

# Fine Structure Lines toward NGC 2024

U.U. Graf



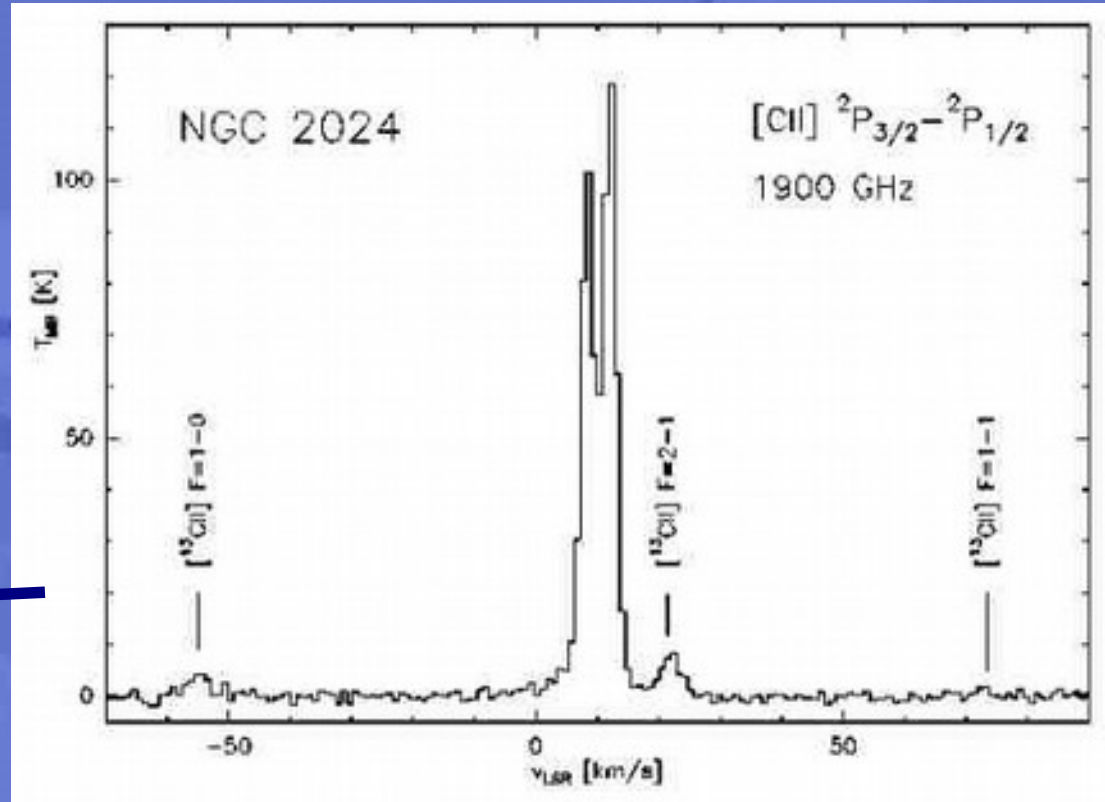
