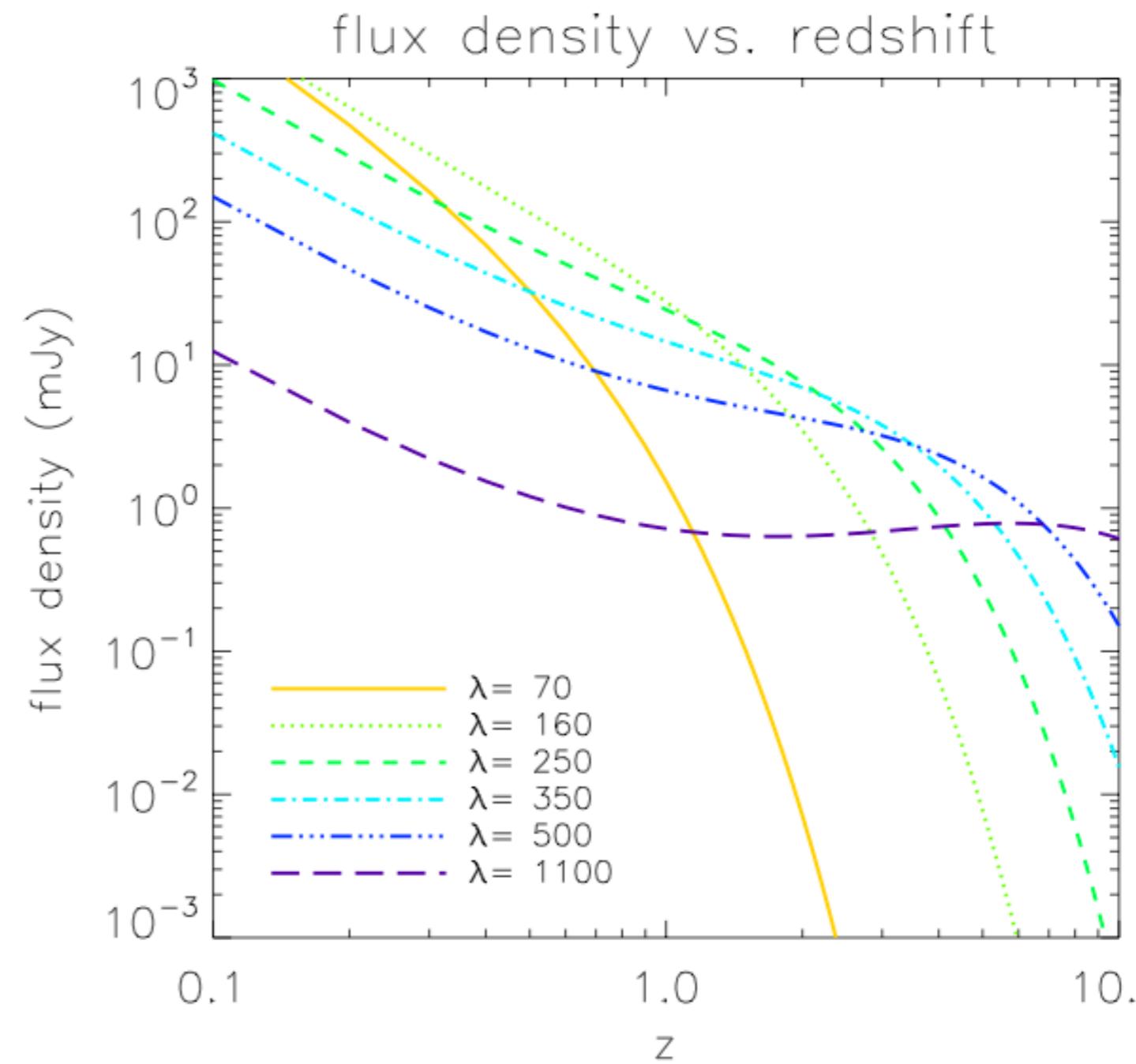
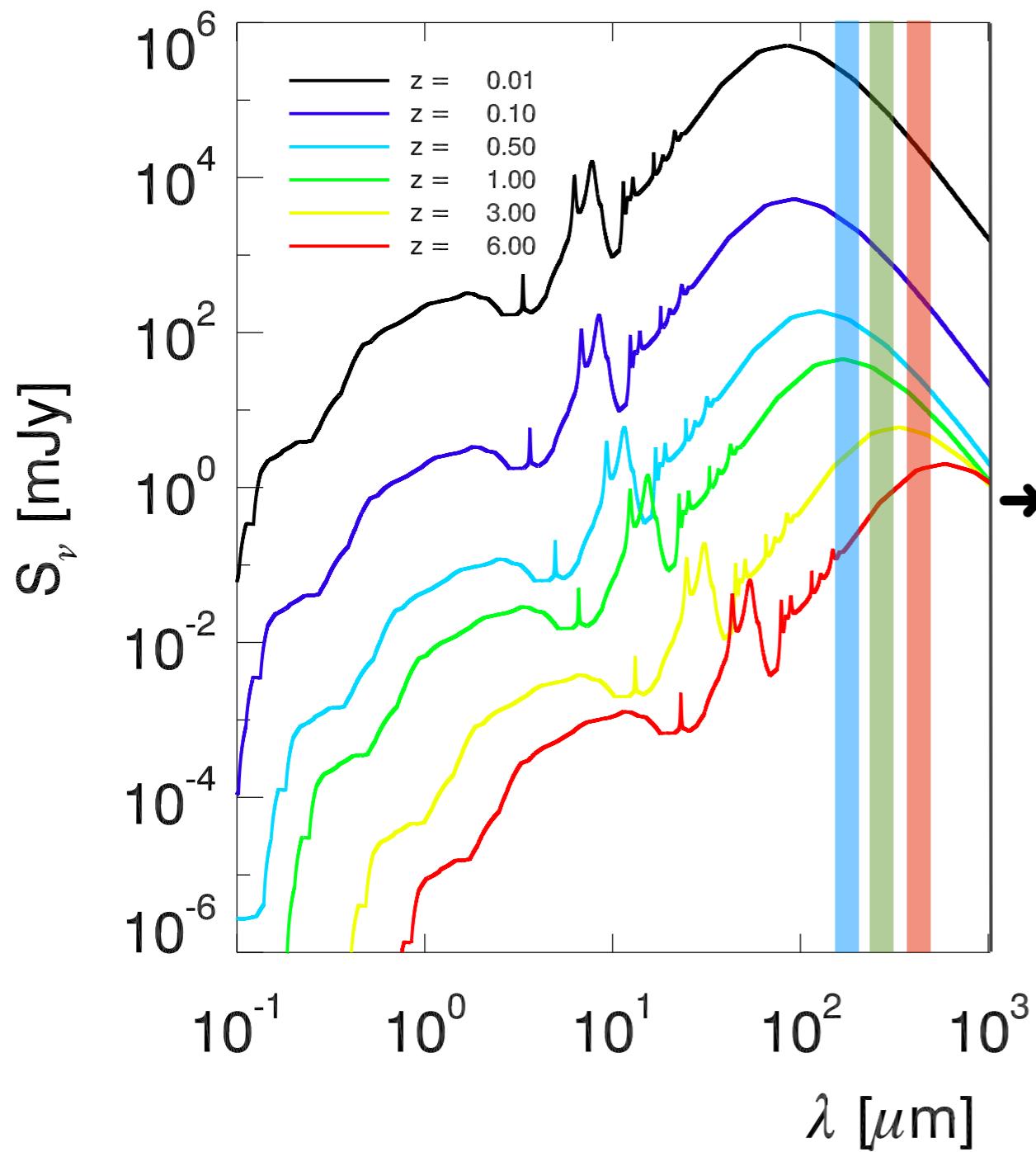
Dust re-emits in the FIR  
Optical/UV Starlight absorbed by dust



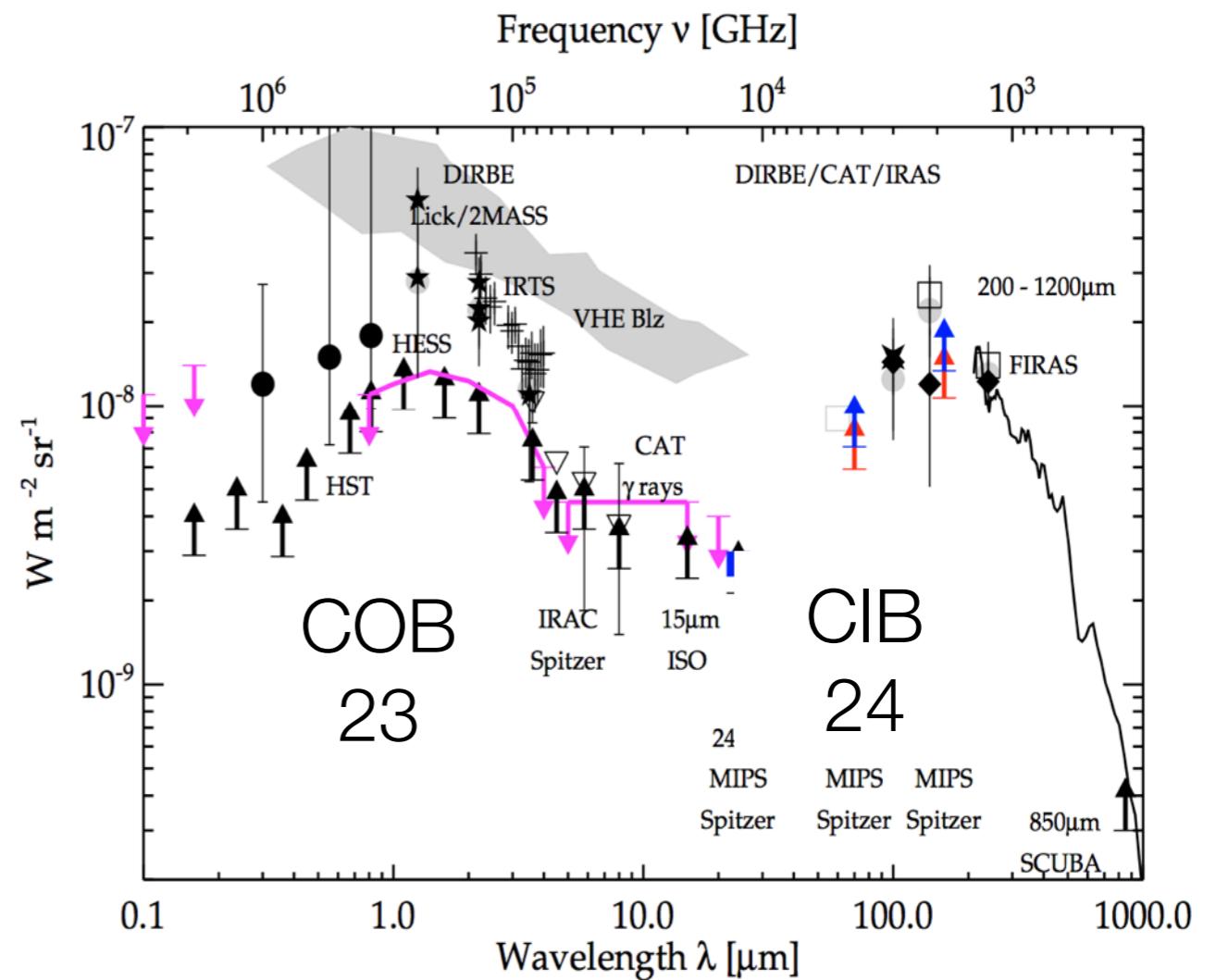
UV/optical to IR relationship



negative  $K$ -correction

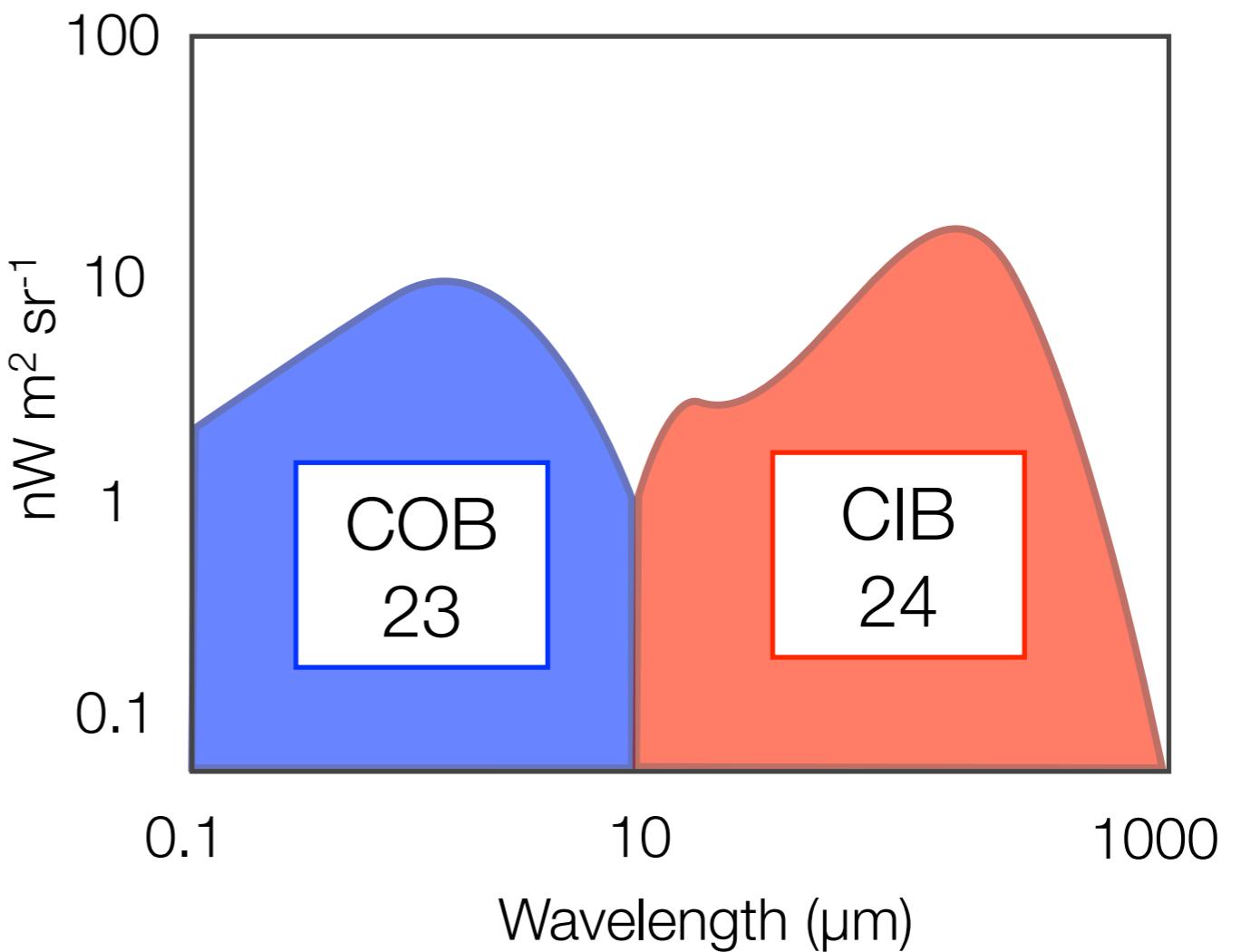


negative  $K$ -correction

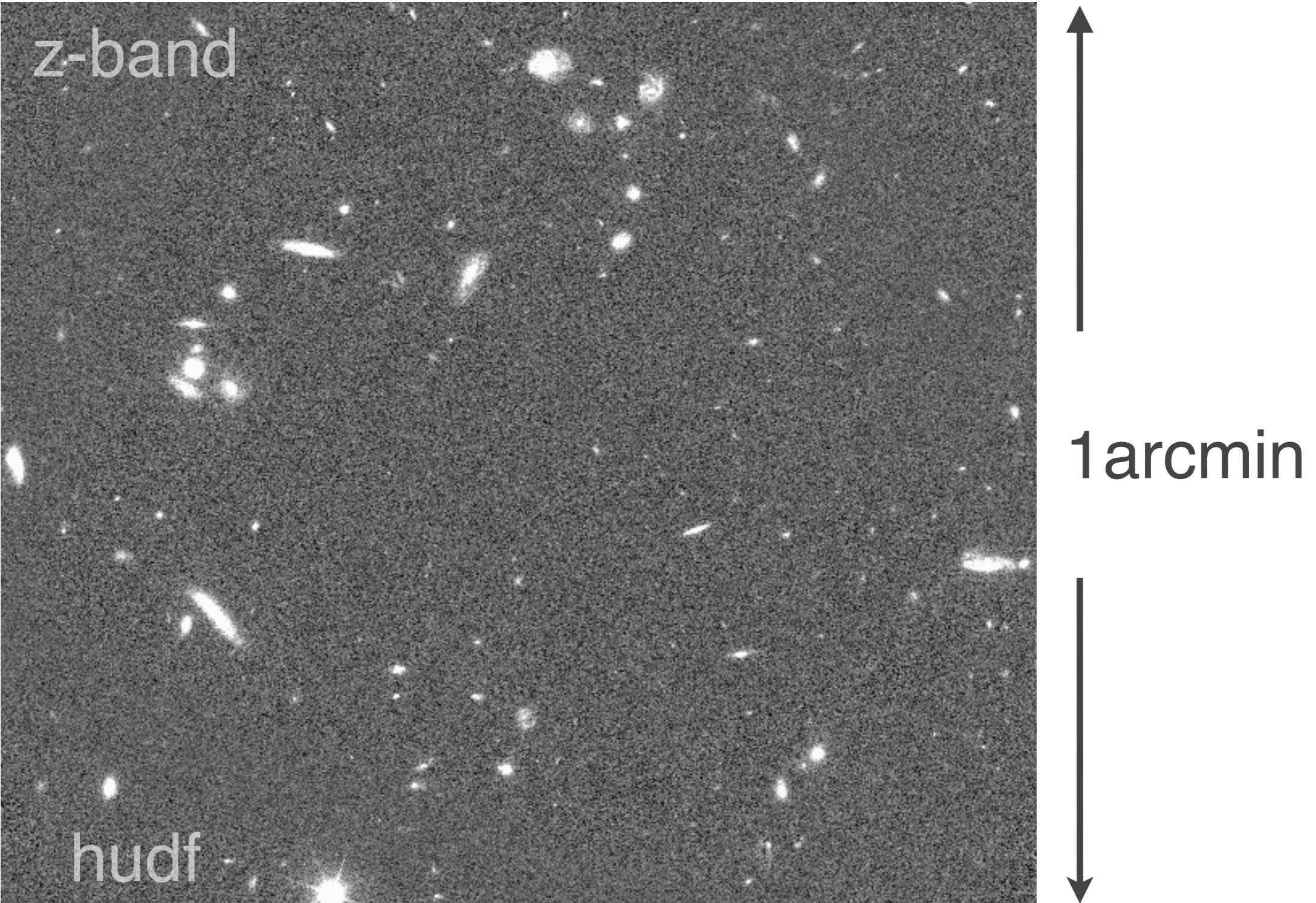


Dole et al (2006)

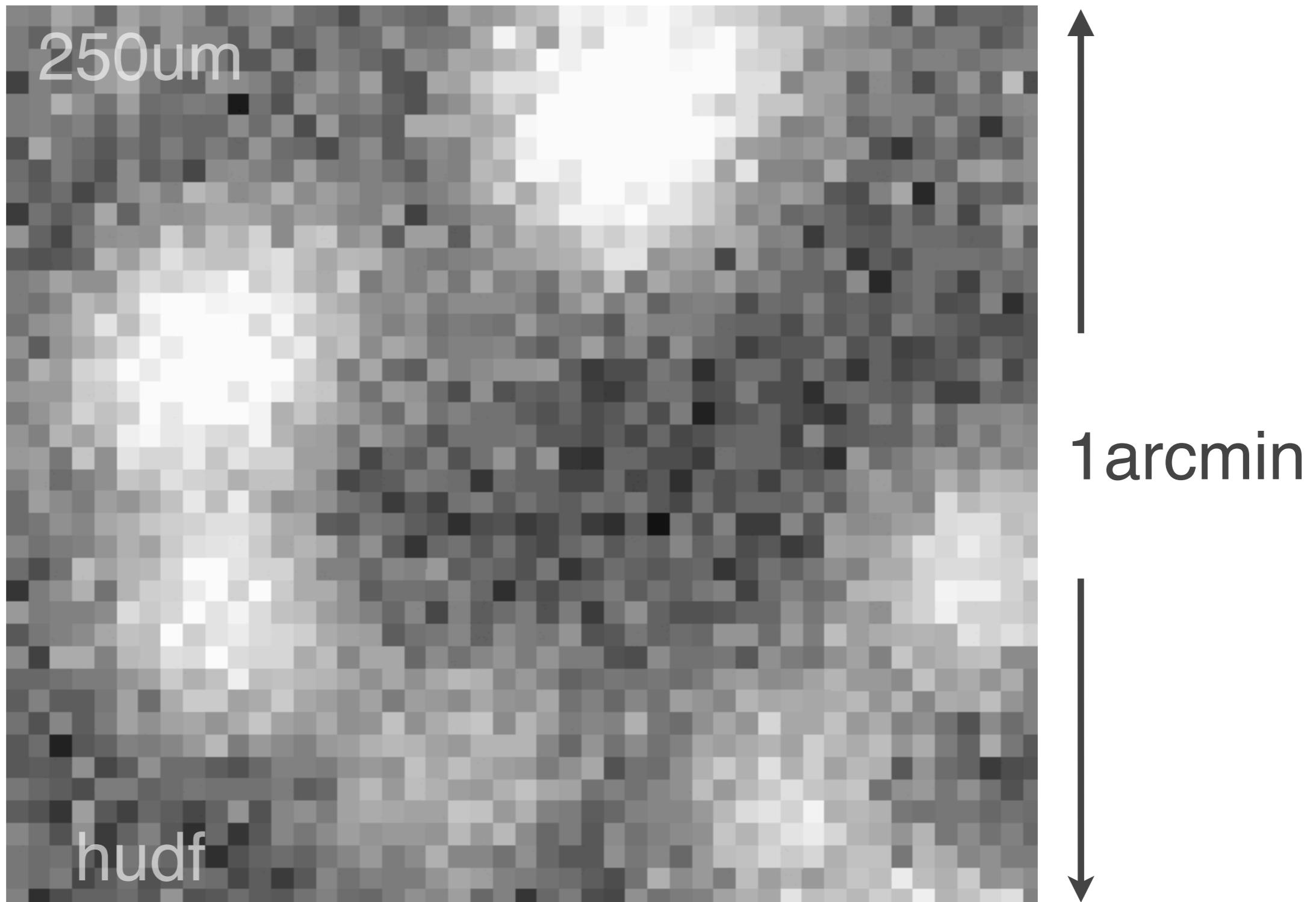
# optical and infrared backgrounds



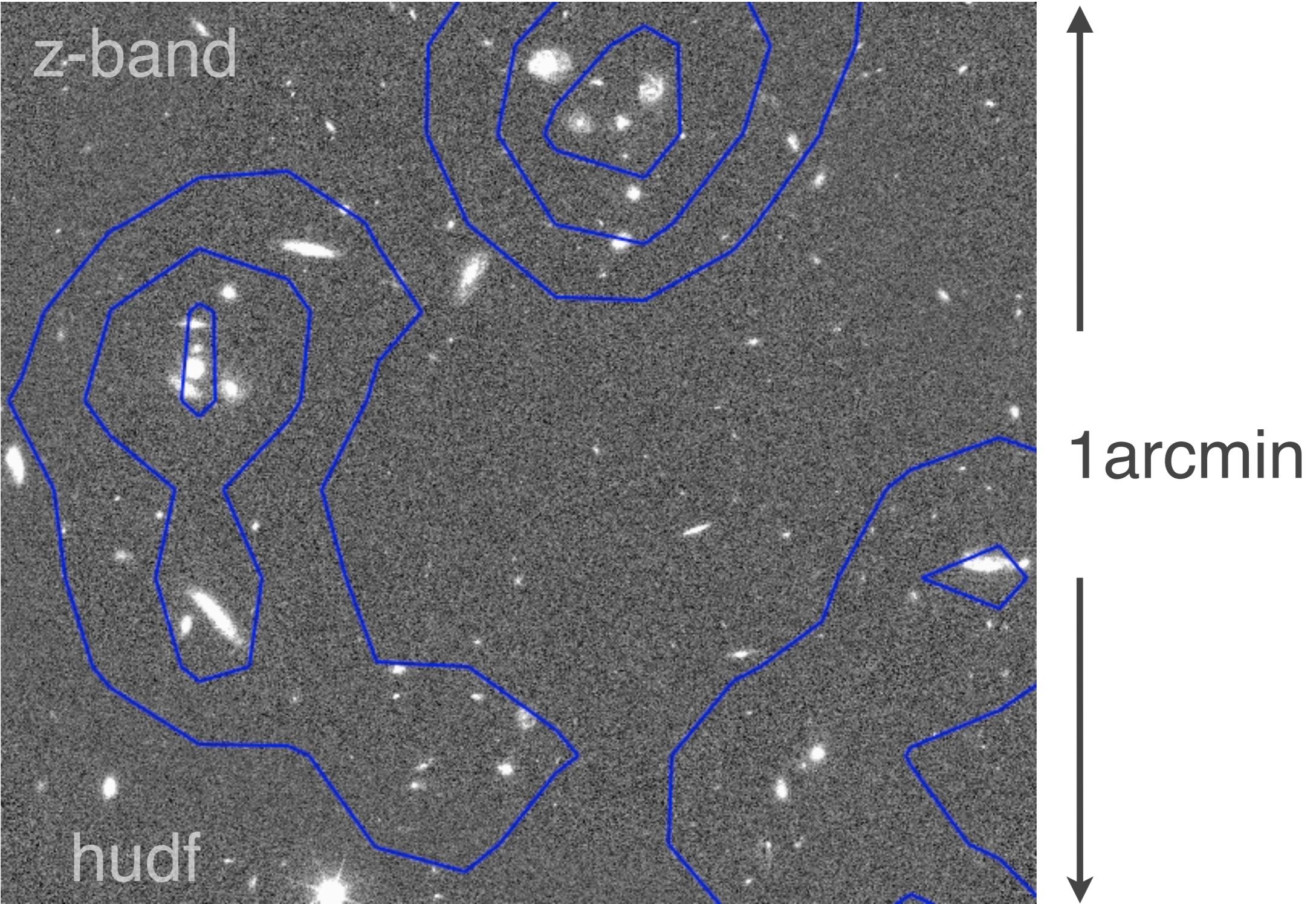
optical and infrared  
backgrounds



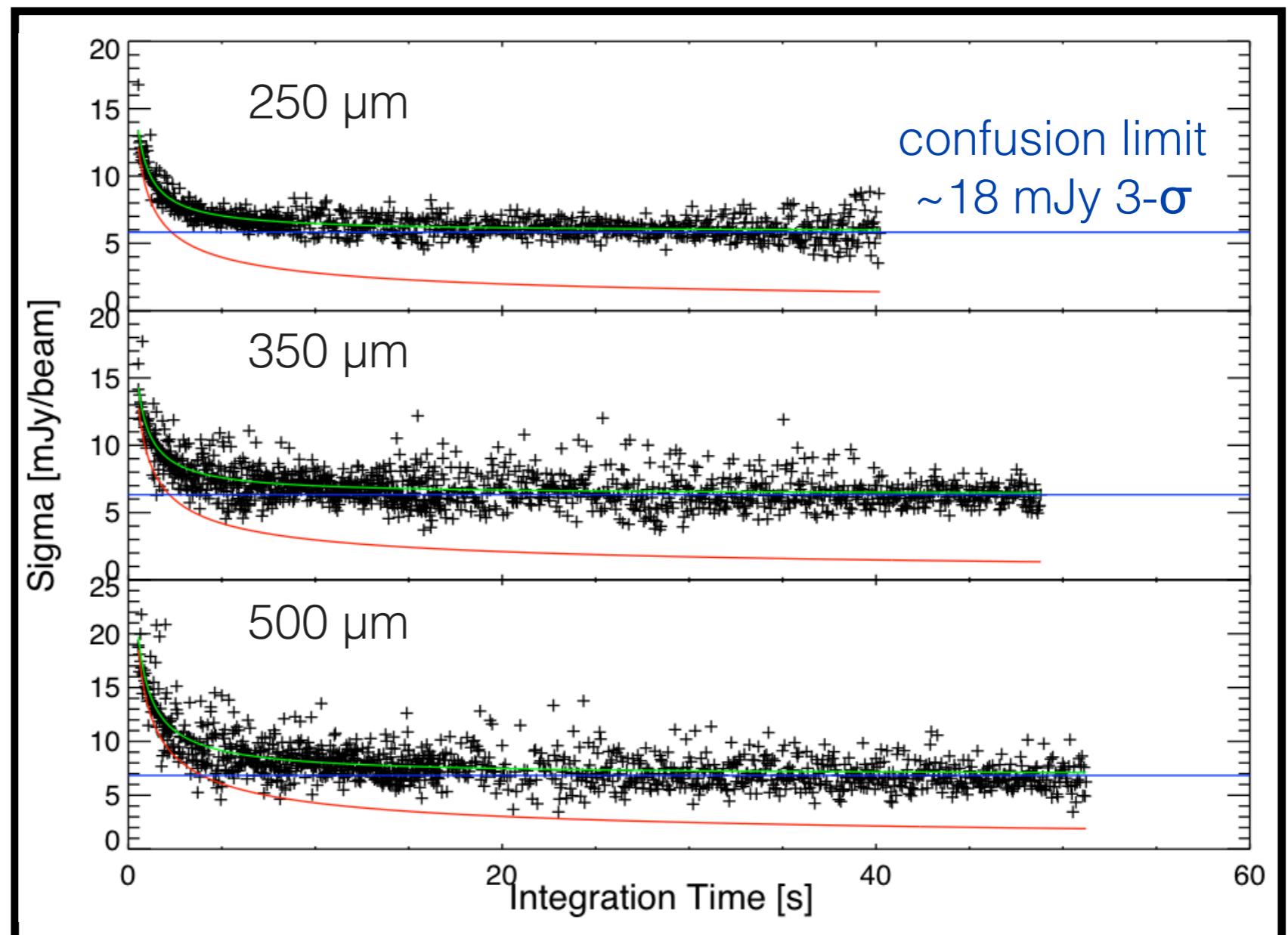
**source confusion**



**source confusion**



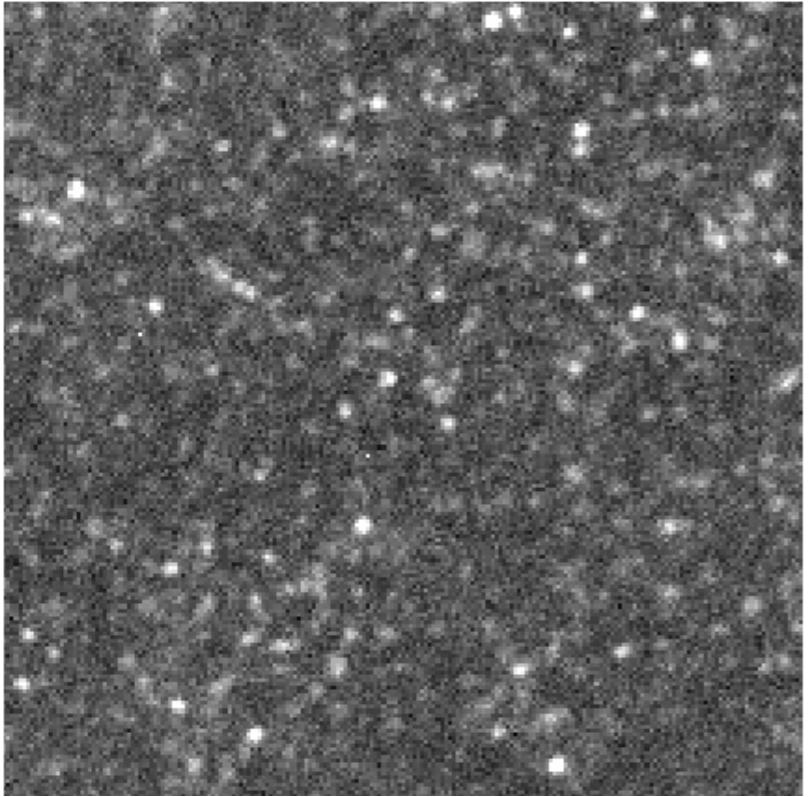
**source confusion**



Nguyen et al. (2009)

# source confusion

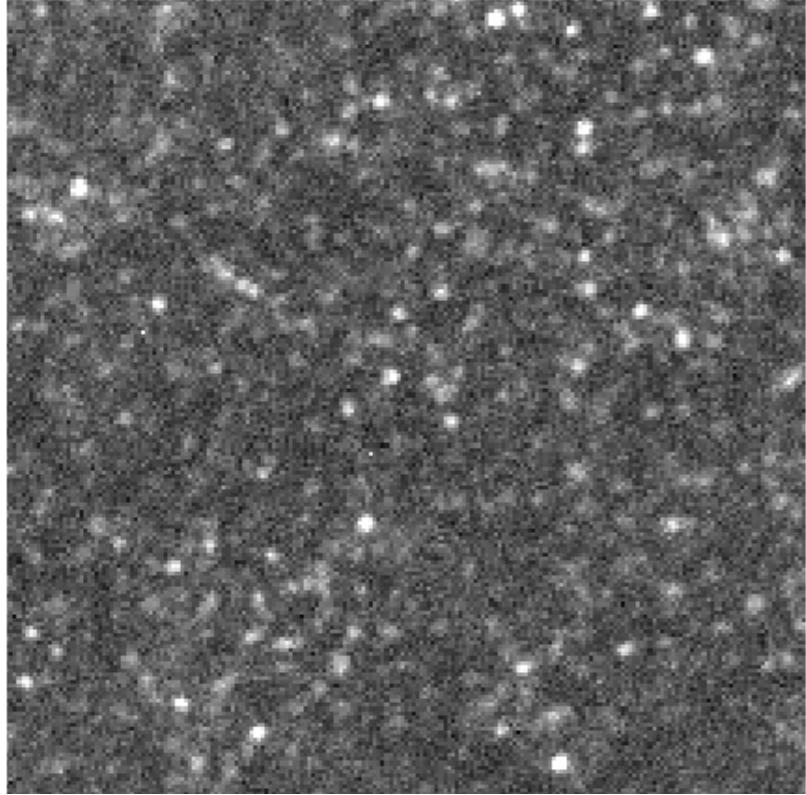
2 repeats (fls)



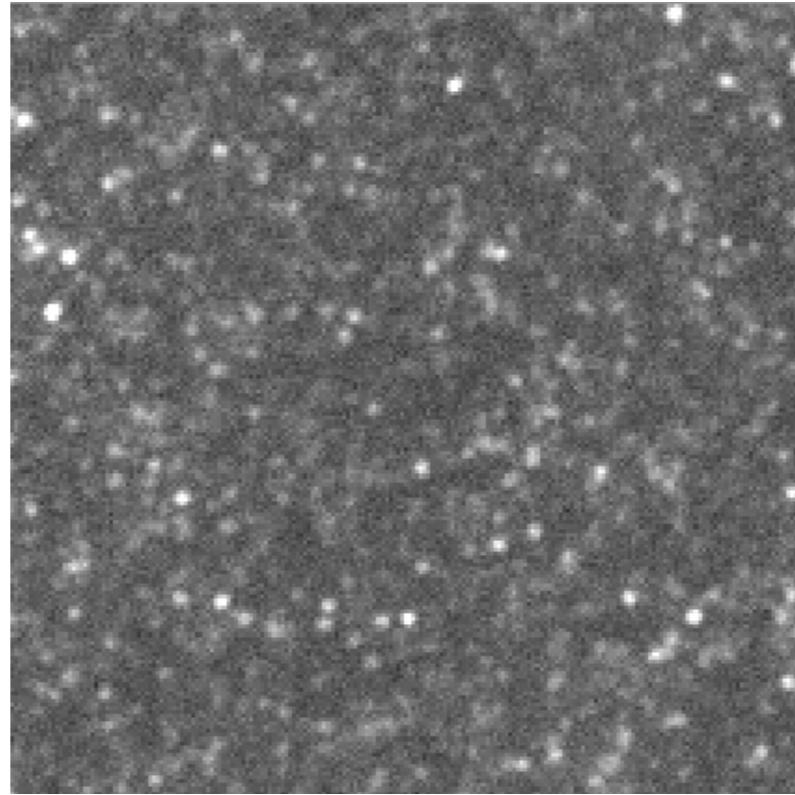
33.5 mJy ( $3\sigma$ )

**source confusion**

2 repeats (fls)



7 repeats (uds)

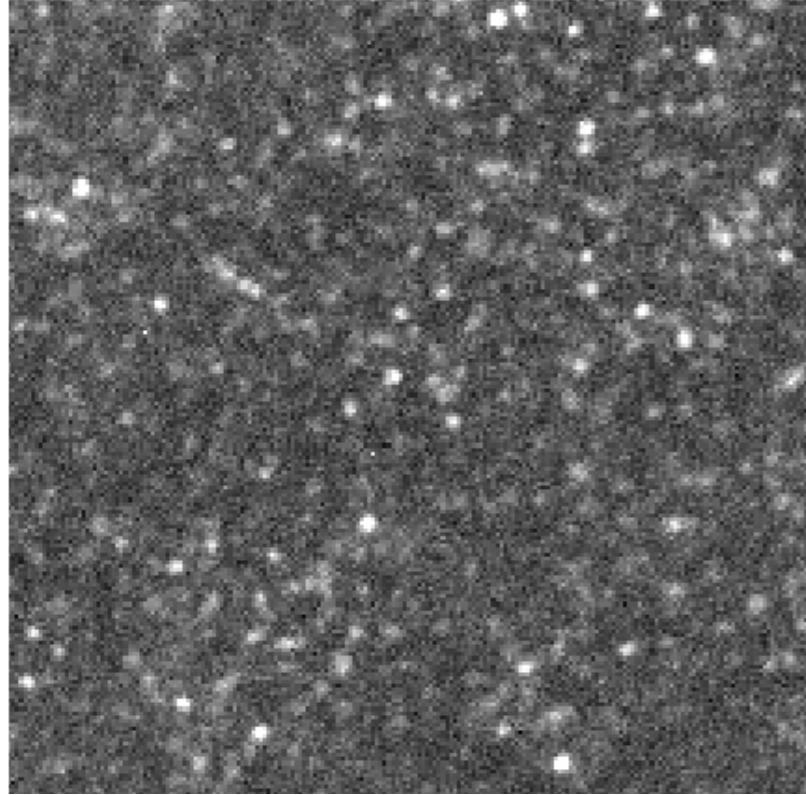


33.5 mJy (3 $\sigma$ )

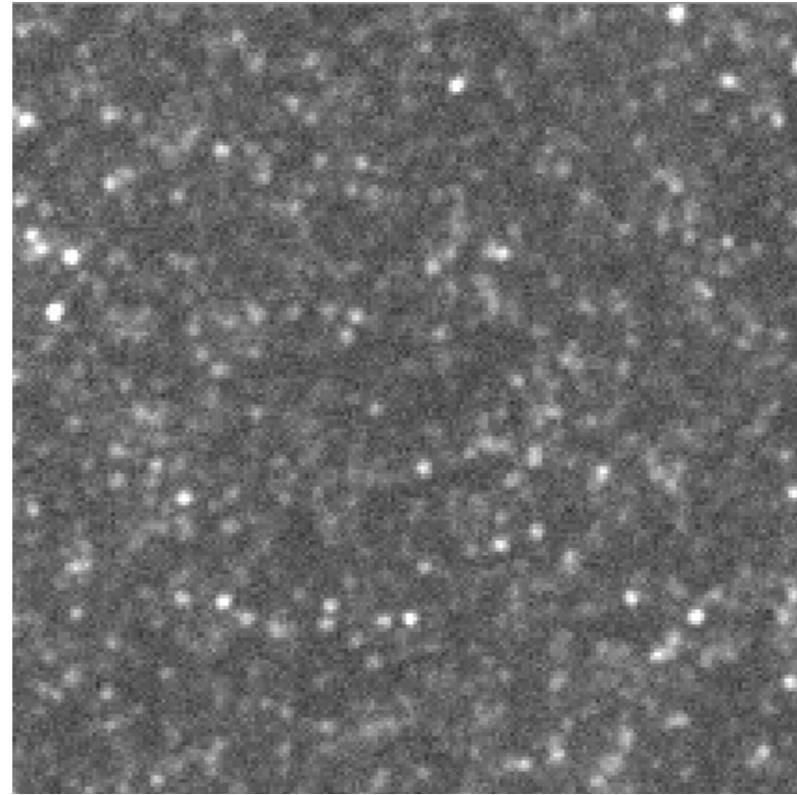
24.7 mJy (3 $\sigma$ )

**source confusion**

2 repeats (fls)



7 repeats (uds)



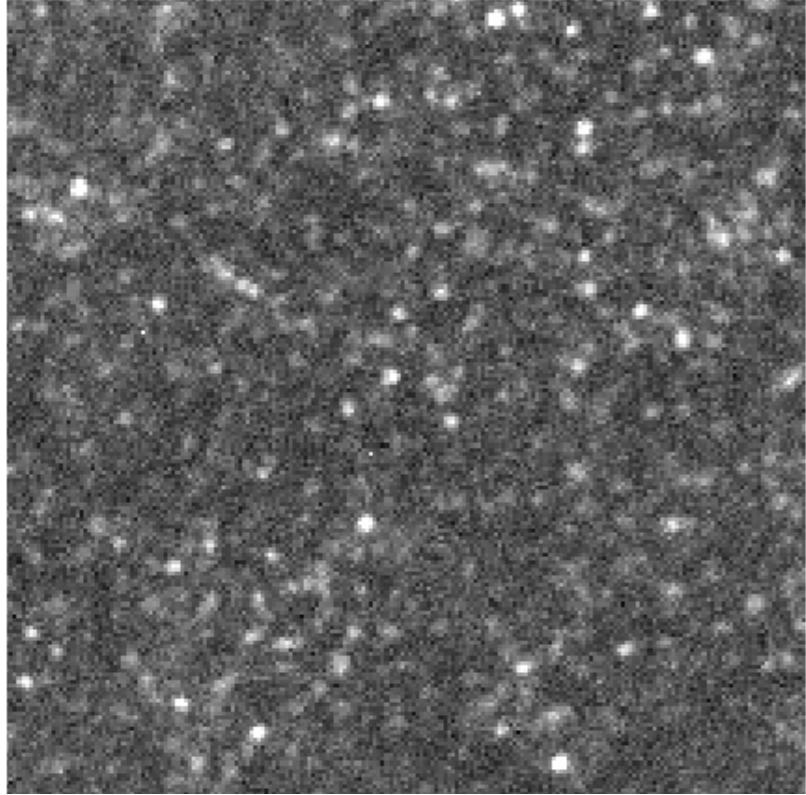
76 repeats (goods-s)

33.5 mJy (3 $\sigma$ )

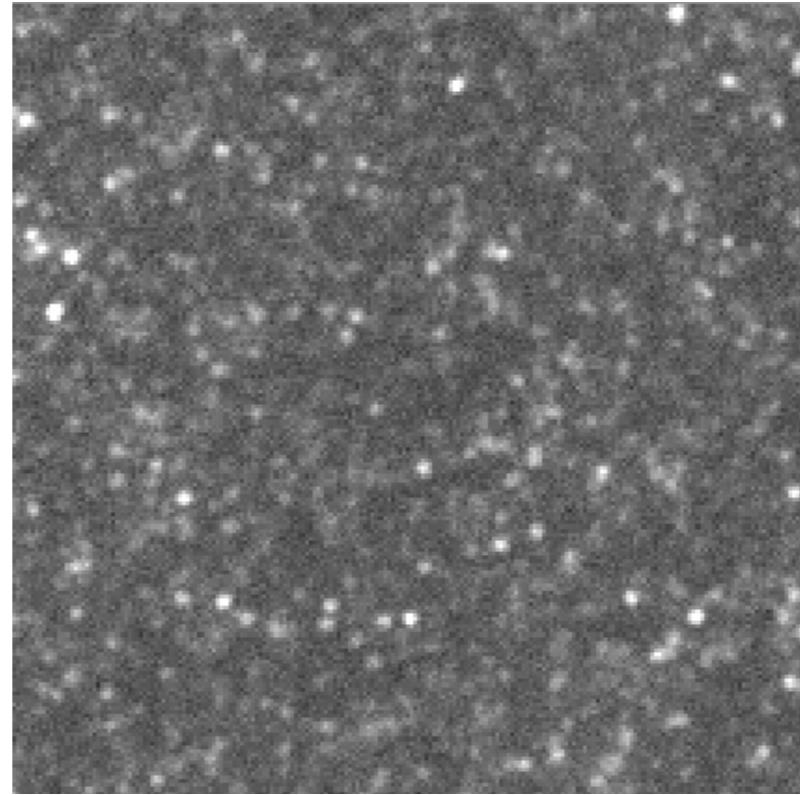
24.7 mJy (3 $\sigma$ )

**source confusion**

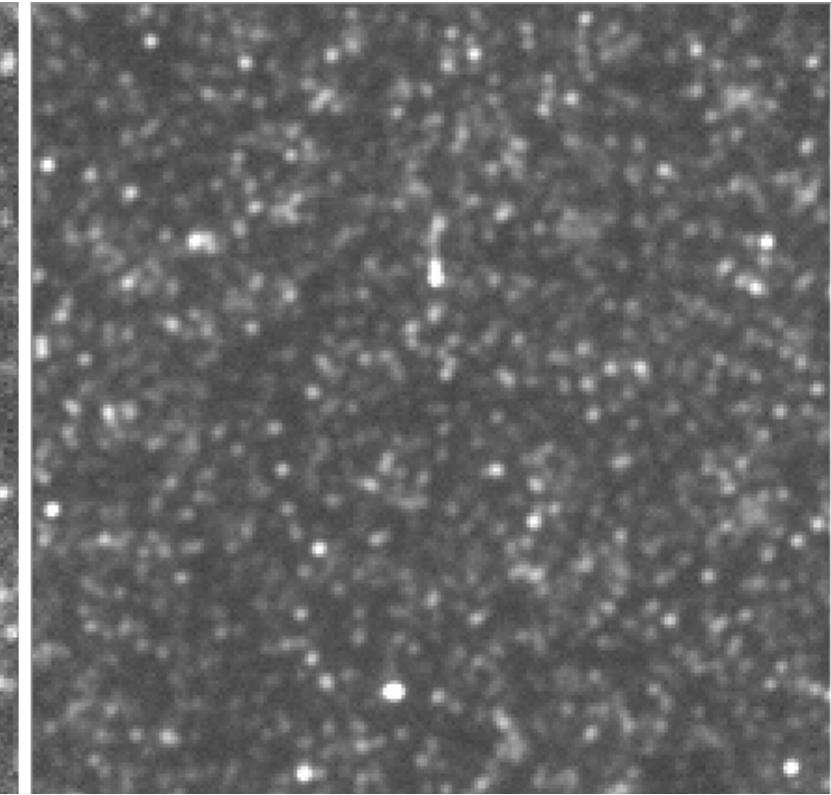
2 repeats (fls)



7 repeats (uds)



76 repeats (goods-s)

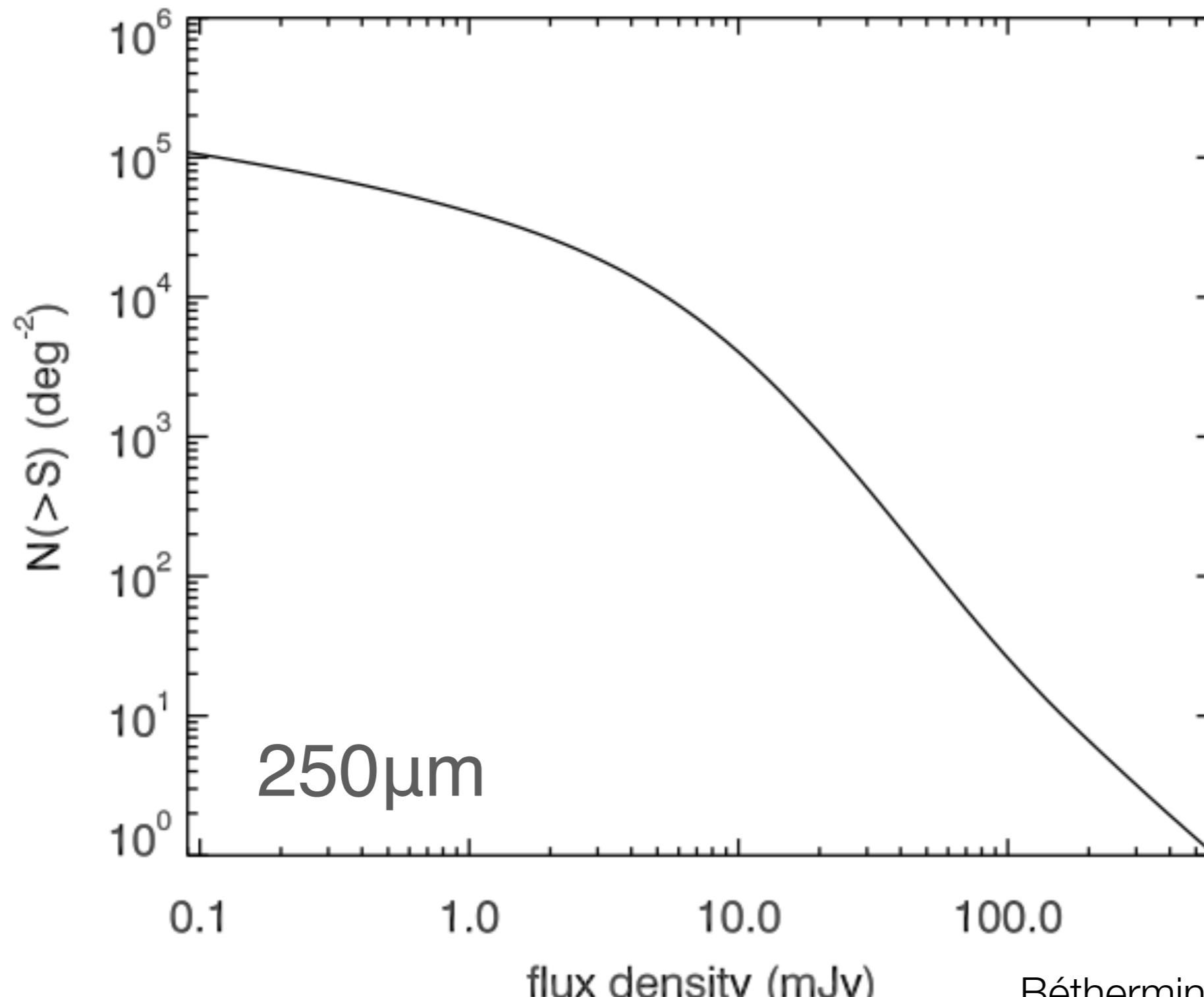


33.5 mJy (3 $\sigma$ )