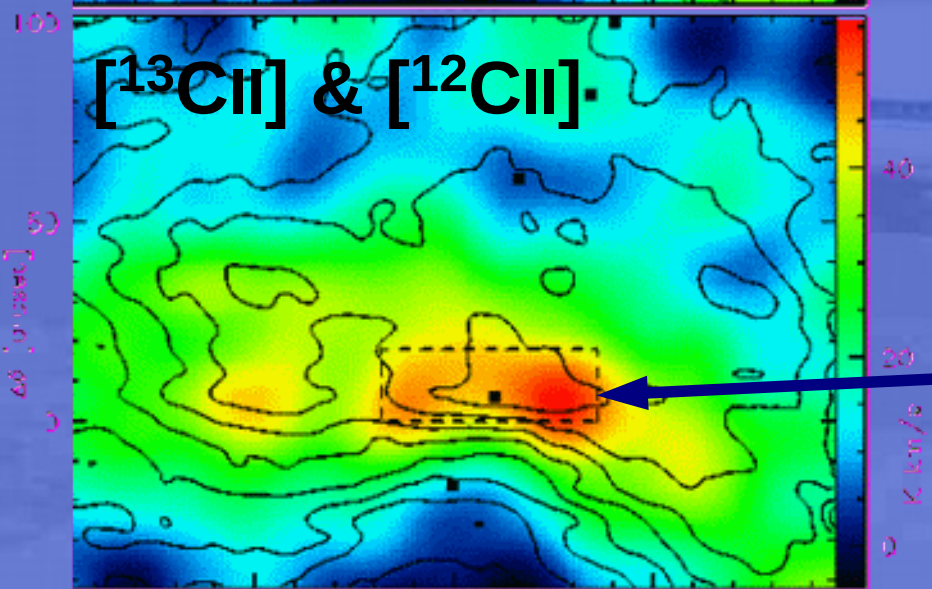
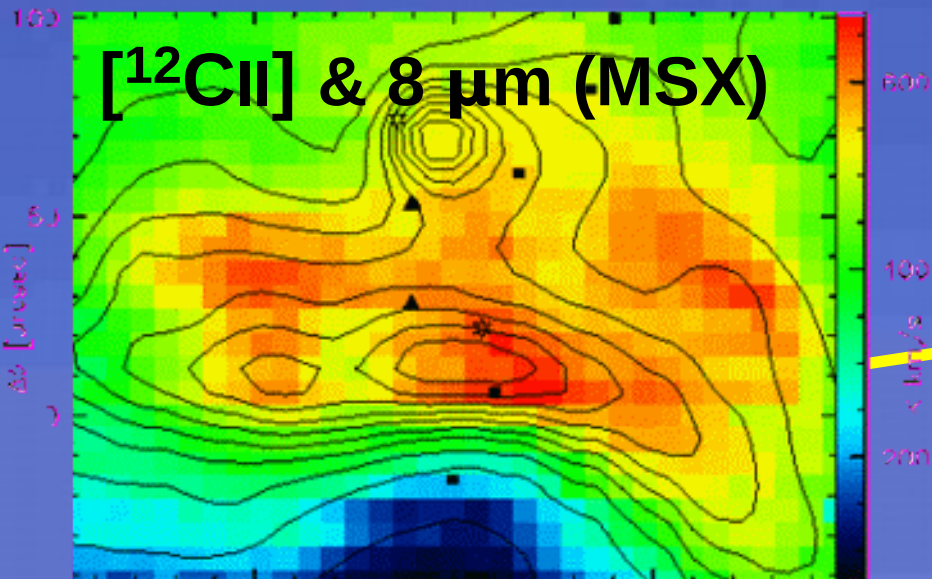
R. Simon, J. Stutzki, R. Güsten et al.