24.7 mJy (3 $\sigma$ )

20.2 mJy (3 $\sigma$ )

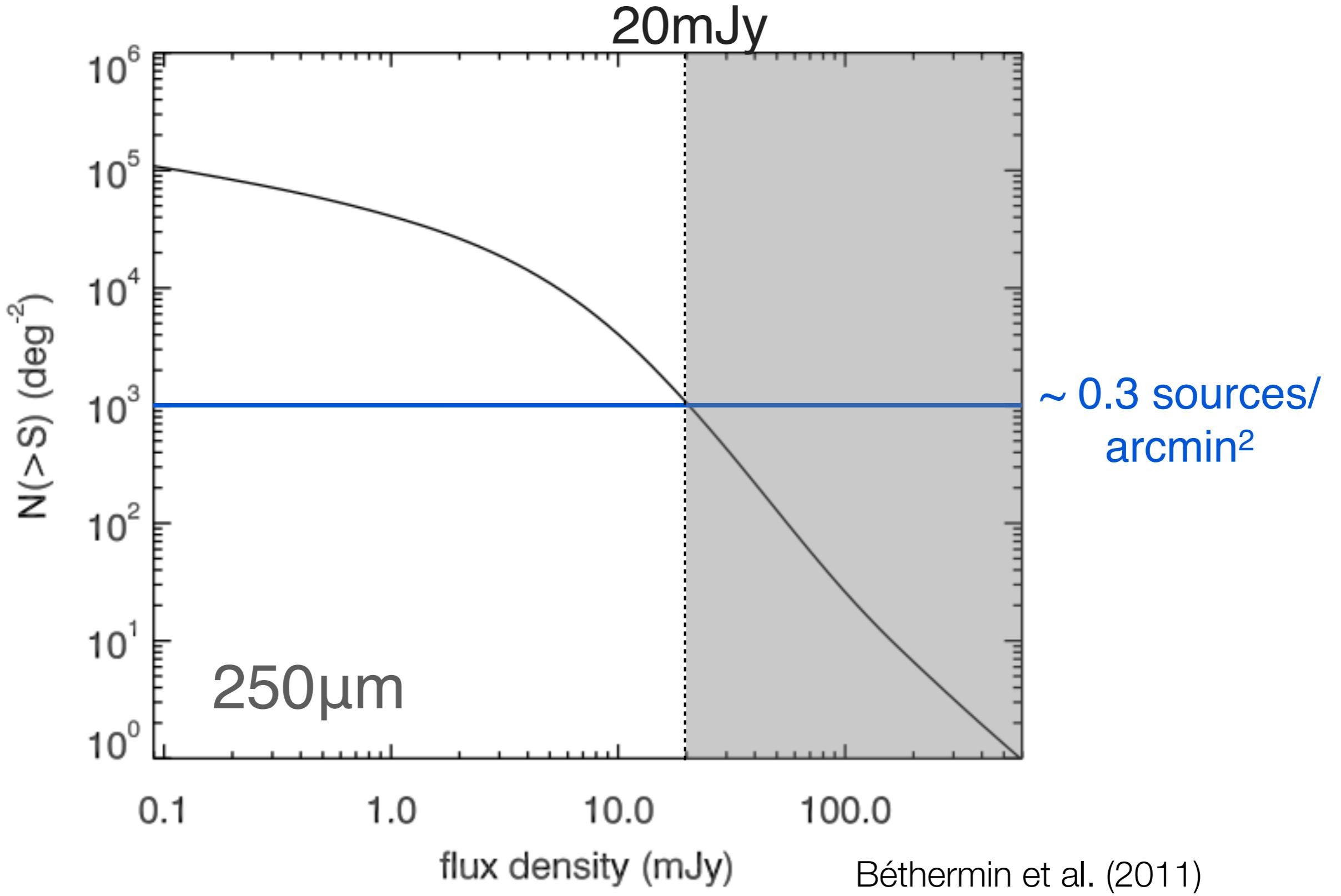
source confusion



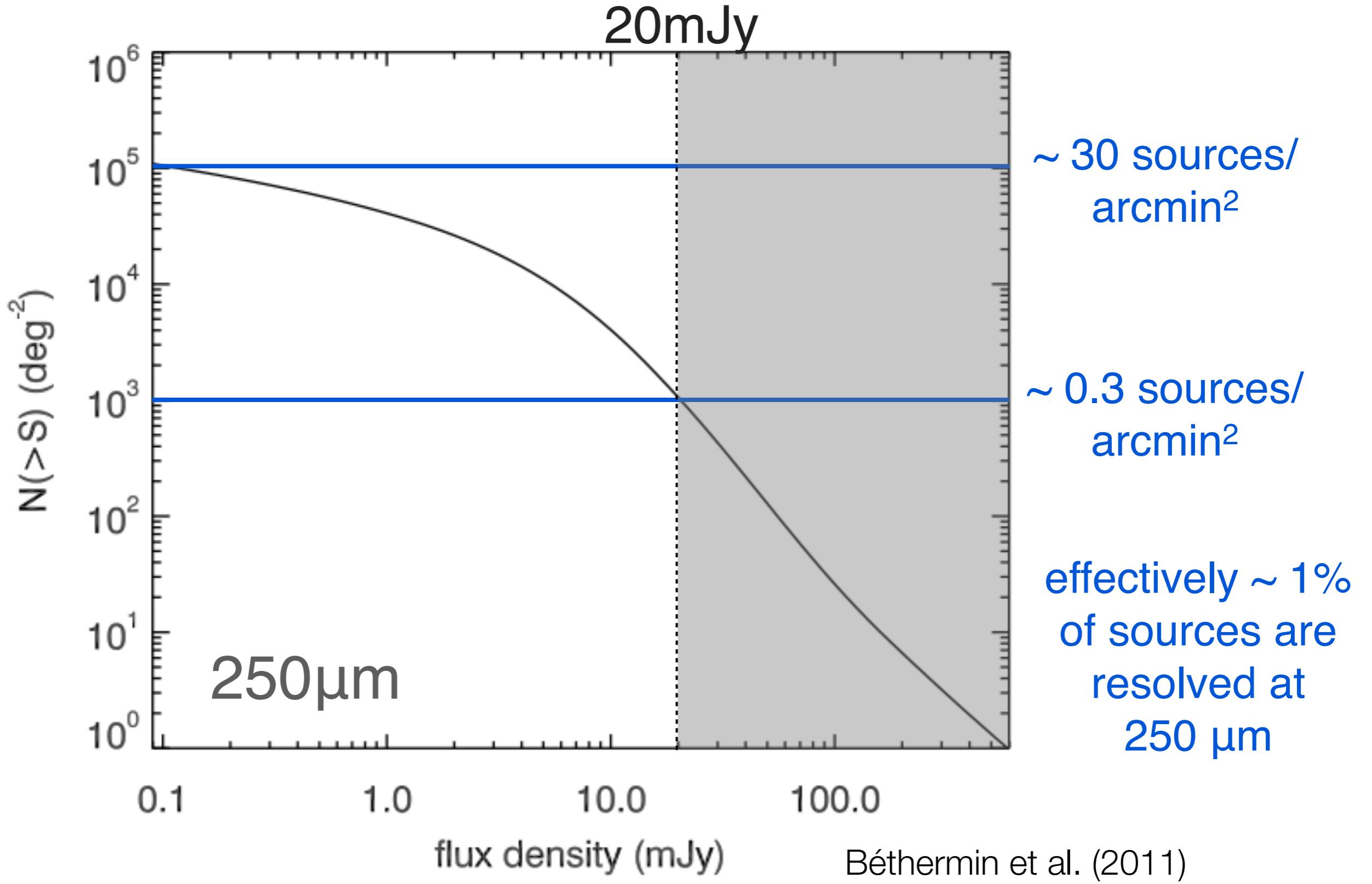
cumulative number counts

flux density (mJy)

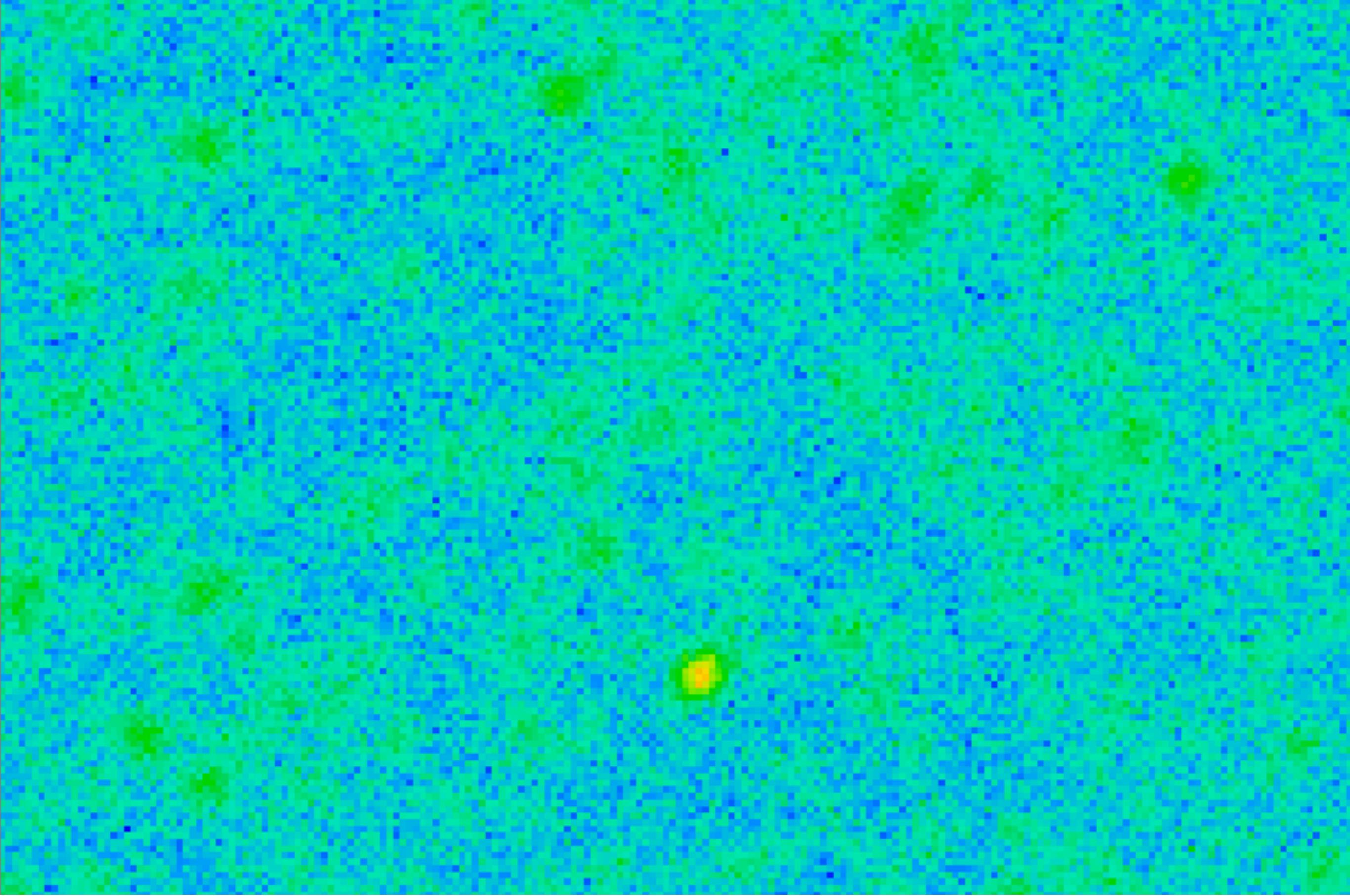
Béthermin et al. (2011)  
arXiv:1010.1150