# Strong $[^{12}\text{CII}]$ & $[^{13}\text{CII}]$ emission



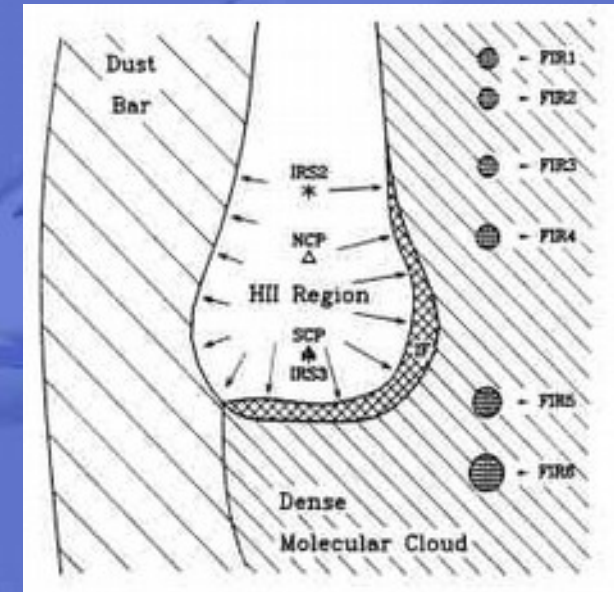
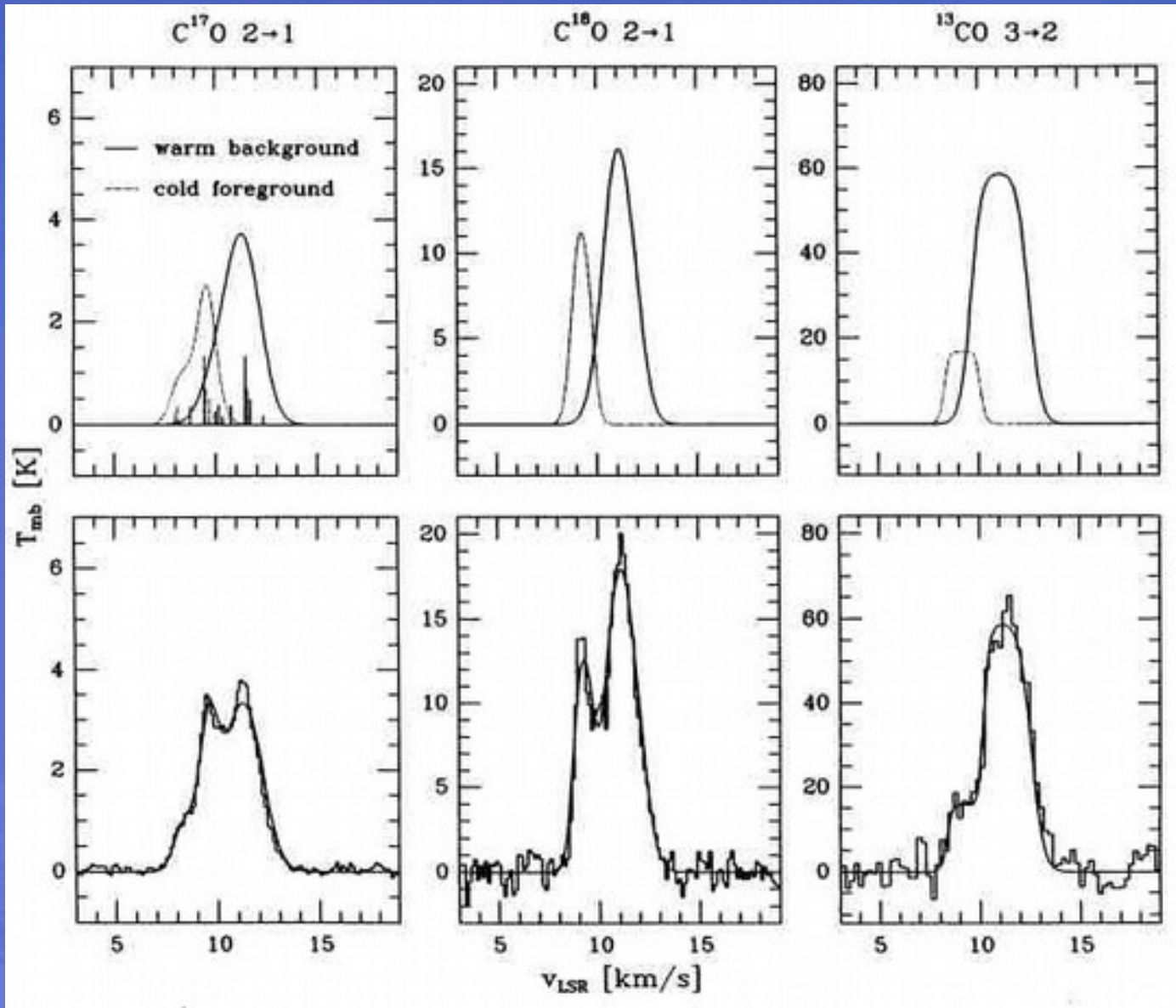
$[^{12}\text{CII}]$  and  $[^{13}\text{CII}]$  mapped in Early Science







# Standard source model: 2 emission components

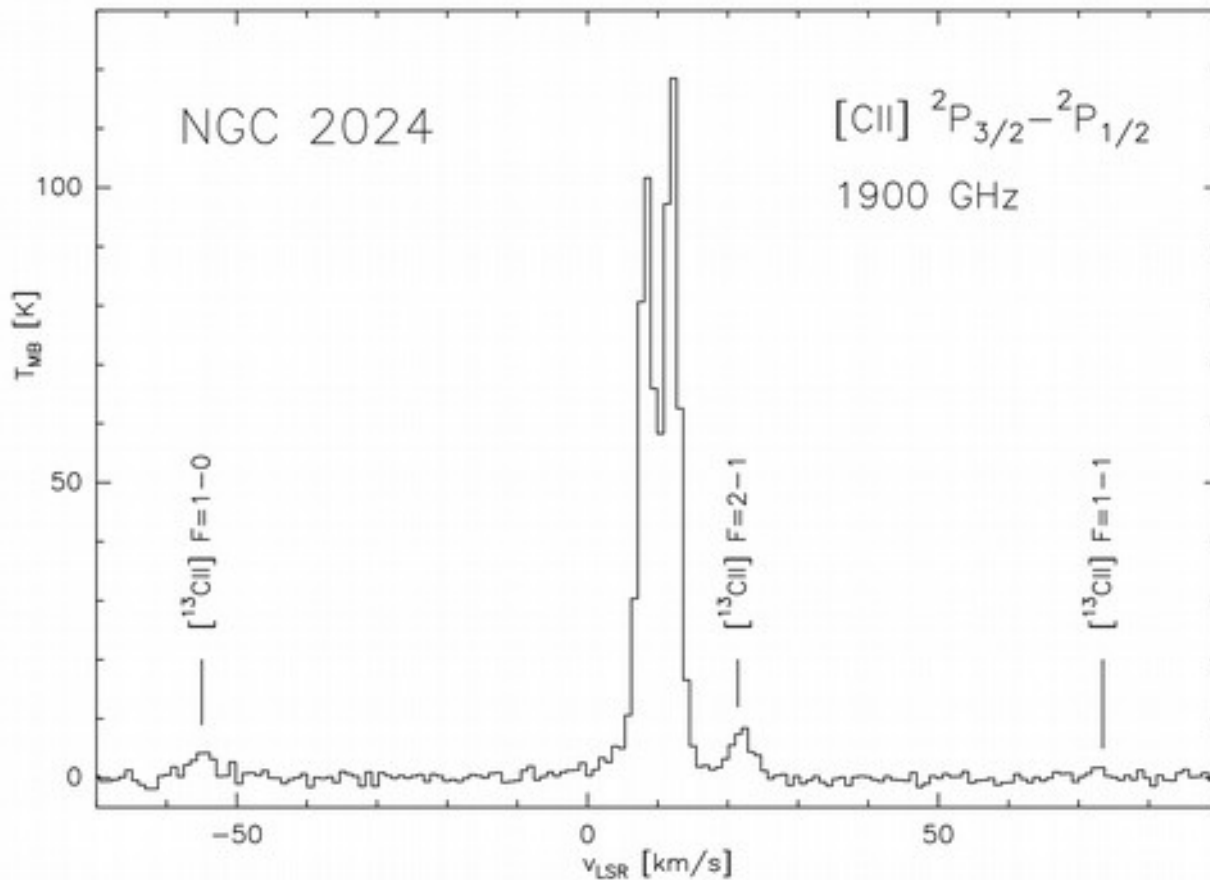


**cold  
fore-  
ground  
@9 km/s**

**warm  
back-  
ground  
@11 km/s**



# [CII]: self-absorbed emission



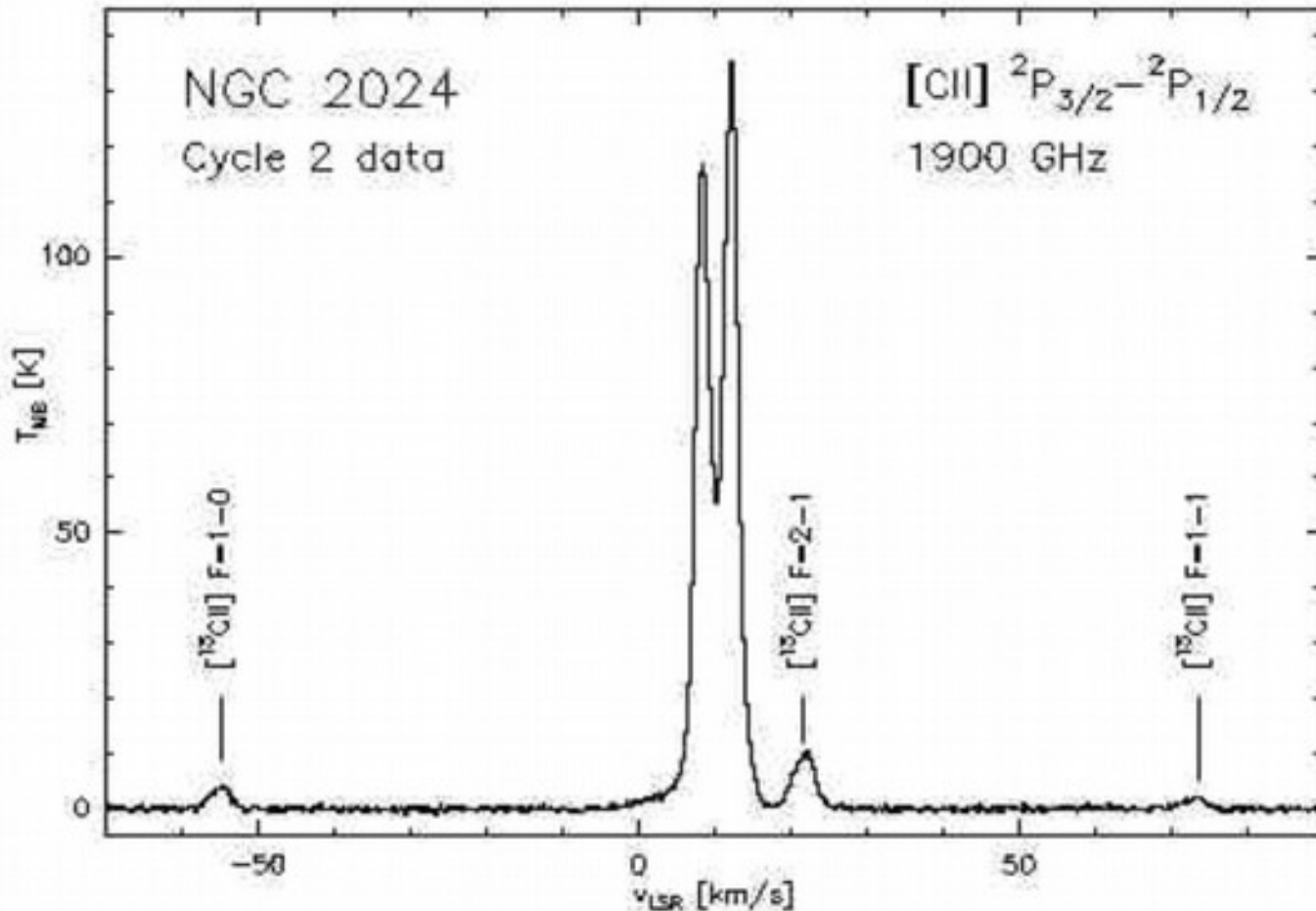
Revisited in

January 2015:

- 1.5' x 1' maps in
  - [OI]
  - [ $^{12}\text{CII}$ ], [ $^{13}\text{CII}$ ]
- 3' x 2' maps in
  - [ $^{12}\text{CII}$ ], [ $^{13}\text{CII}$ ]
  - $^{12}\text{CO}$  13→12
  - $^{13}\text{CO}$  13→12



# [CII]: self-absorbed emission



Revisited in  
January 2015:

- 1.5' x 1' maps in
  - [OI]
  - [12CII], [13CII]
- 3' x 2' maps in
  - [12CII], [13CII]
  - 12CO 13→12
  - 13CO 13→12



# Two step line modelling: 1: [ $^{13}\text{CII}$ ] - HFS - fit

Fit optically thin [ $^{13}\text{CII}$ ] HFS components with [ $^{12}\text{CII}$ ] blanked.