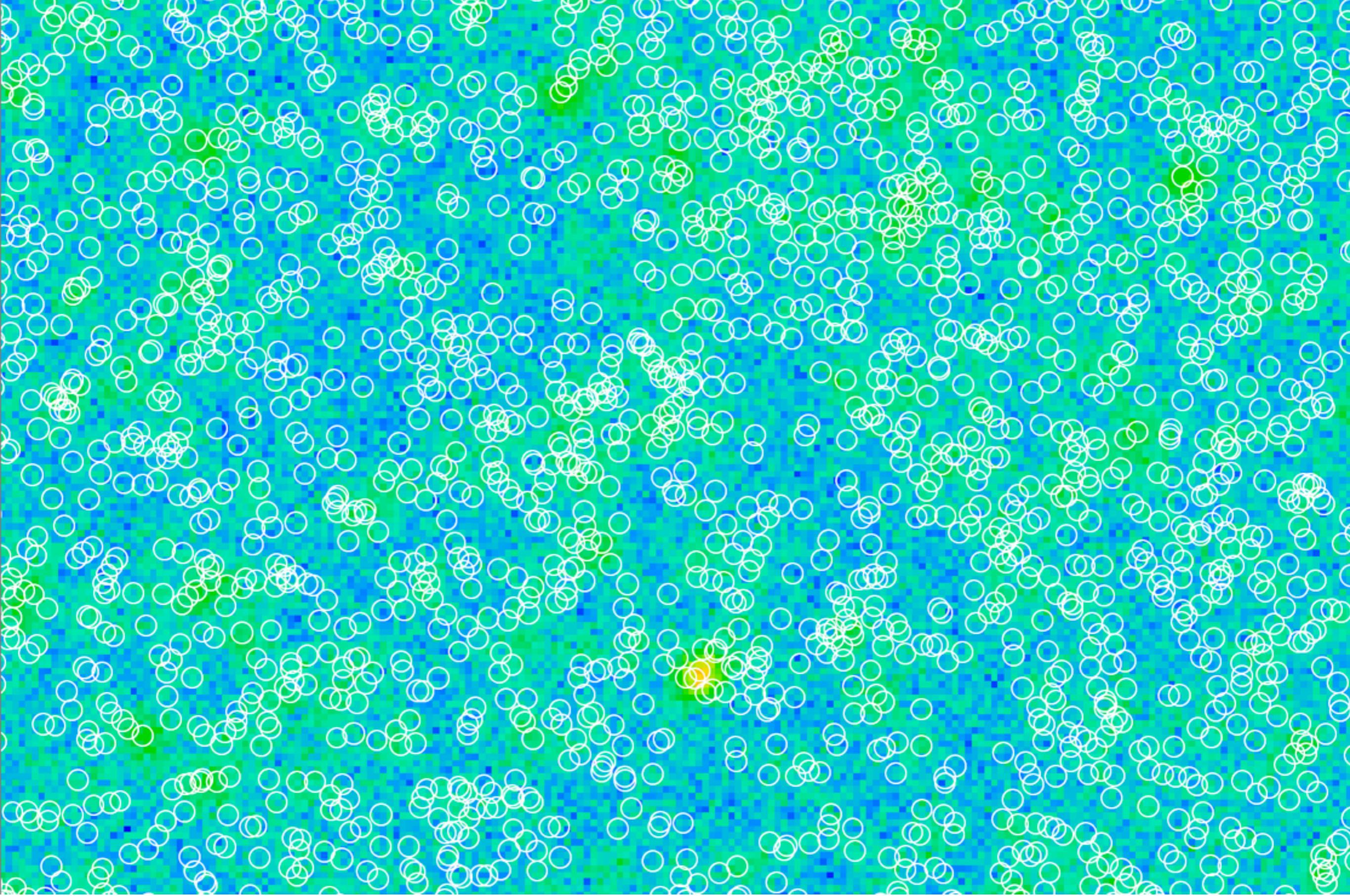
cumulative number counts



# cumulative number counts



# The Cosmic Infrared Background

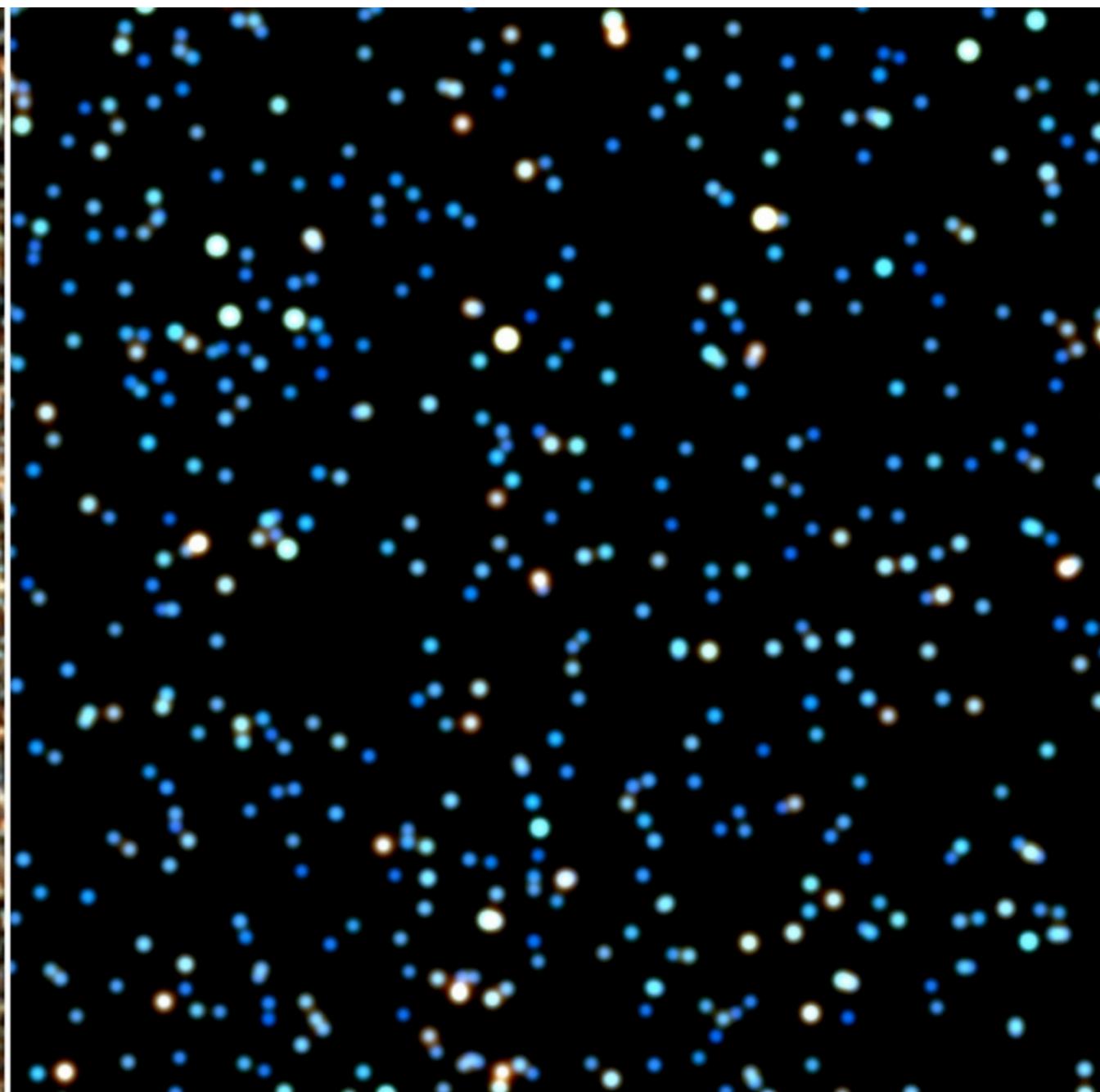


# The Cosmic Infrared Background

# CIB Results in 2 parts: Unresolved/Resolved

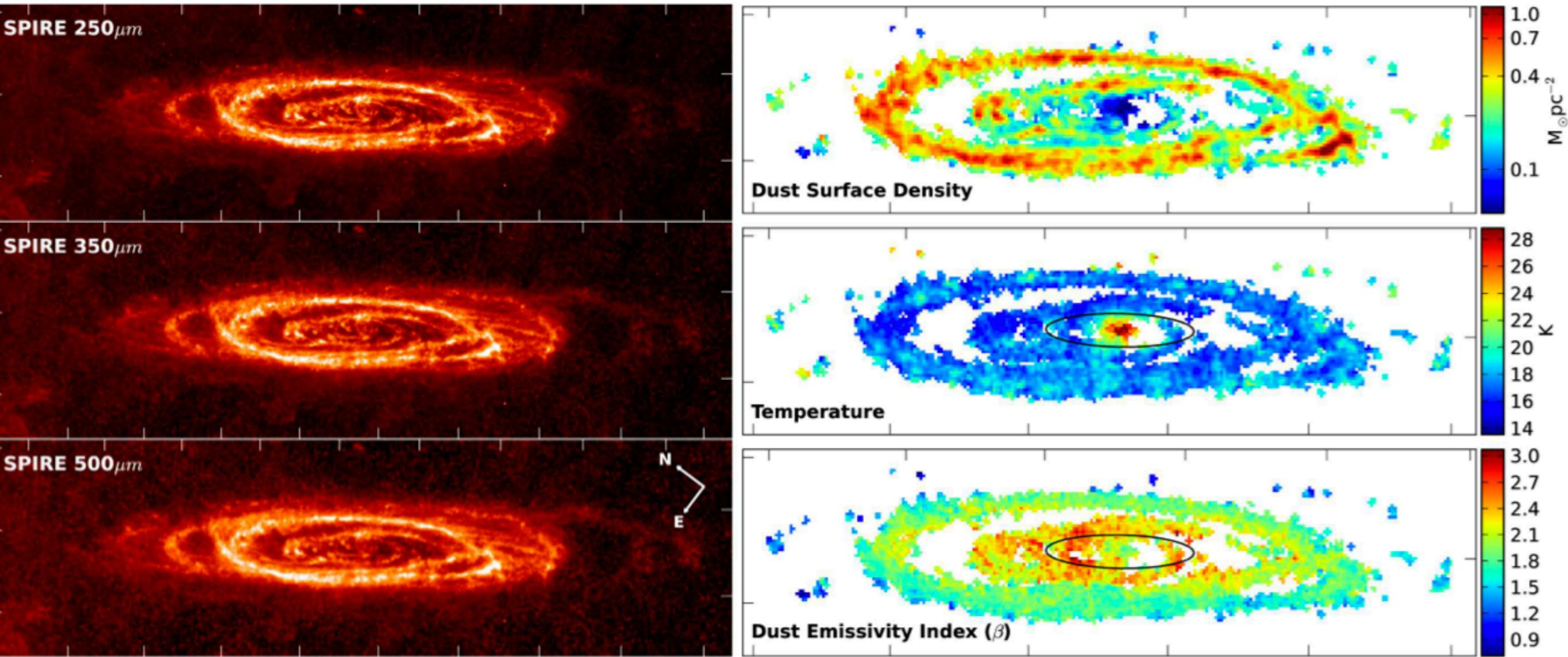


$S < 20 \text{ mJy} : 36,000/\text{deg}^2$



$S > 20 \text{ mJy} : 1,200/\text{deg}^2$

# Local Resolved Galaxies

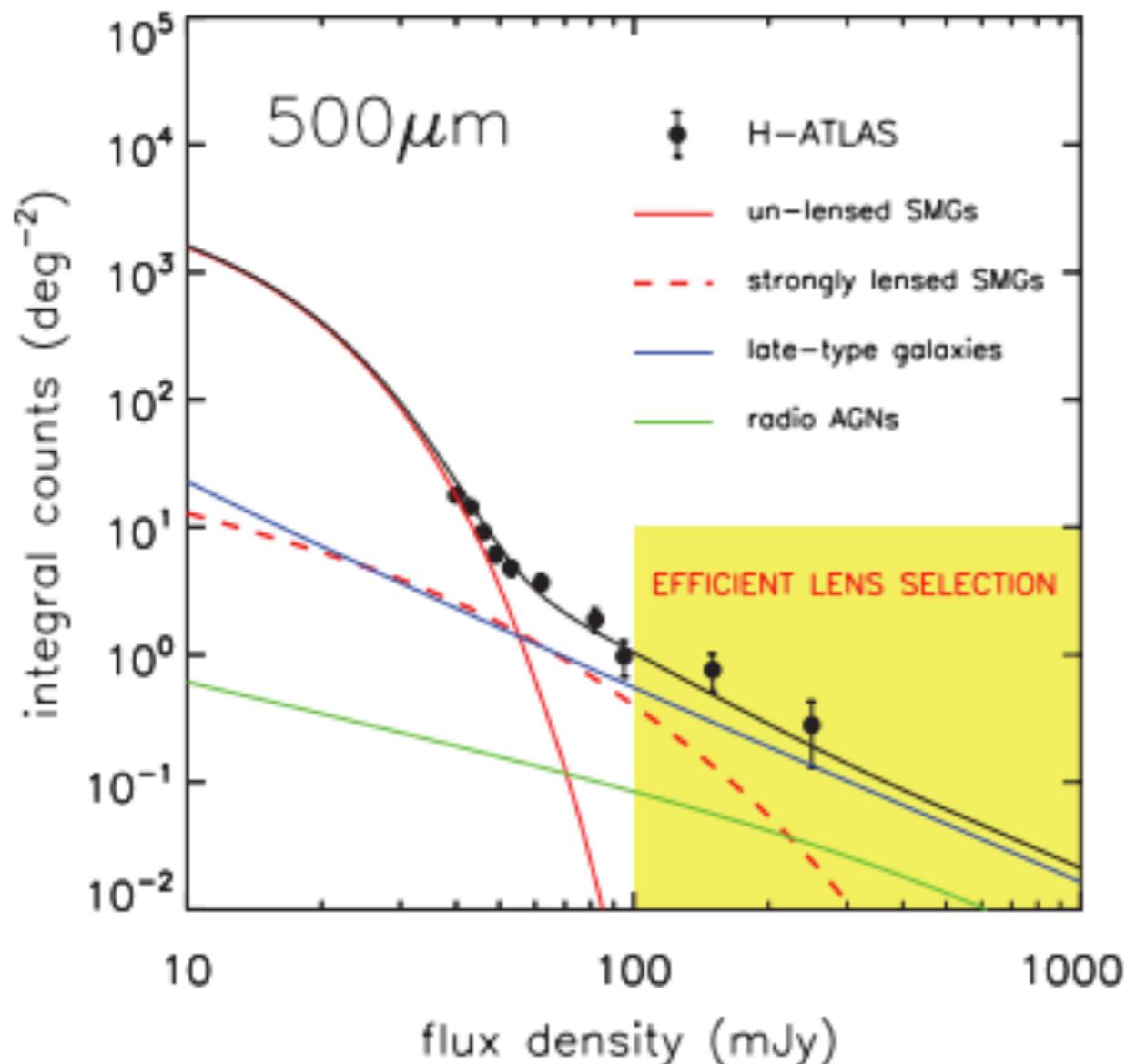


Smith++ 2012, *ApJ*, 756, 40

The Herschel Exploitation of Local Galaxy Andromeda (HELGA) II

See also: Mentuch Cooper++2013, Foyle++2013, SINGS: Wilson++2012

# Lensed Sources



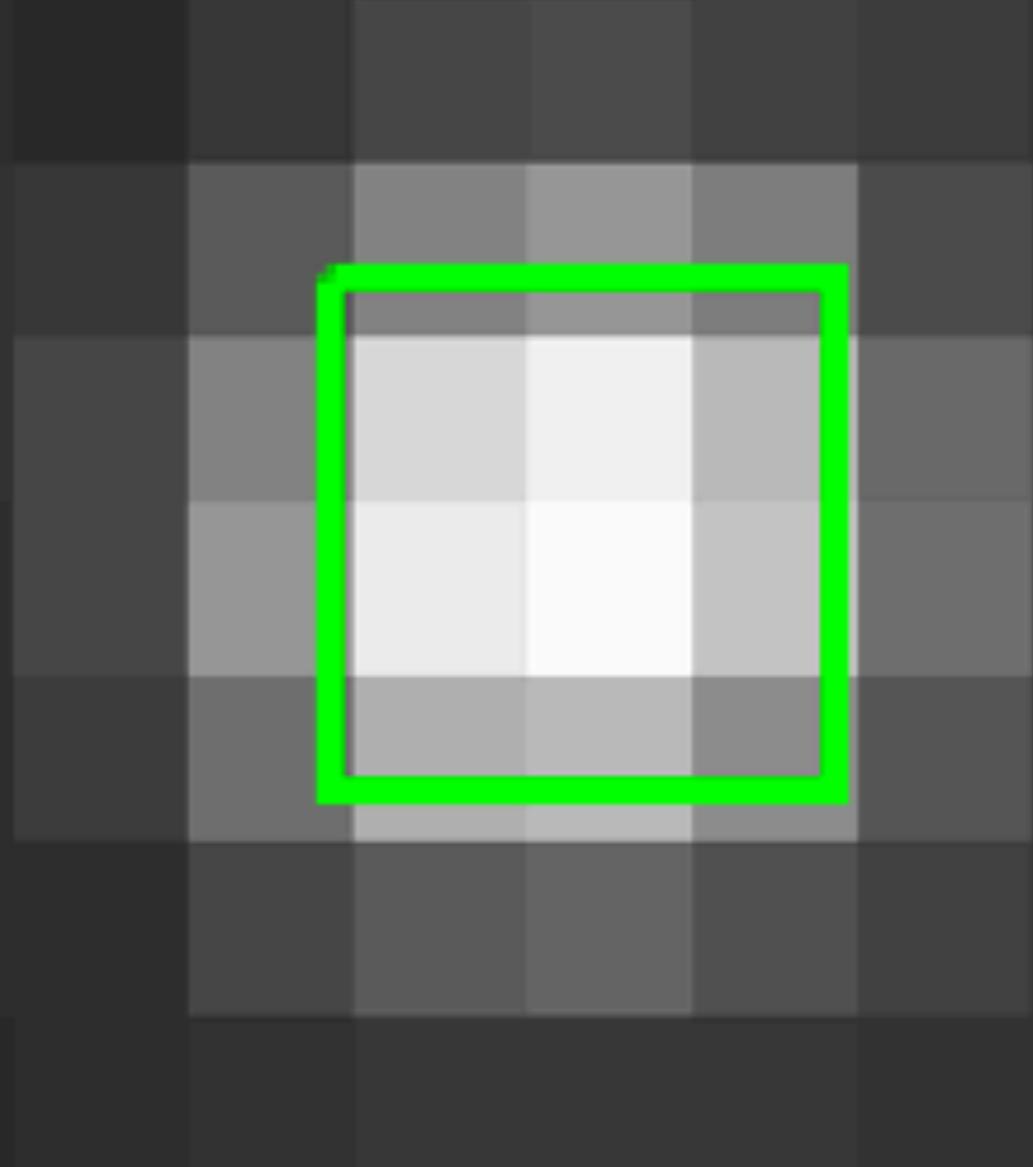
- Sources with flux density  $S > 100\text{mJy}$  at 500 $\mu$ m have high probability of being lensed

Negrello++ 2010

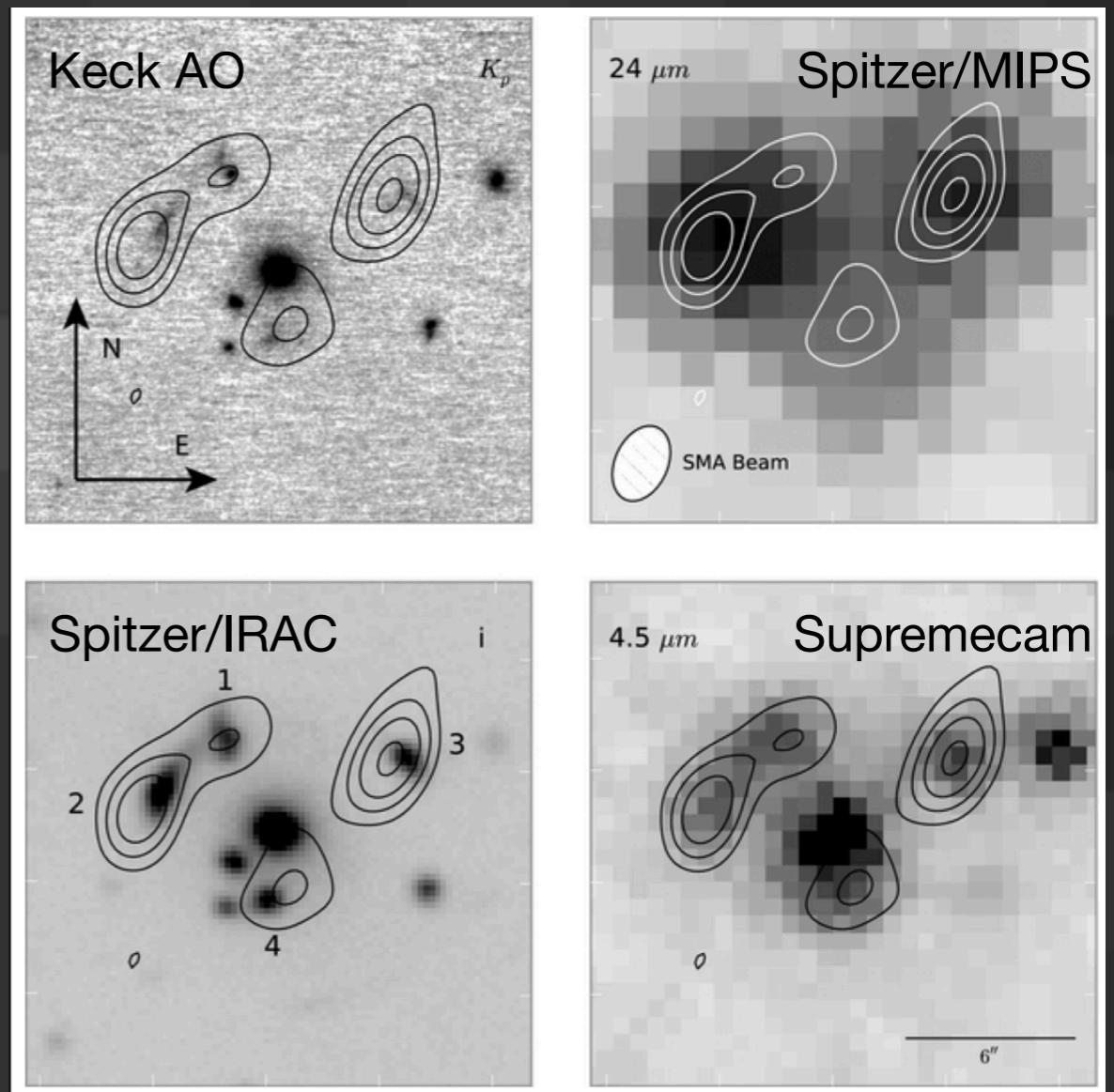
The Detection of a Population of Submillimeter-Bright, Strongly Lensed Galaxies. Science 330, 800.

# Lensed Sources

SPIRE 250 $\mu$ m (6" pixels)



$z=2.97$  from spectroscopic follow-up



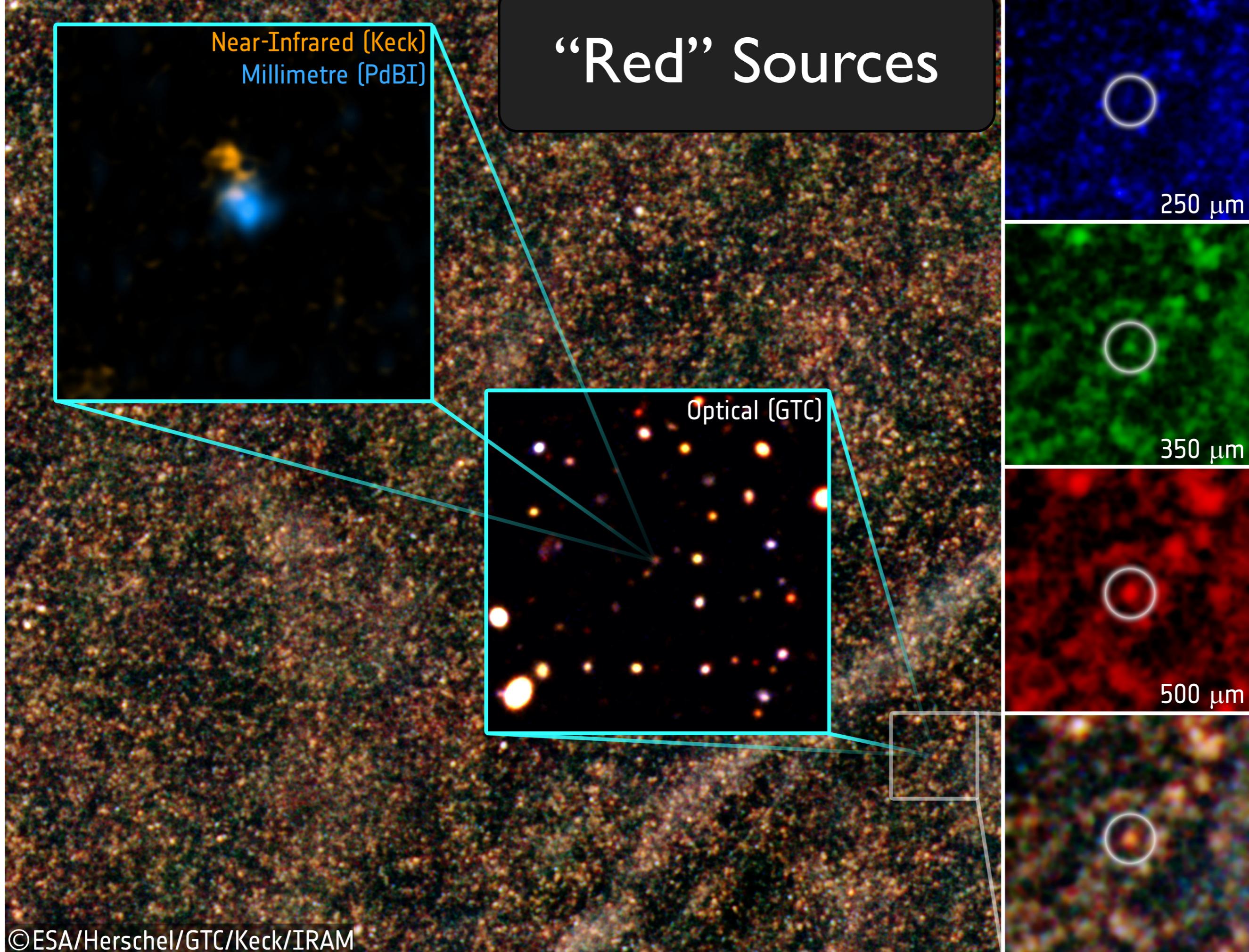
Contours From Submillimeter Array (SMA)

Conley++ 2011, DISCOVERY OF A MULTIPLY LENSED SUBMILLIMETER GALAXY  
IN EARLY HerMES *HERSCHEL/SPIRE* DATA  
also see: Negrello++ 2010, Gonzalez-Nuevo++ 2012, Wardlow++ 2012, Fu++ 2013

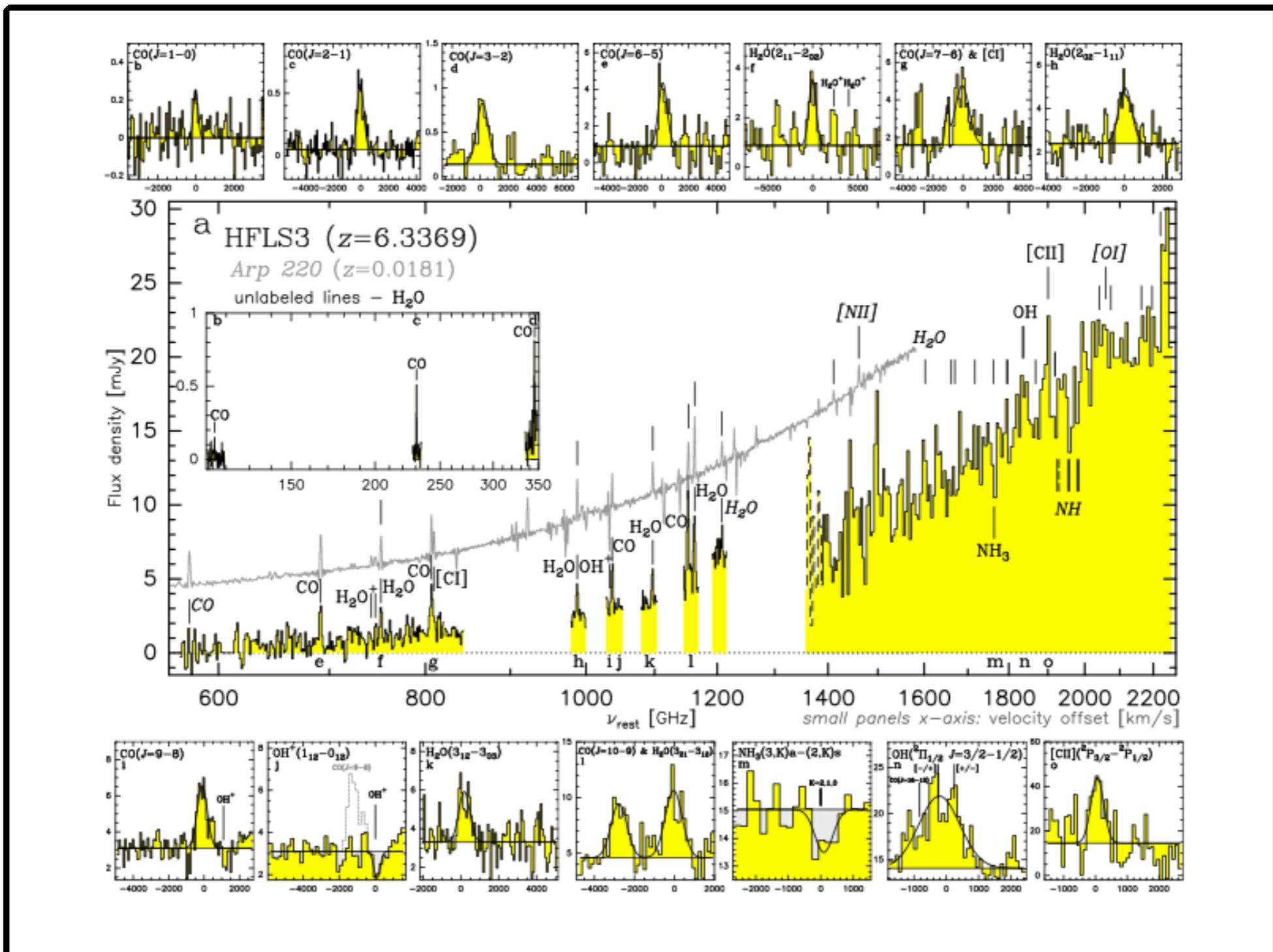
Near-Infrared (Keck)  
Millimetre (PdBI)

# “Red” Sources

Optical (GTC)



# $z=6.337$ “red” source

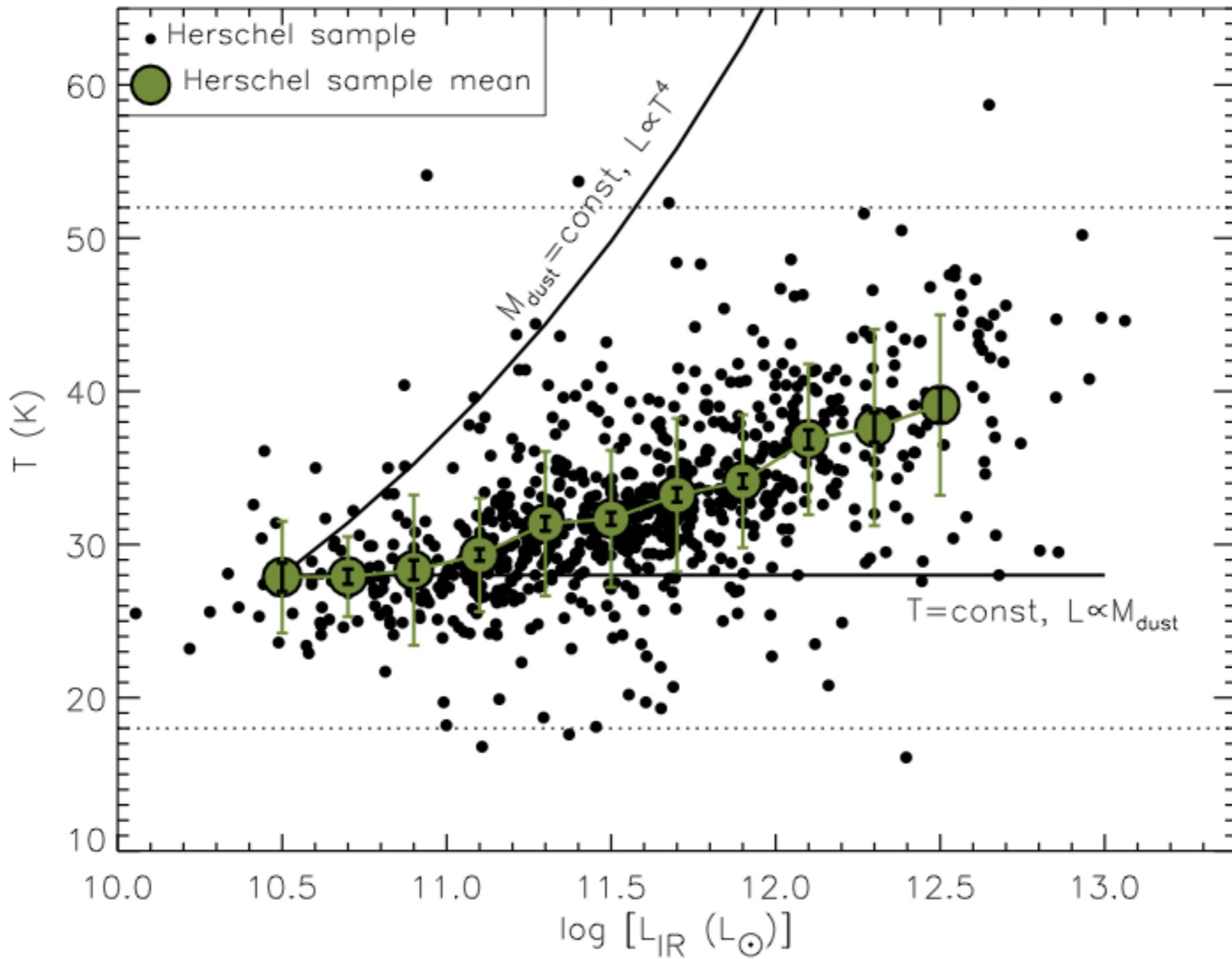


Riechers++ 2013, *Nature*, 496(7), pp.329–333

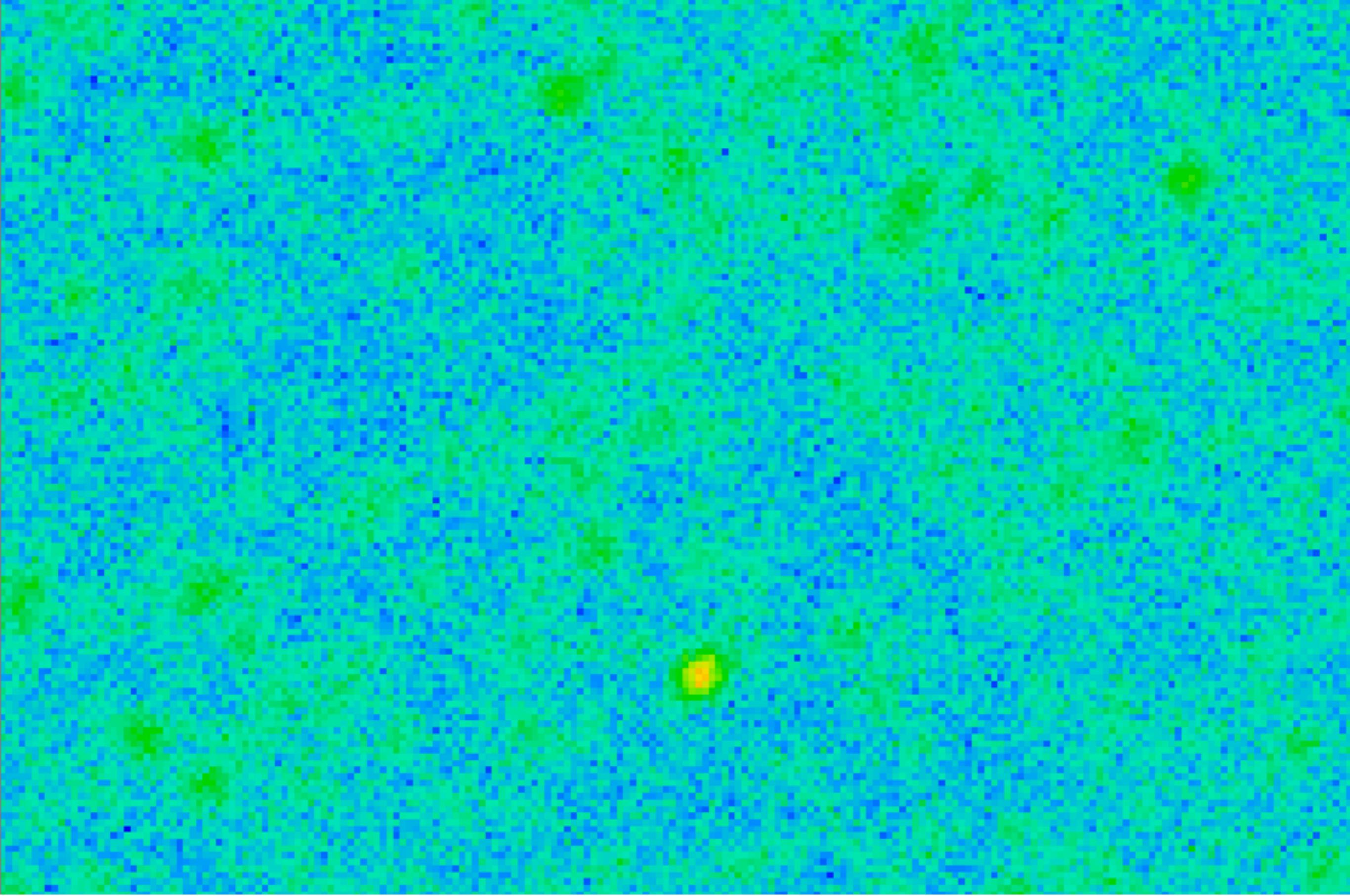
A dust-obscured massive maximum-starburst galaxy at a redshift of 6.34

See also: Dowell++ in prep., Gill++ in prep.

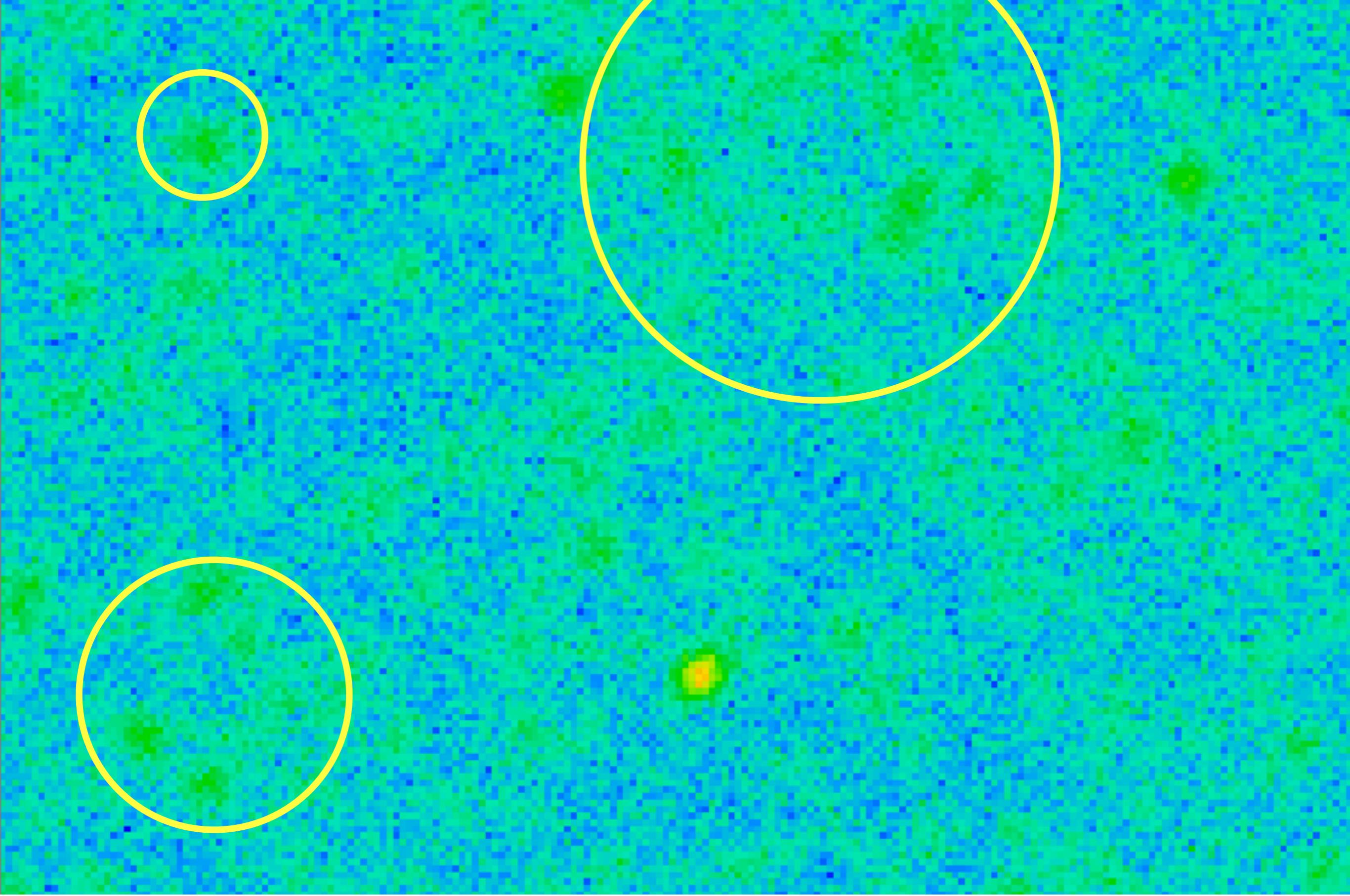
# SEDs



Symeonidis++ 2013, The *Herschel* Census of Infrared SEDs through cosmic time, *MNRAS*, 431(3), pp. 2317–2340.  
See also: Casey++ 2012a,b, Magdis++ 2010, 2012, Canalog++ 2013

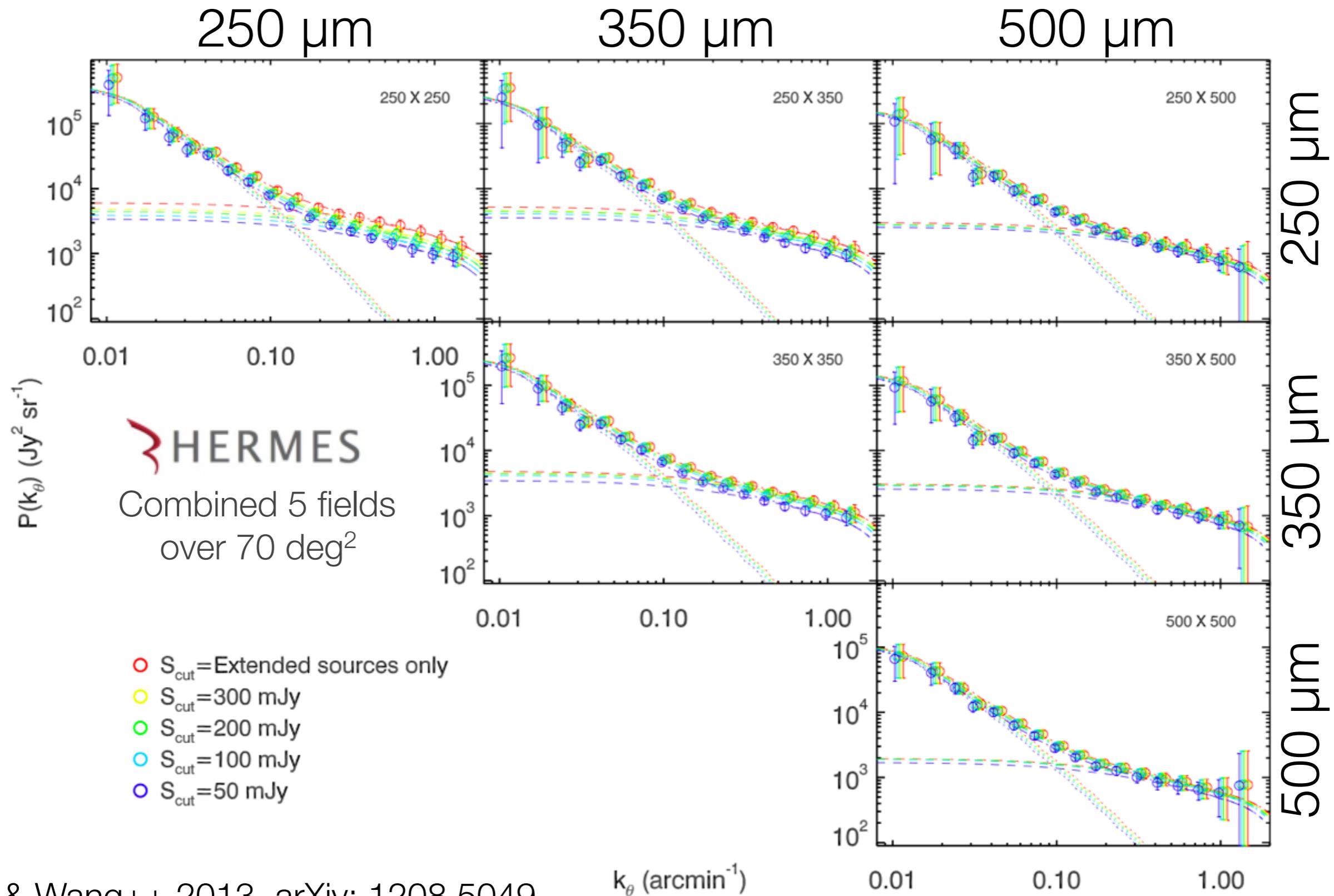


ii) The Unresolved Background

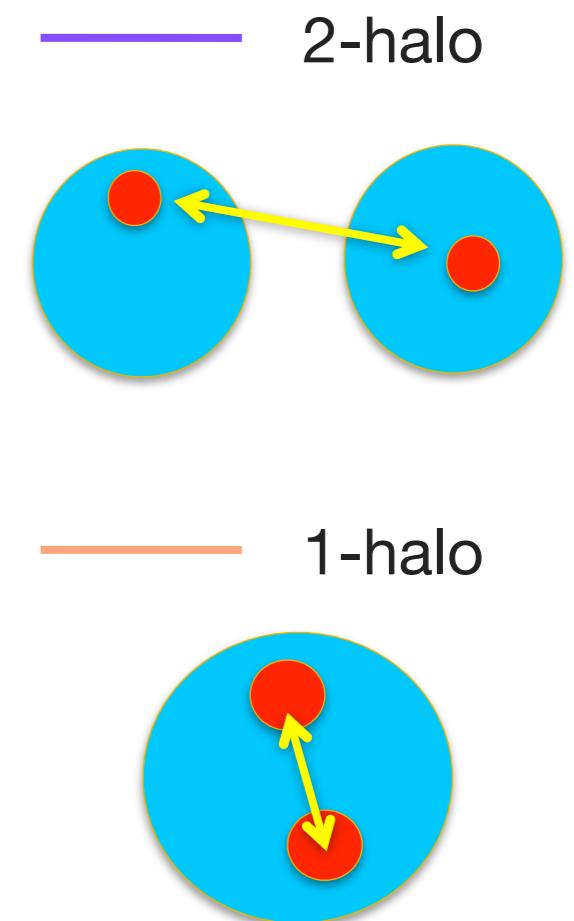
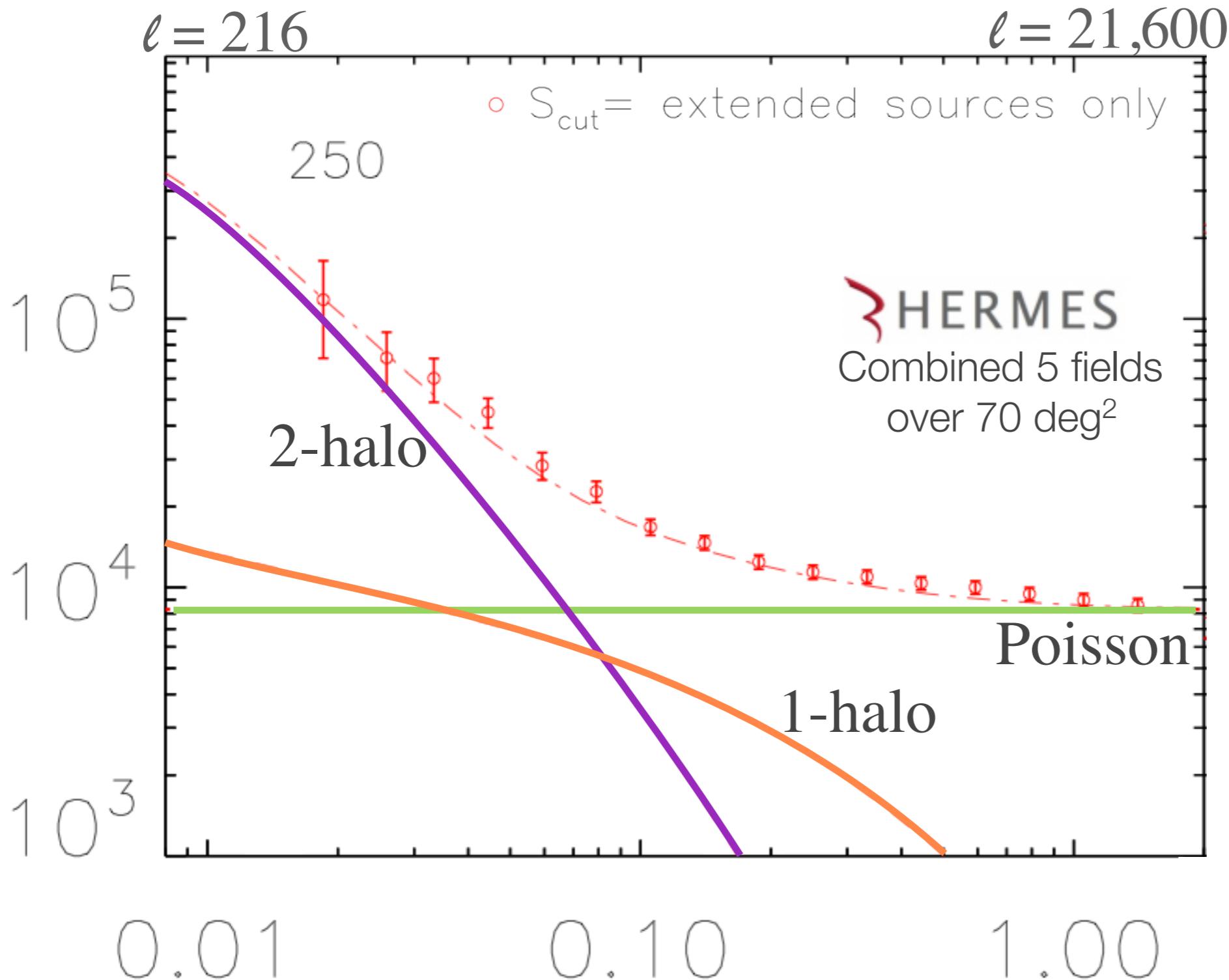


CIB Anisotropies (CIBA)

# CIBA power spectra

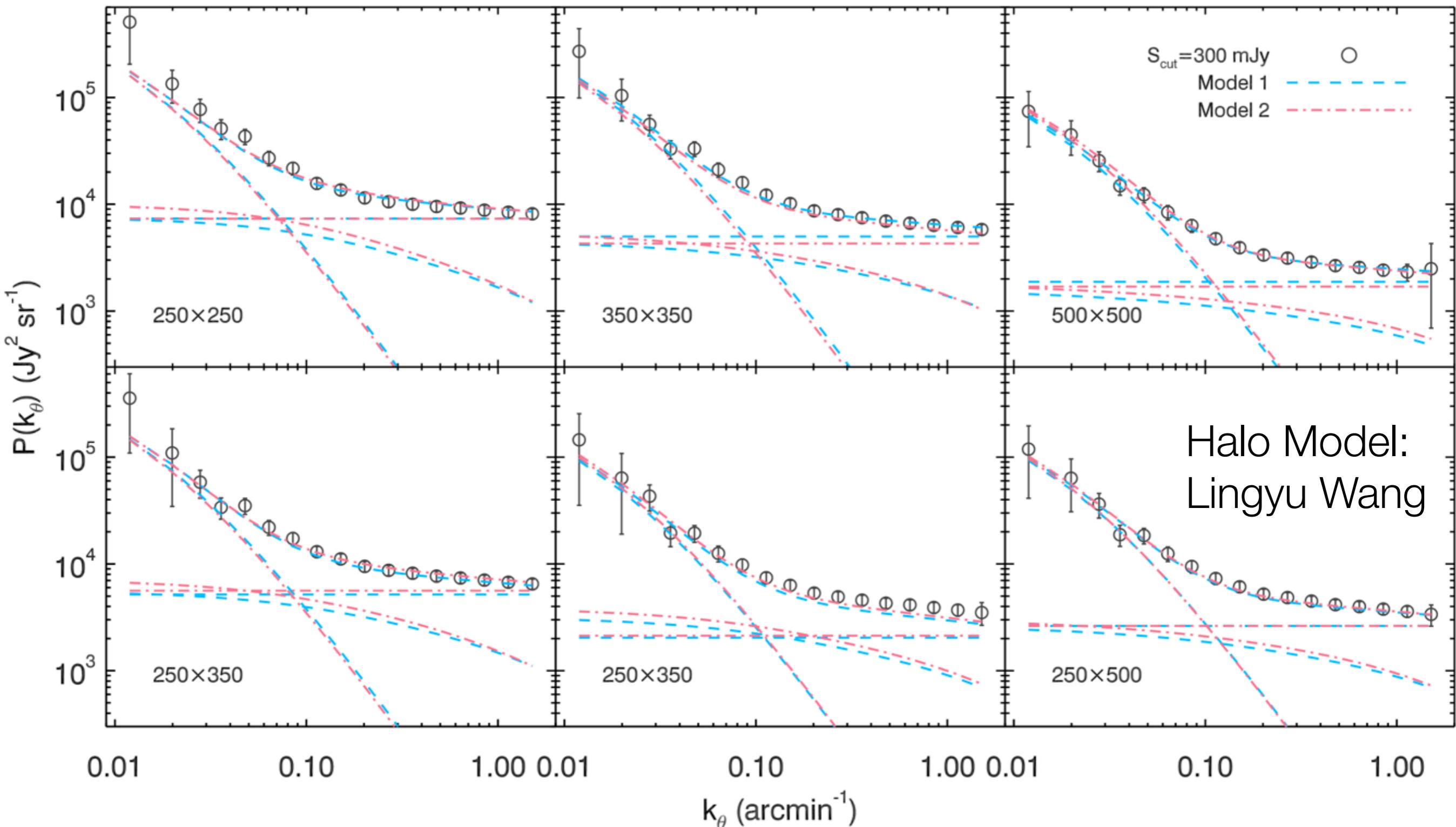


# Clustering of DSFGs



Viero & Wang++ 2012, arXiv: 1208.5049  
See also: Bethermin++ 2013, Wang++ 2013

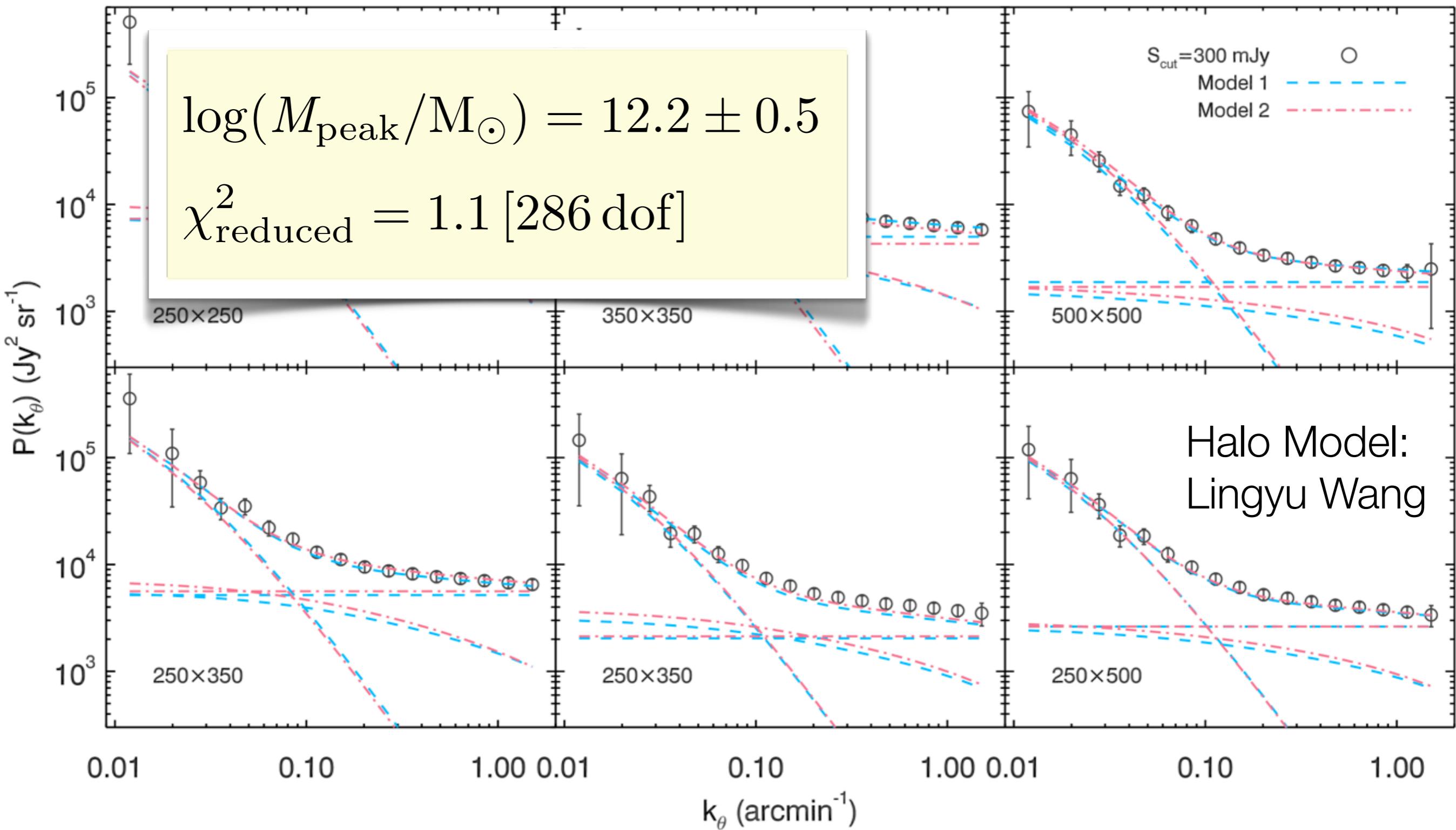
# Best-Fit Halo Model



Viero & Wang++ 2012, arXiv: 1208.5049

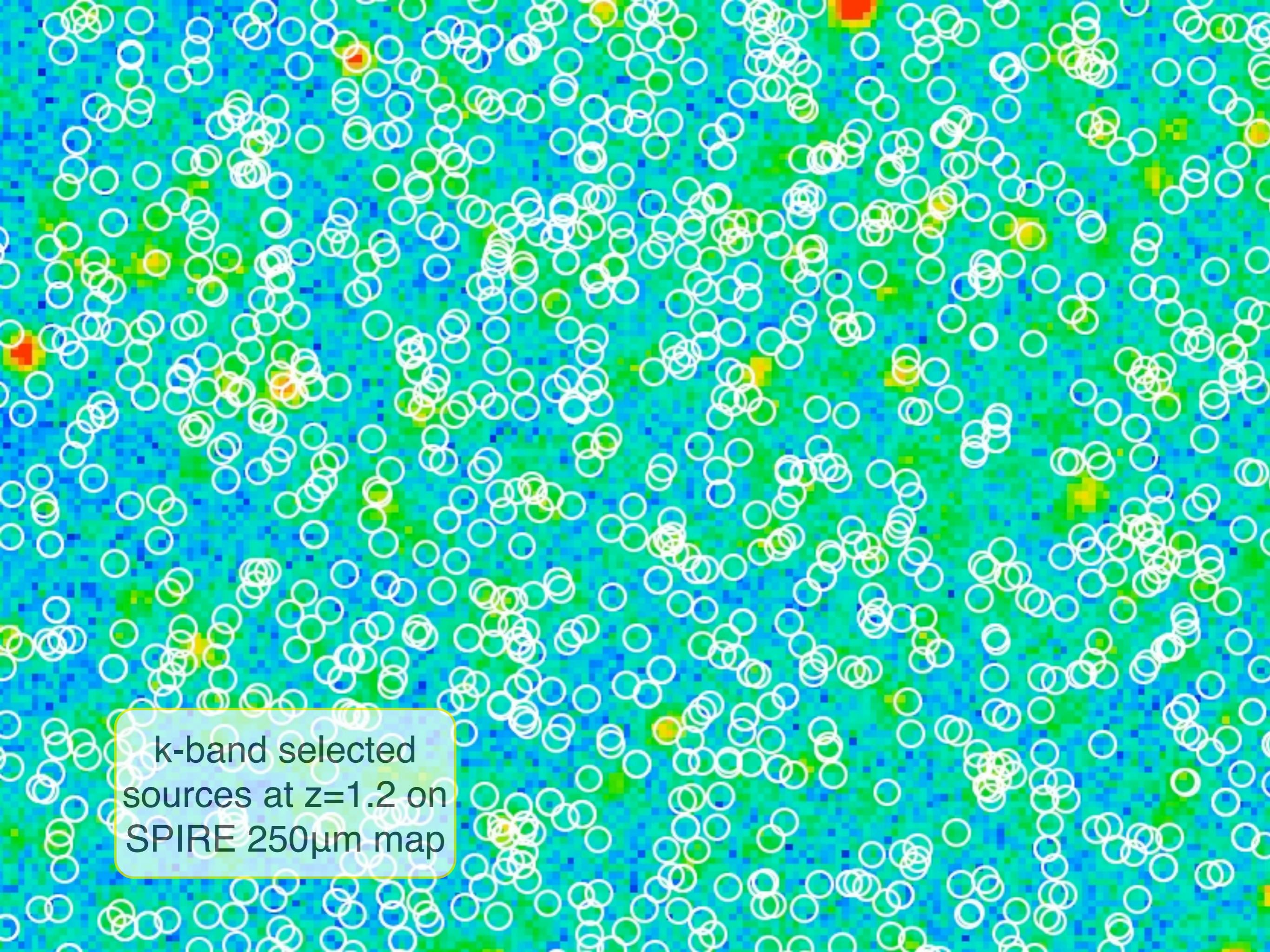
See also: Bethermin++ 2013, Wang++ 2013

# Best-Fit Halo Model

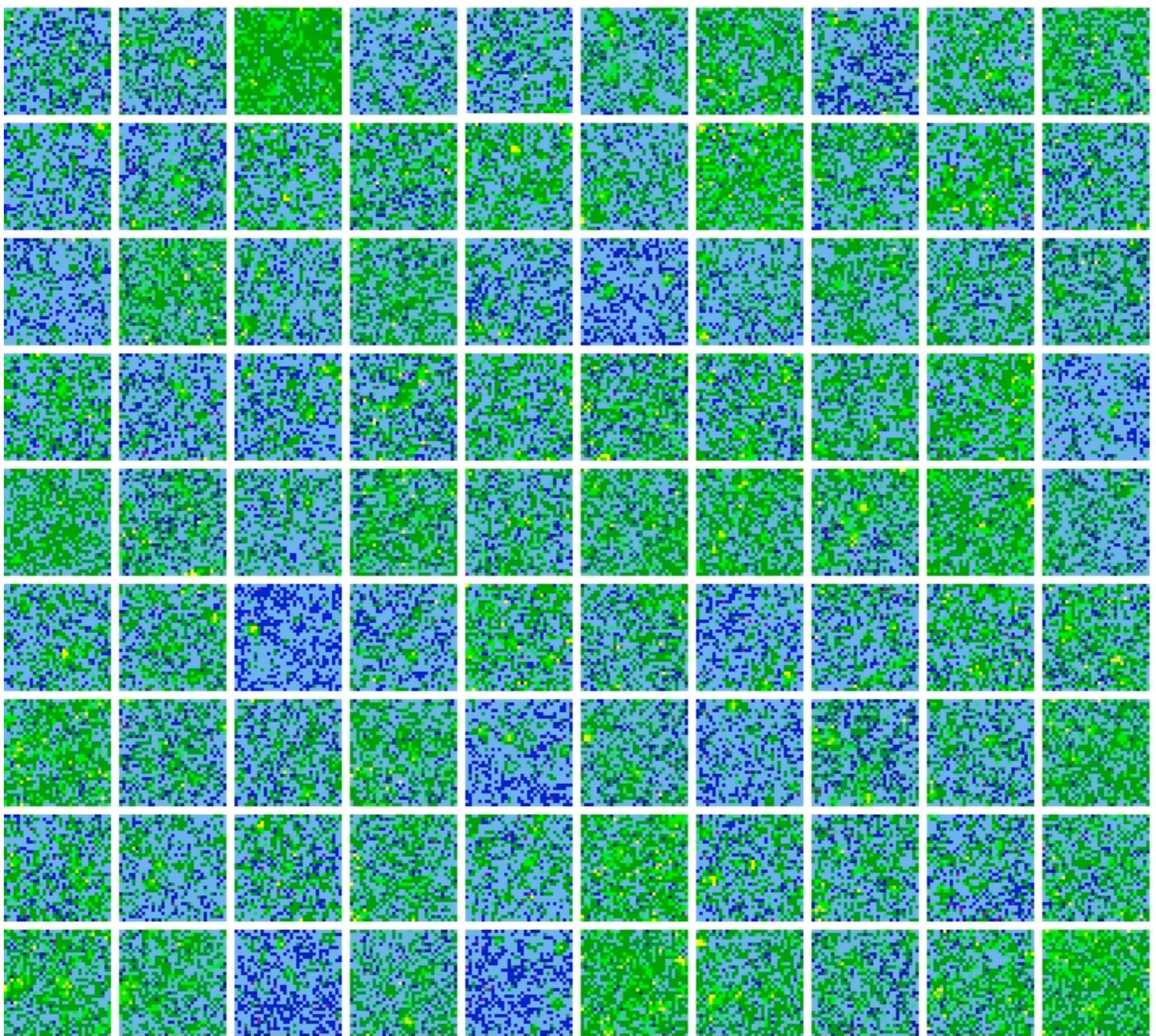


Viero & Wang++ 2012, arXiv: 1208.5049

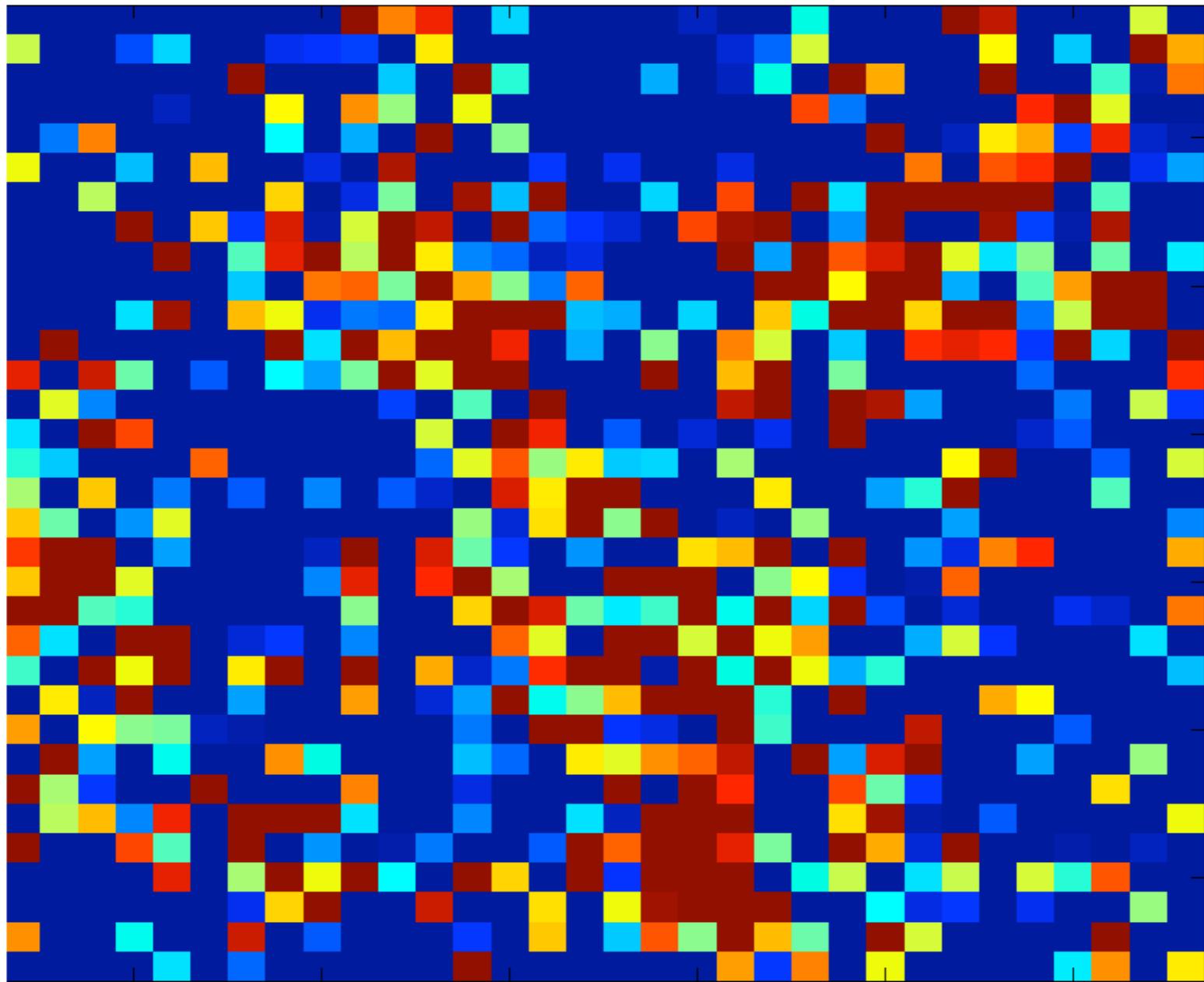
See also: Bethermin++ 2013, Wang++ 2013



k-band selected  
sources at  $z=1.2$  on  
SPIRE 250 $\mu\text{m}$  map



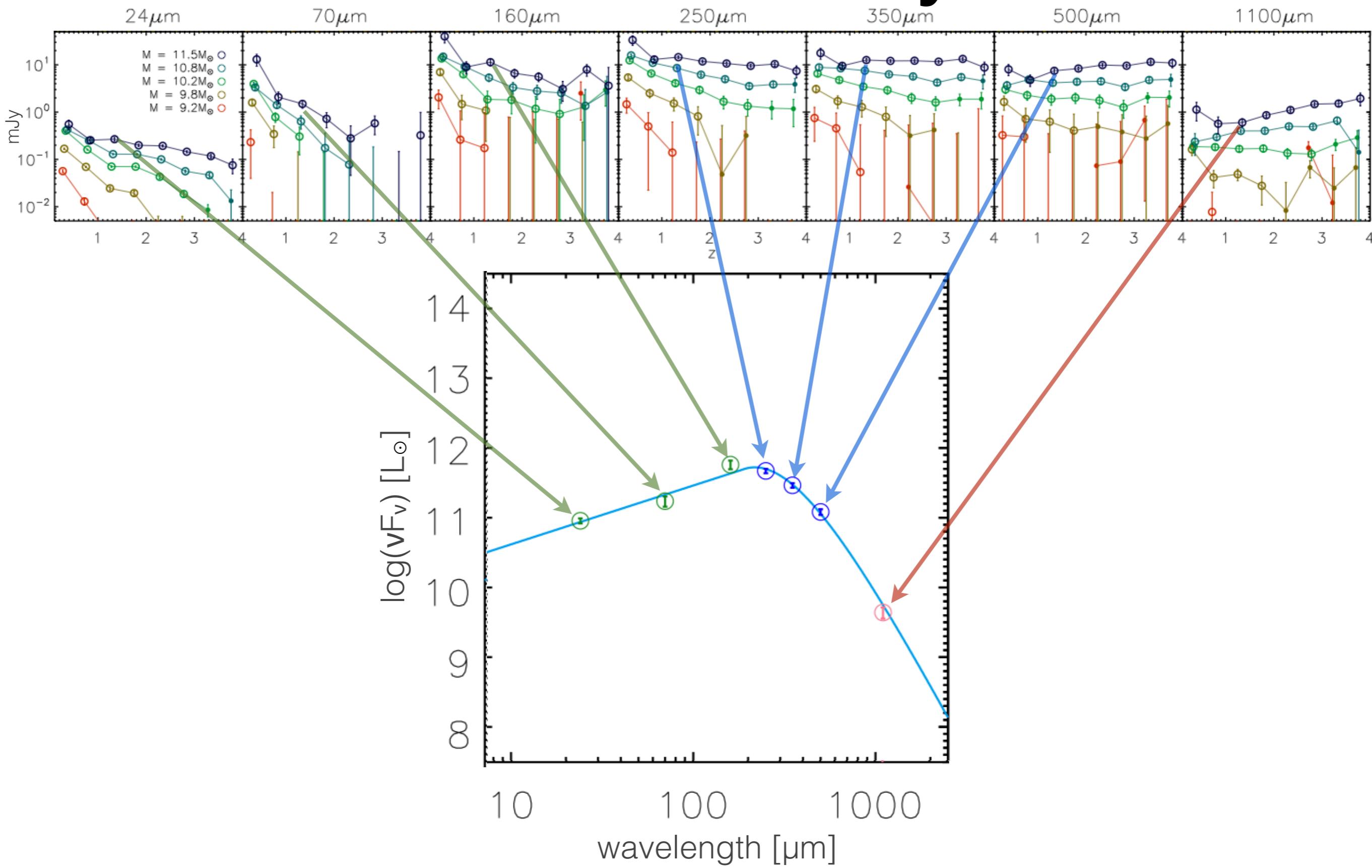
stacking



Phil Korngut (Caltech)

stacking

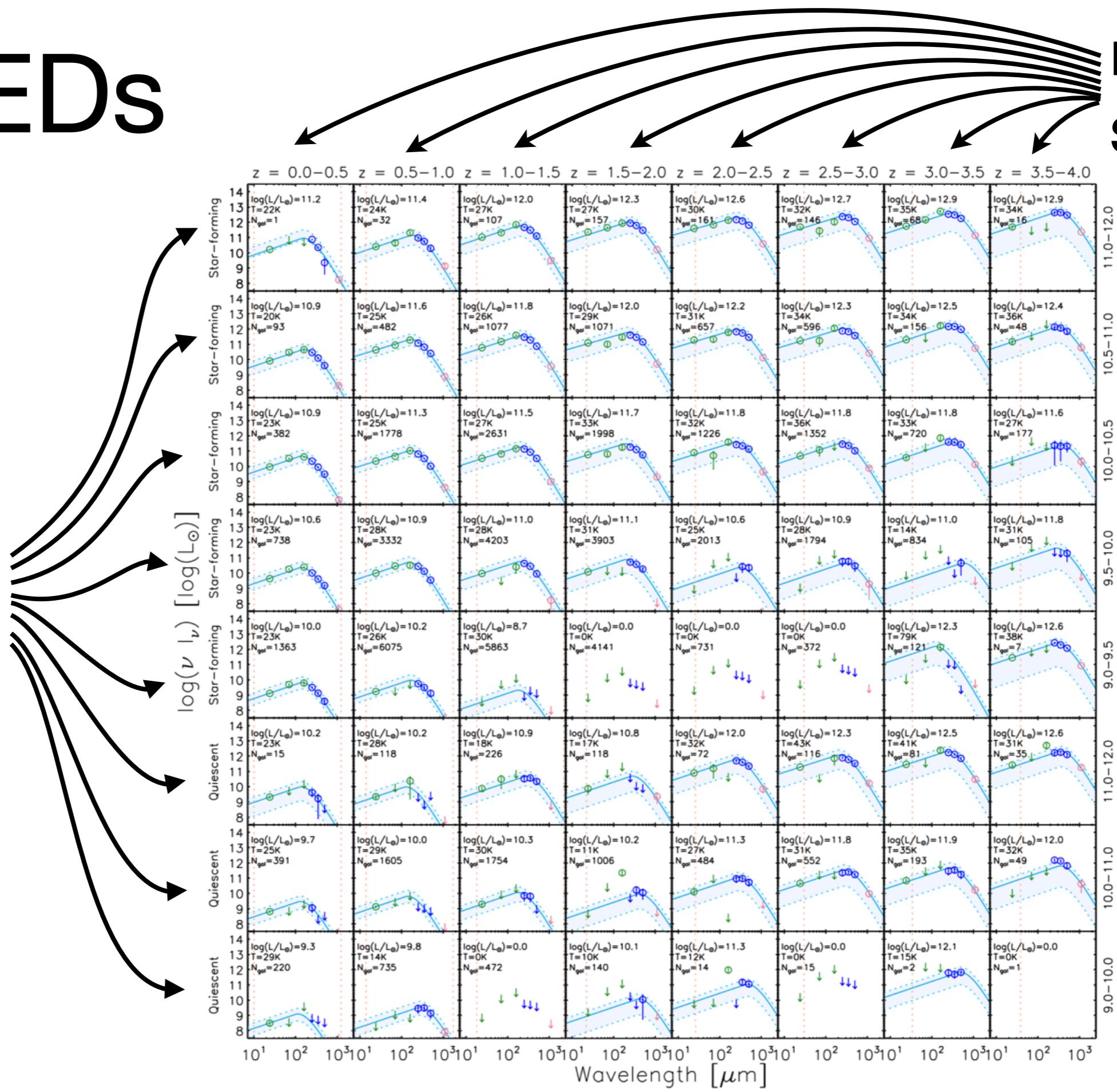
# stacked flux density



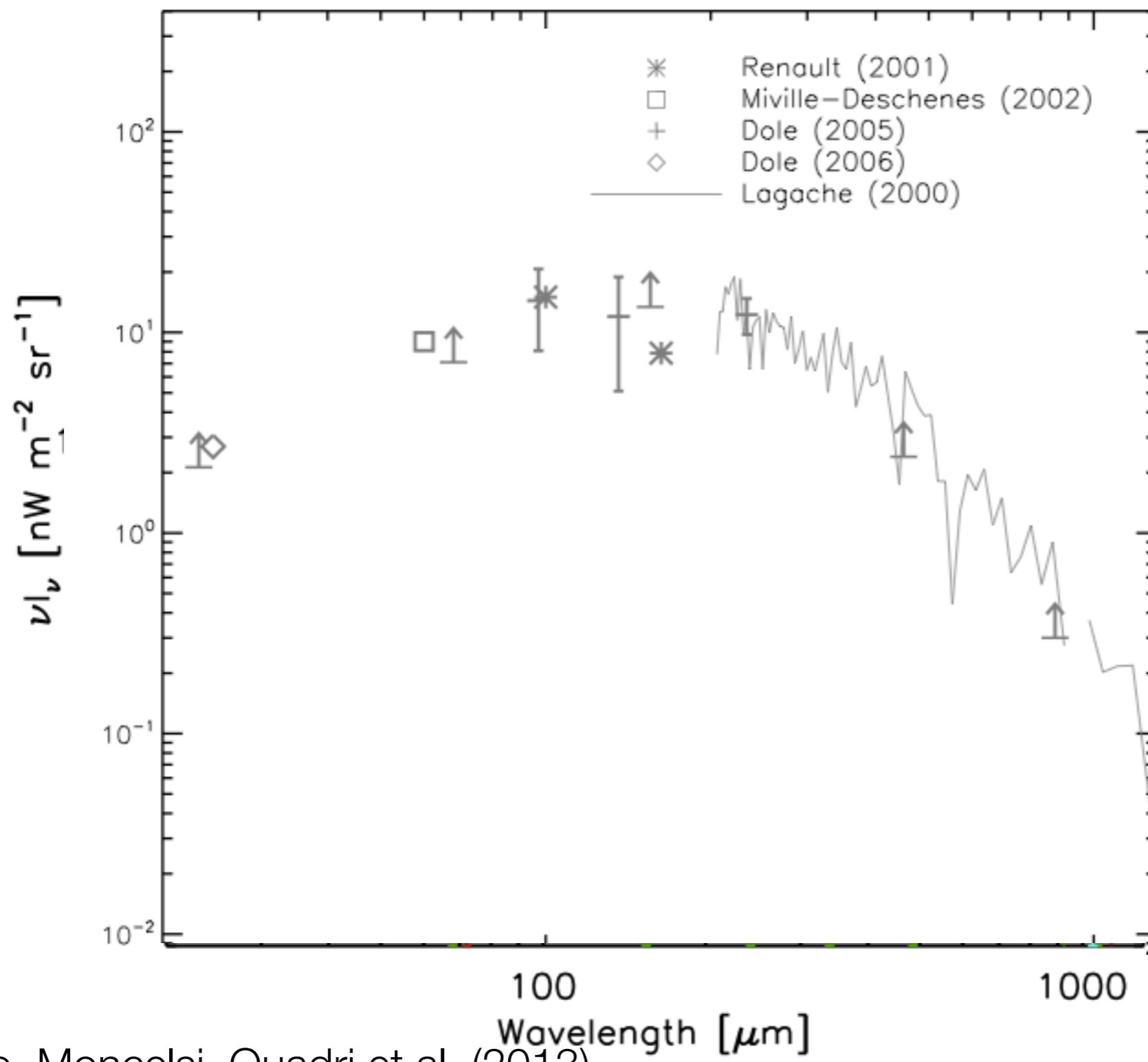
# SEDS

redshift  
slices

mass  
slices



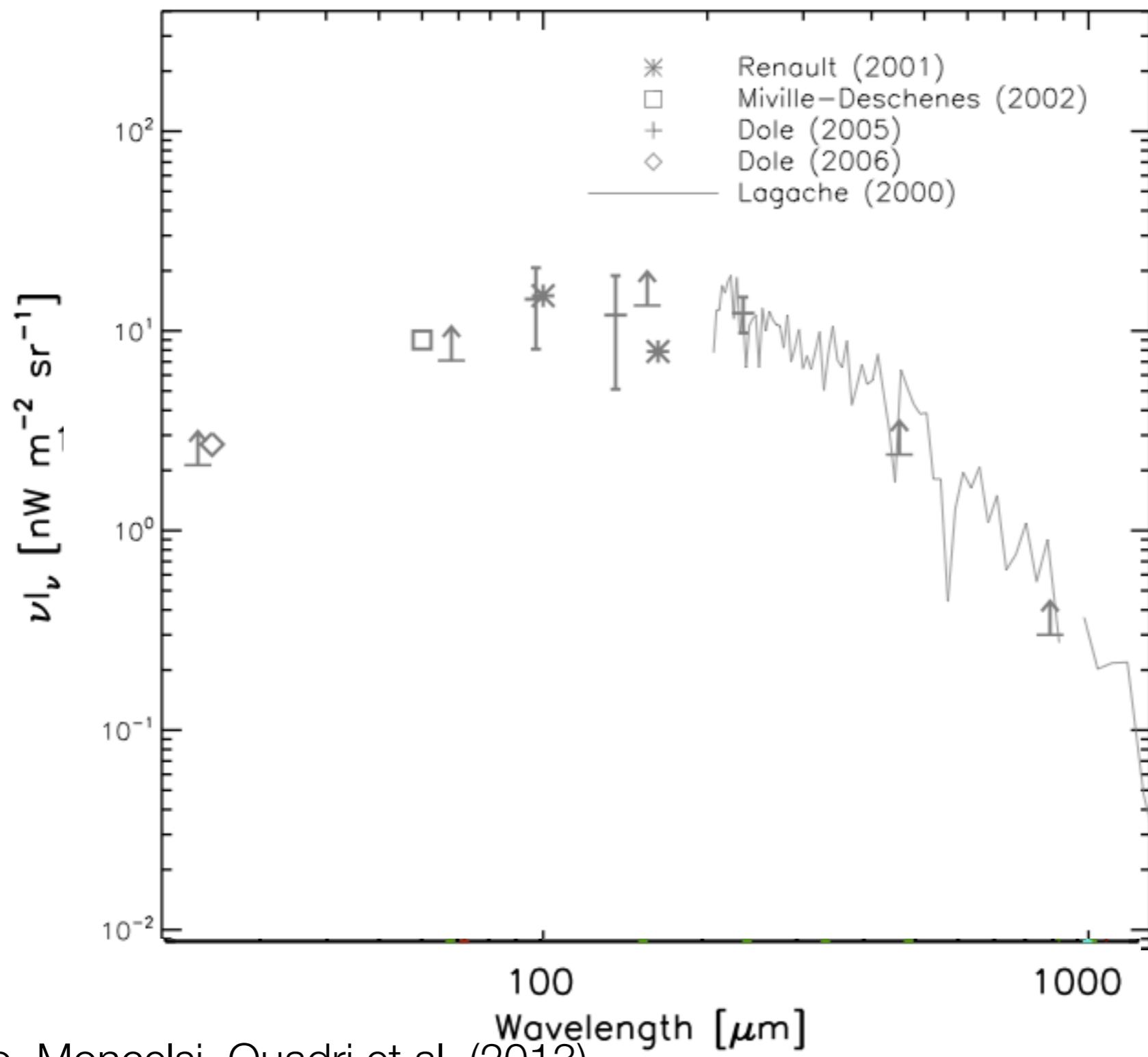
# stacked CIB



Viero, Moncelsi, Quadri et al. (2013)

arXiv:1304.0446

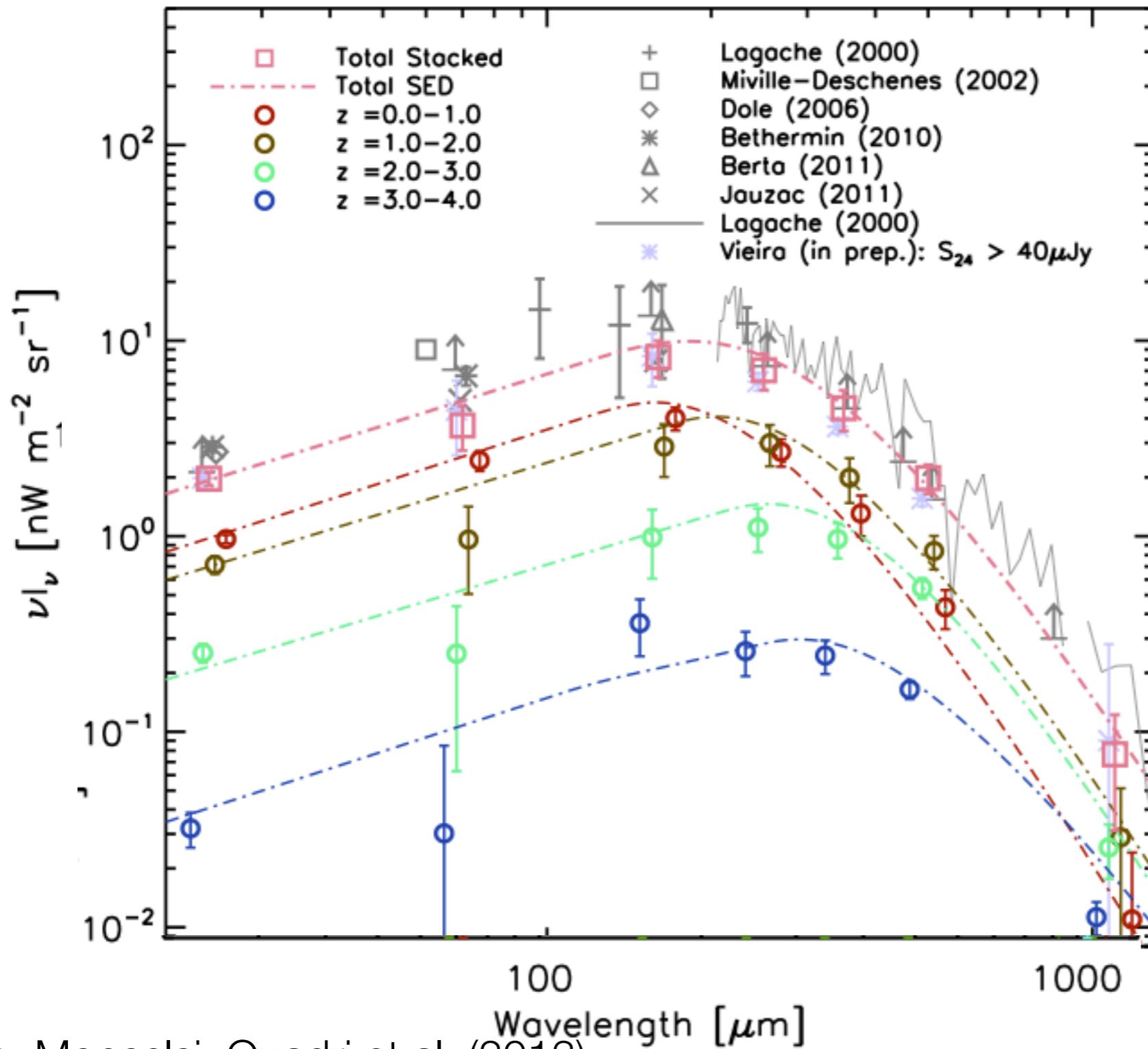
# stacked CIB



Viero, Moncelsi, Quadri et al. (2013)

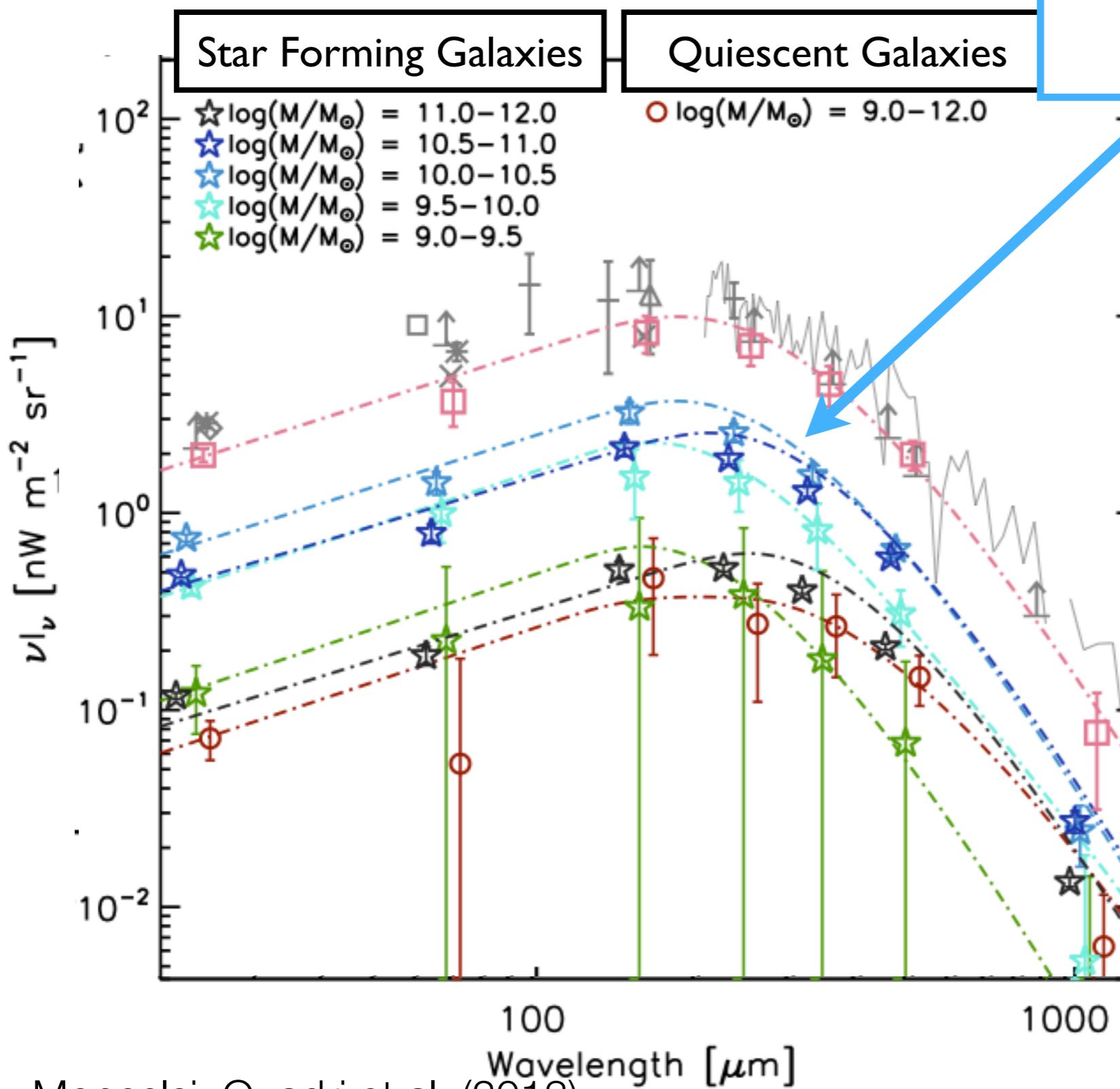
arXiv:1304.0446

# stacked CIB



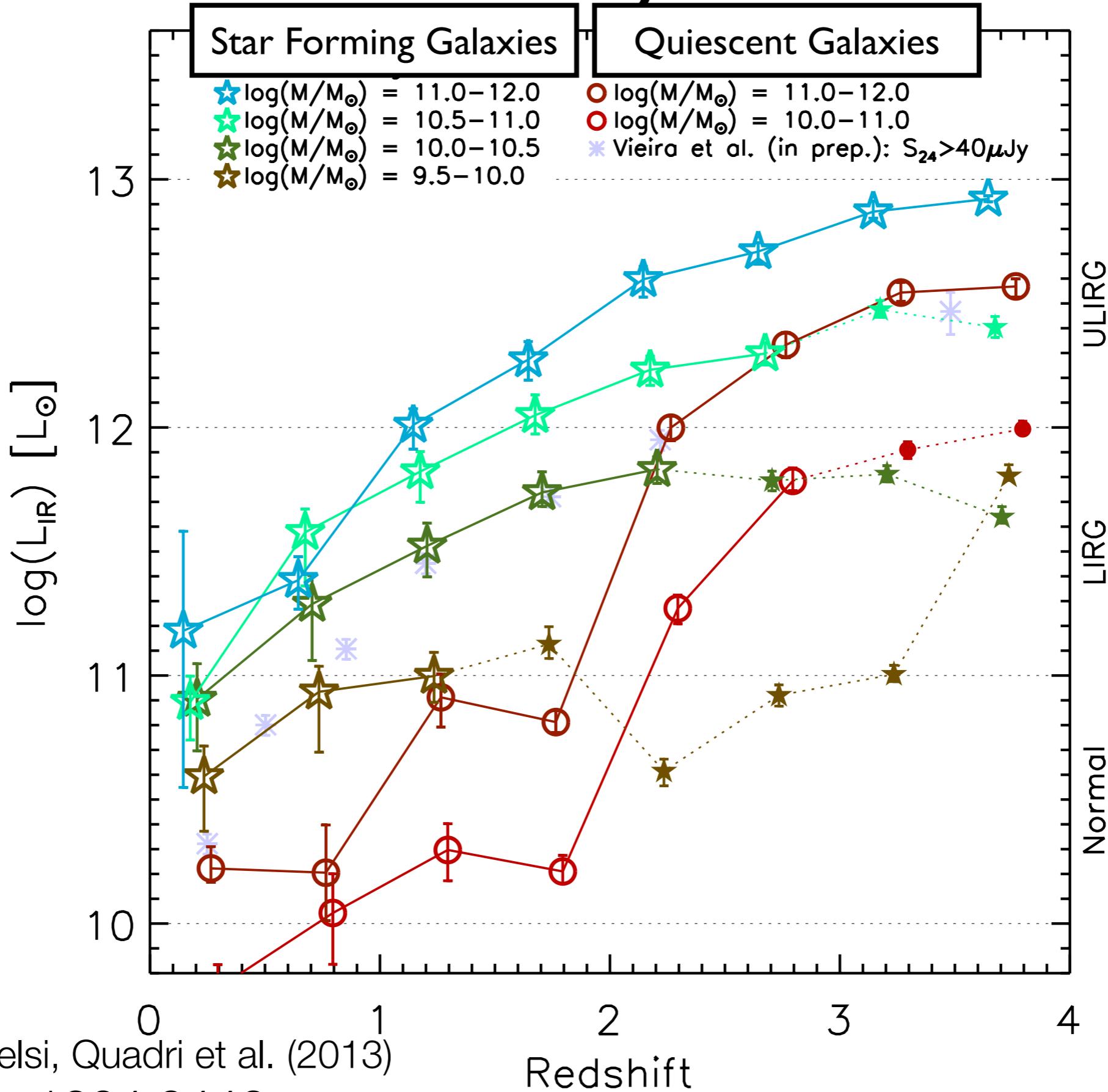
~80% at SPIRE  
wavelengths

# stacked CIB



$z \sim 1.5 - 2.5$   
 $\log(M/M_\odot \sim 10 - 11)$

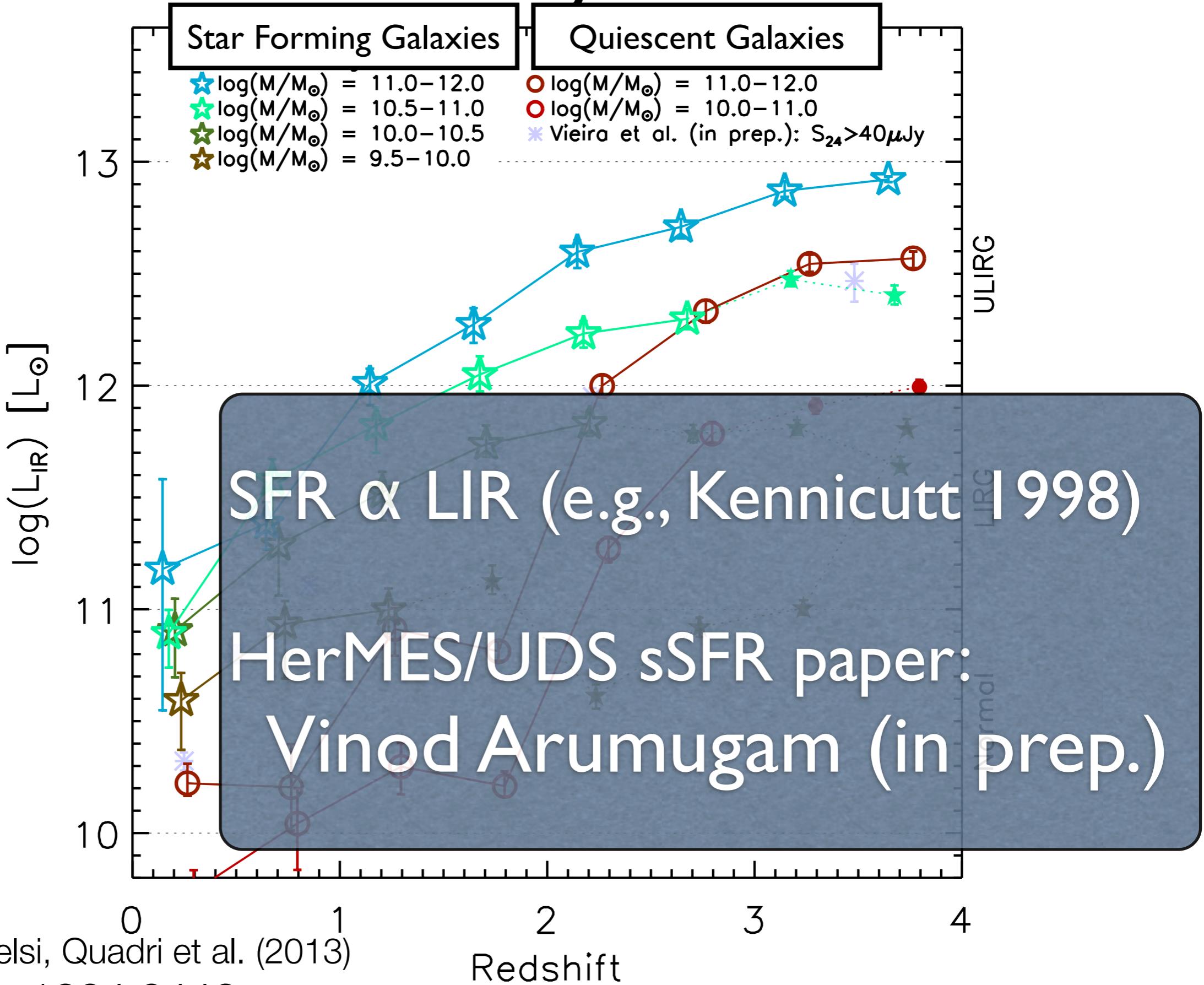
# Infrared Luminosity



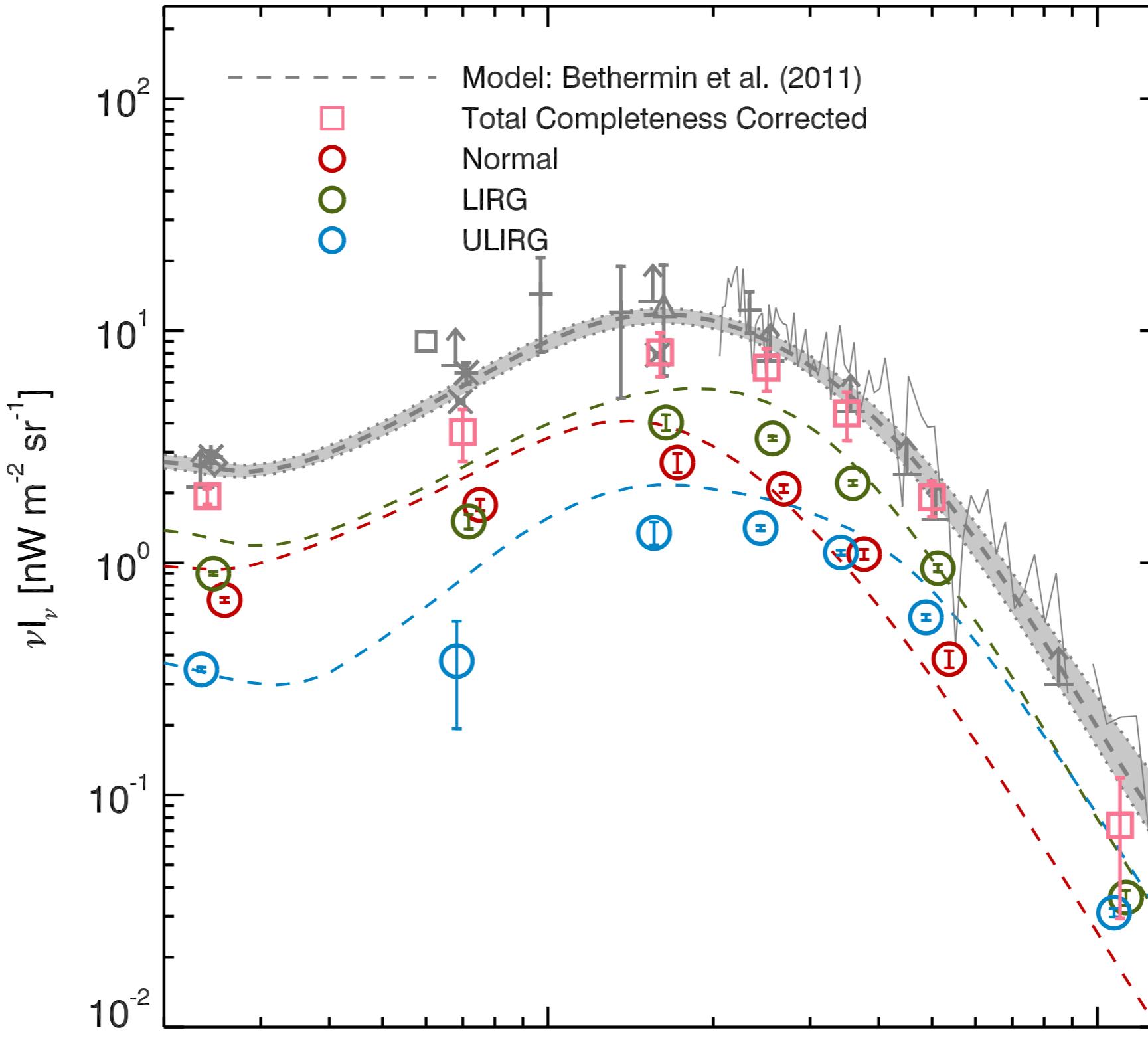
Viero, Moncelsi, Quadri et al. (2013)

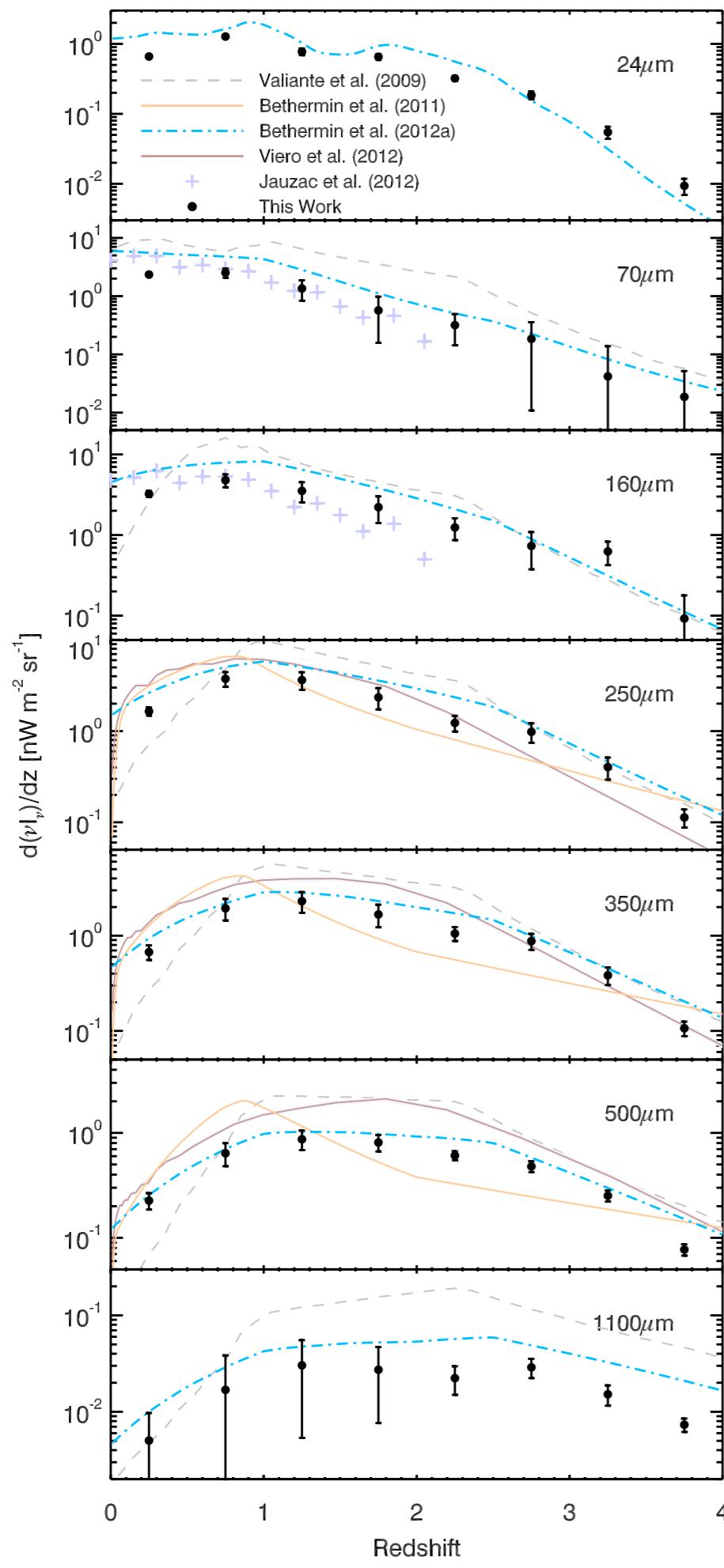
arXiv:1304.0446

# Infrared Luminosity

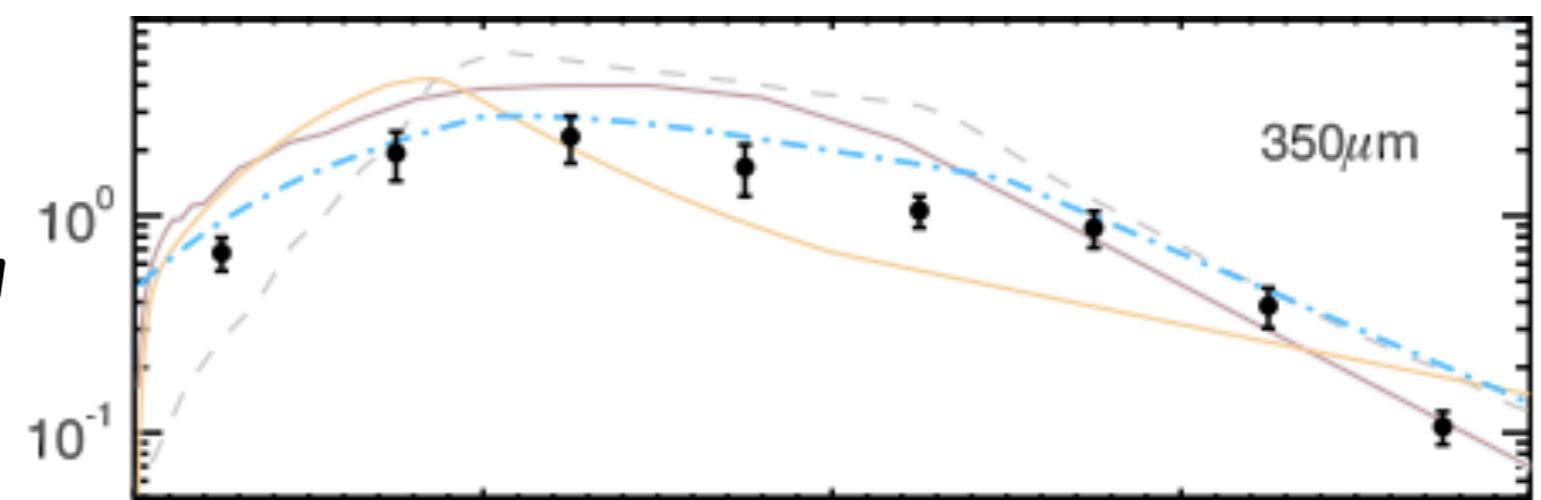


# CIB by Luminosity Class



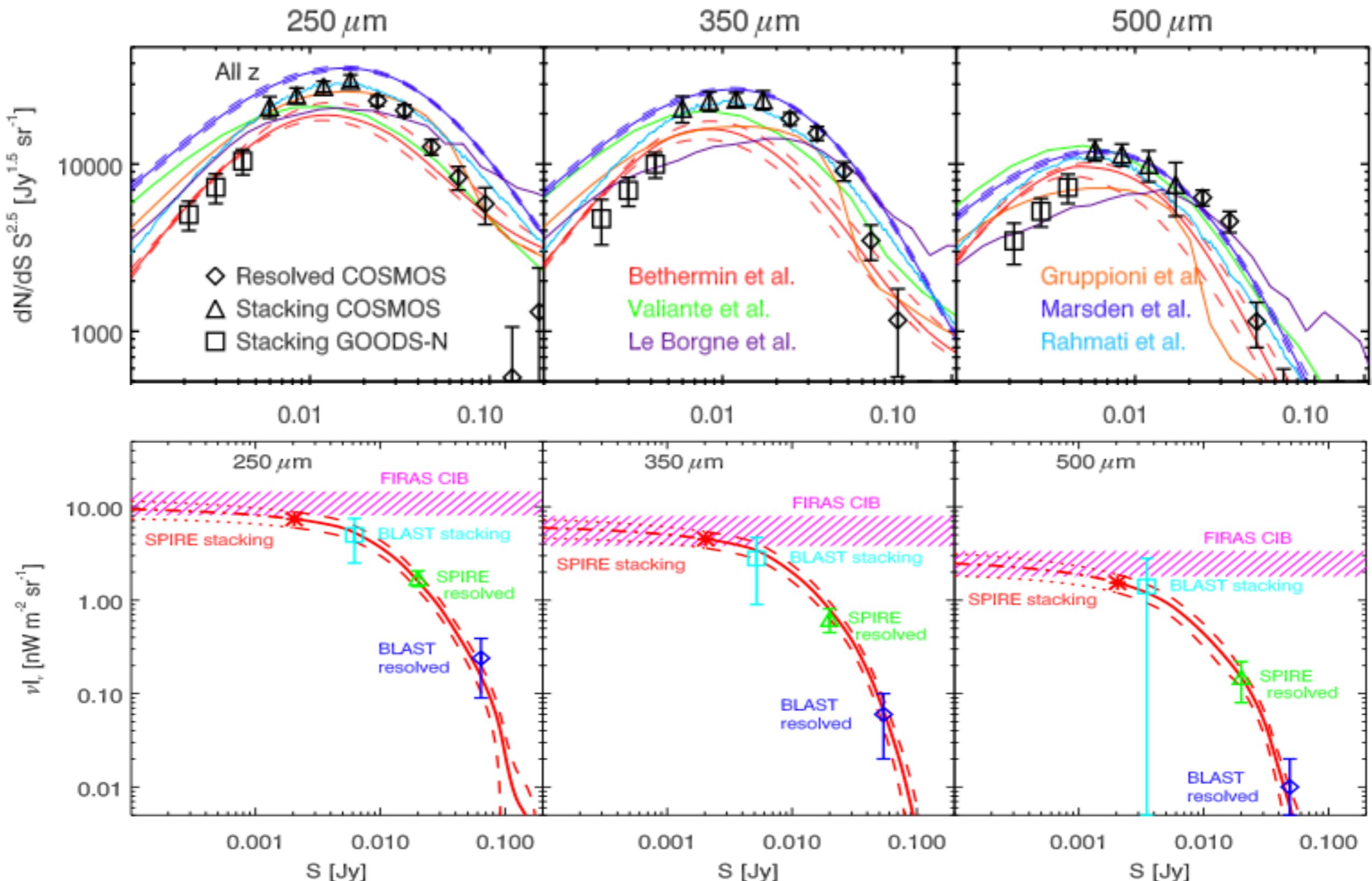


# Redshift Distribution of CIB



Viero, Moncelsi, Quadri et al. (2013)  
arXiv:1304.0446

# Deep Counts and the CIB



Bethermin et al. (2012) arXiv: 1203.1925

See also: Heinis++ 2013, Hilton++ 2012

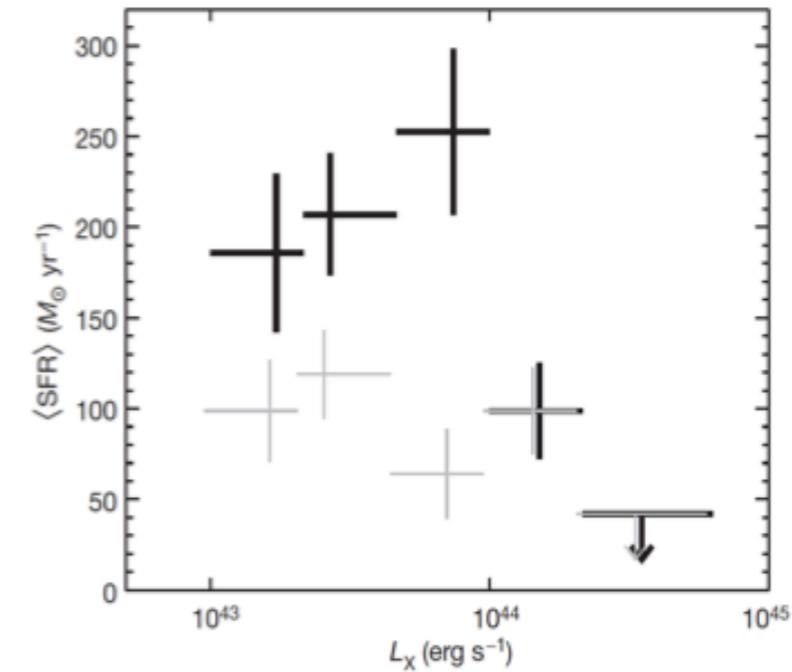
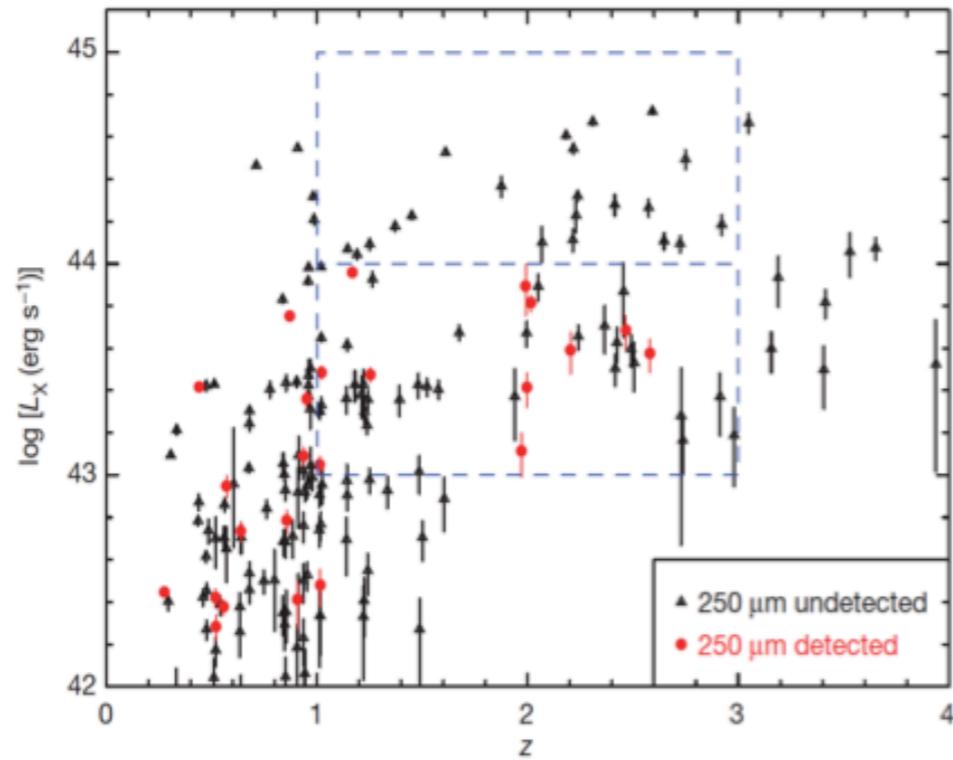
# AGN Stacking

## LETTER

doi:10.1038/nature11096

### The suppression of star formation by powerful active galactic nuclei

M. J. Page<sup>1</sup>, M. Symeonidis<sup>1</sup>, J. D. Vieira<sup>2</sup>, B. Altieri<sup>3</sup>, A. Amblard<sup>4</sup>, V. Arumugam<sup>5</sup>, H. Aussel<sup>6</sup>, T. Babbedge<sup>7</sup>, A. Blain<sup>8</sup>, J. Bock<sup>2,9</sup>, A. Boselli<sup>10</sup>, V. Buat<sup>10</sup>, N. Castro-Rodríguez<sup>11,12</sup>, A. Cava<sup>13</sup>, P. Chanial<sup>6</sup>, D. L. Clements<sup>7</sup>, A. Conley<sup>14</sup>, L. Conversi<sup>3</sup>, A. Cooray<sup>2,15</sup>, C. D. Dowell<sup>2,9</sup>, E. N. Dubois<sup>16</sup>, J. S. Dunlop<sup>5</sup>, E. Dwek<sup>17</sup>, S. Dye<sup>18</sup>, S. Eales<sup>19</sup>, D. Elbaz<sup>6</sup>, D. Farrah<sup>16</sup>, M. Fox<sup>7</sup>, A. Franceschini<sup>20</sup>, W. Gear<sup>19</sup>, J. Glenn<sup>14,21</sup>, M. Griffin<sup>19</sup>, M. Halpern<sup>22</sup>, E. Hatziminaoglou<sup>23</sup>, E. Ibar<sup>24</sup>, K. Isaak<sup>25</sup>, R. J. Ivison<sup>5,24</sup>, G. Lagache<sup>26</sup>, L. Levenson<sup>2,9</sup>, N. Lu<sup>2,27</sup>, S. Madden<sup>6</sup>, B. Maffei<sup>28</sup>, G. Mainetti<sup>20</sup>, L. Marchetti<sup>20</sup>, H. T. Nguyen<sup>2,9</sup>, B. O'Halloran<sup>7</sup>, S. J. Oliver<sup>16</sup>, A. Omont<sup>29</sup>, P. Panuzzo<sup>6</sup>, A. Papageorgiou<sup>19</sup>, C. P. Pearson<sup>30,31</sup>, I. Pérez-Fournon<sup>11,12</sup>, M. Pohlen<sup>19</sup>, J. I. Rawlings<sup>1</sup>, D. Rigopoulou<sup>30,32</sup>, L. Riguccini<sup>6</sup>, D. Rizzo<sup>7</sup>, G. Rodighiero<sup>20</sup>, I. G. Roseboom<sup>5,16</sup>, M. Rowan-Robinson<sup>7</sup>, M. Sánchez Portal<sup>3</sup>, B. Schulz<sup>2,27</sup>, D. Scott<sup>22</sup>, N. Seymour<sup>1,33</sup>, D. L. Shupe<sup>2,27</sup>, A. J. Smith<sup>16</sup>, J. A. Stevens<sup>34</sup>, M. Trichas<sup>35</sup>, K. E. Tugwell<sup>1</sup>, M. Vaccari<sup>20</sup>, I. Valtchanov<sup>3</sup>, M. Viero<sup>2</sup>, L. Vigroux<sup>29</sup>, L. Wang<sup>16</sup>, R. Ward<sup>16</sup>, G. Wright<sup>24</sup>, C. K. Xu<sup>2,27</sup> & M. Zemcov<sup>2,9</sup>



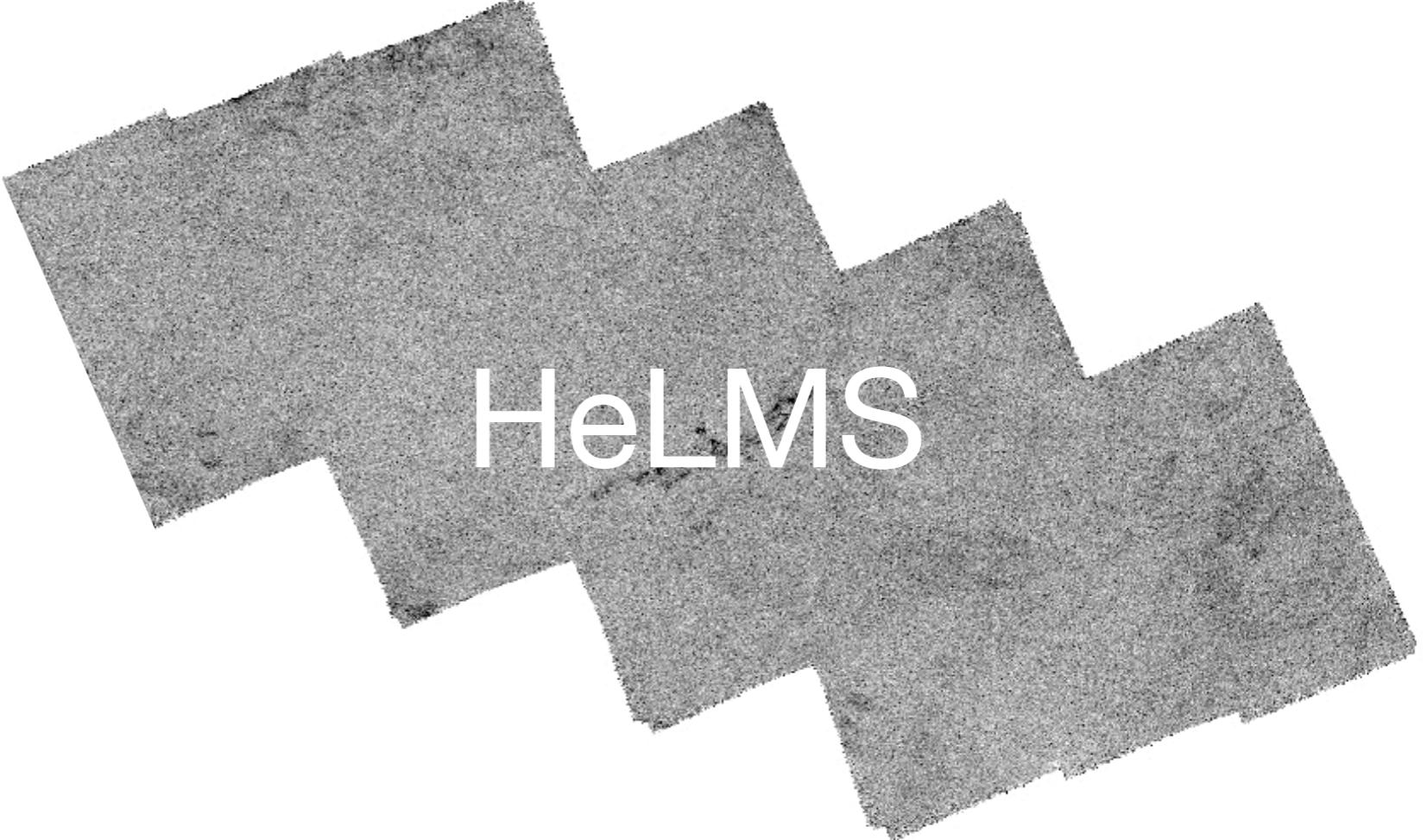
Page++ 2012, *Nature*, 485(7), pp.213–216

See also: Seymour++ 2011, Hatziminaoglou++ 2010, Dai++ 2012

SDSS Stripe 82

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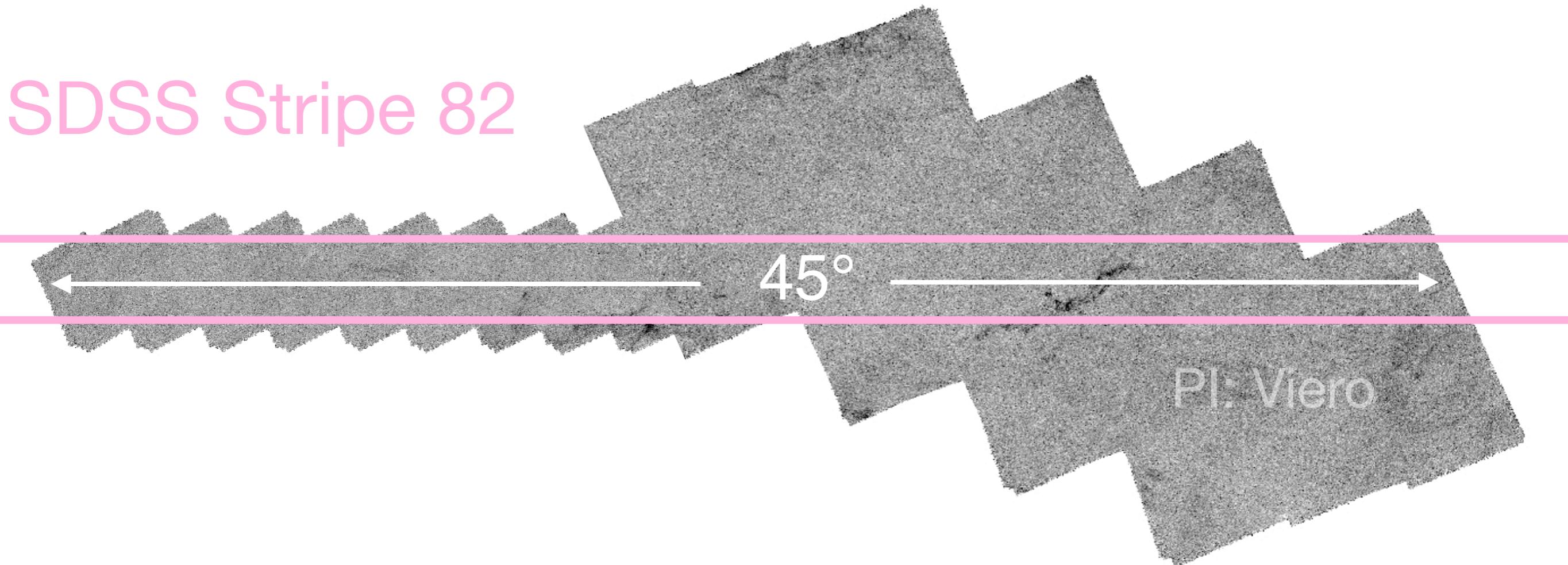


HeLMS

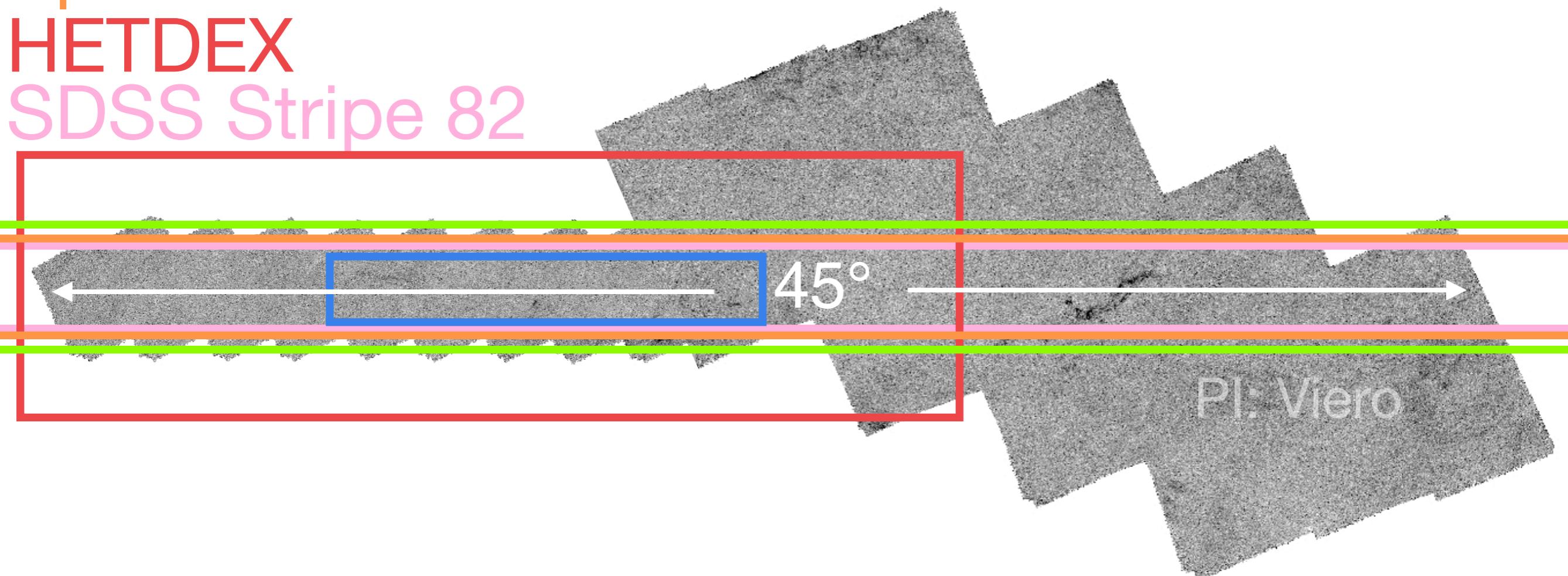


HeRS

SDSS Stripe 82



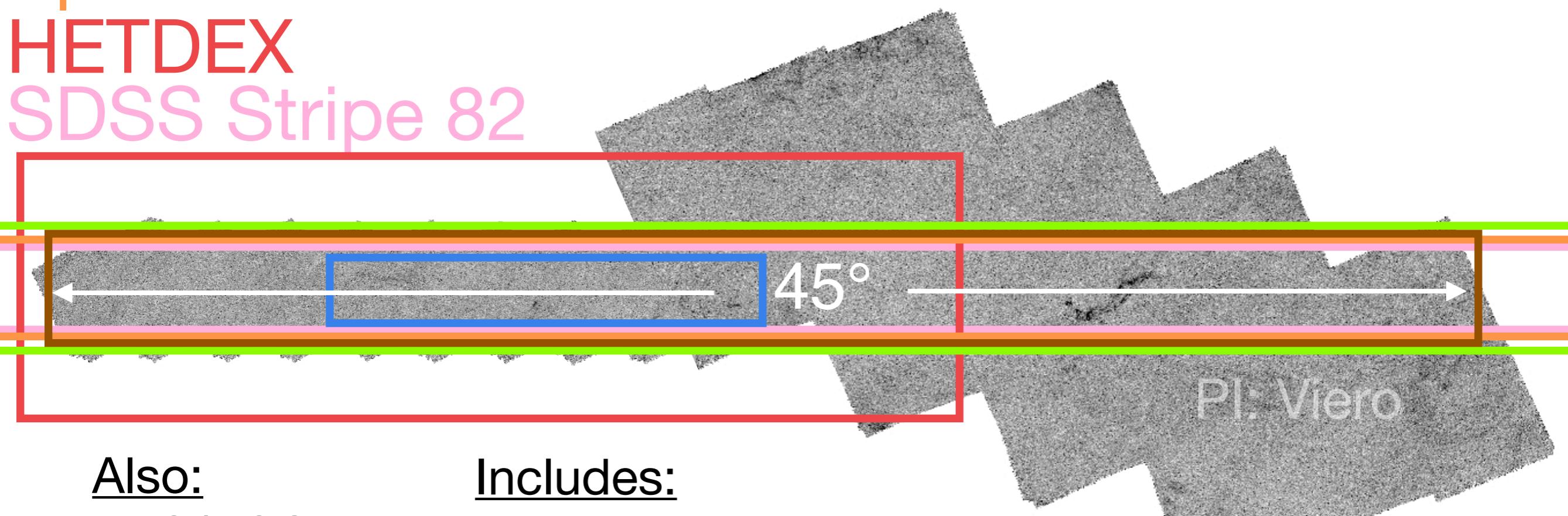
ACT  
SHELA  
SpIES  
HETDEX  
SDSS Stripe 82



Viero++2013, Herschel Stripe 82 Survey; arXiv:1308.4399

Find Maps/Catalogs at: <http://www.astro.caltech.edu/hers>

ACT  
SHELA  
SpIES  
HETDEX  
SDSS Stripe 82



Also:

- DES/HSC
- VHS/MICS82
- VLA
- Wiggle-z
- LSST

Includes:

- Clusters
- QSOs
- LRGs
- maxBCGs
- HI

• Optical Spectra:

- Lyman Alpha Forest
- DLAs/Mg2/CIV

**END!**