Yields

- $T_{\text{ex}} \times \tau([\text{CII}]) = 962 \text{ K}$





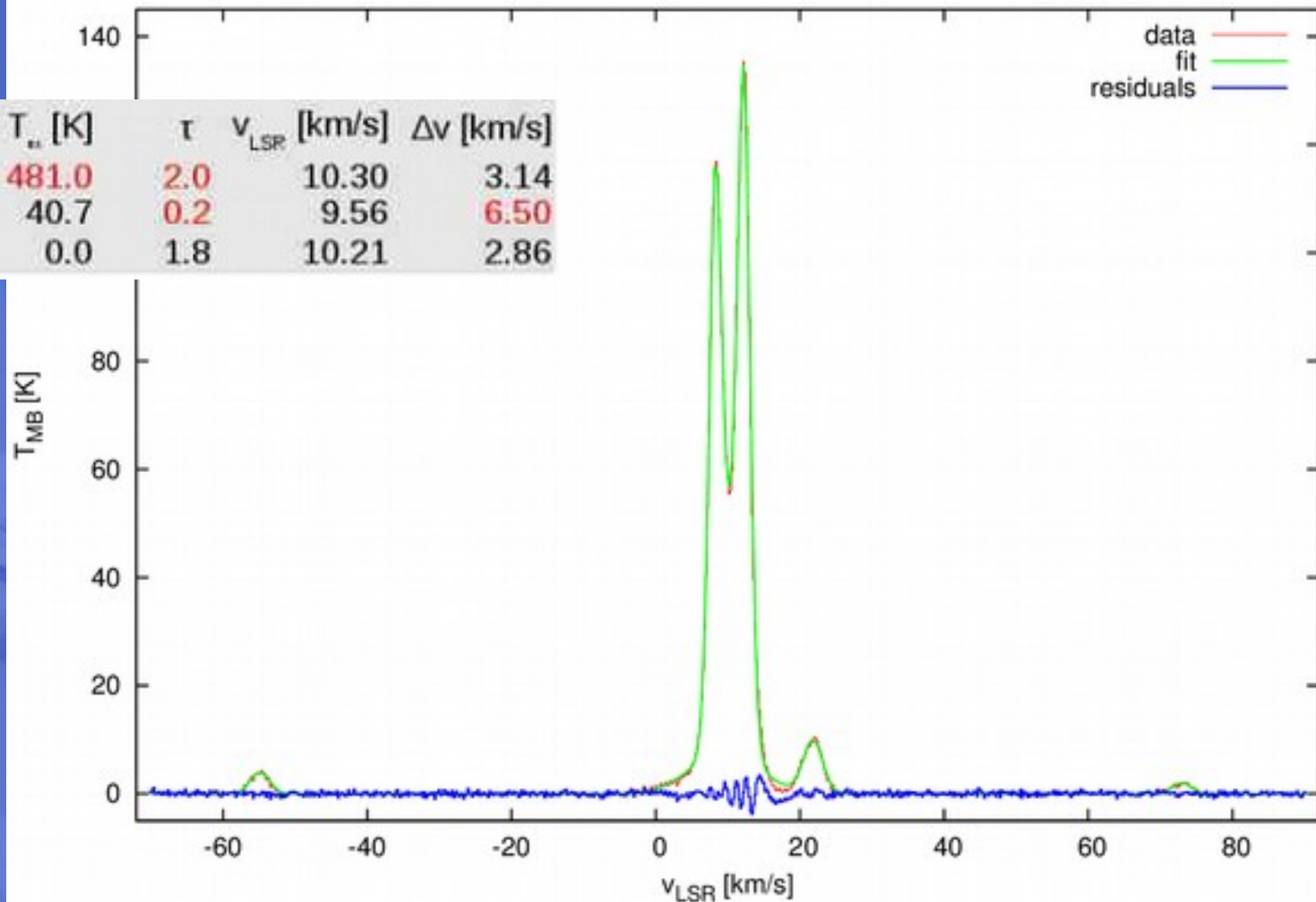
# Two step line modelling: 2: 3-component full line fit

Component	$T_{\text{a}}$ [K]	$\tau$	$v_{\text{LSR}}$ [km/s]	$\Delta v$ [km/s]
Background	<b>481.0</b>	<b>2.0</b>	10.30	3.14
broad emission	40.7	<b>0.2</b>	9.56	<b>6.50</b>
Foreground	0.0	1.8	10.21	2.86

Fixed values  
in **RED**

$N_{\text{BG}}(^{12}\text{C}^+) >$   
 $10^{19} \text{ cm}^{-2}$

$N_{\text{FG}}(^{12}\text{C}^+) >$   
 $10^{18} \text{ cm}^{-2}$



**Note: Rest frequencies of HFS-components revised**



# [OI] 63 $\mu\text{m}$ map

[OI] integrated intensity

overlayed on [ $^{13}\text{CII}$ ] map



Peaks just north of ionization front - as expected, but...





# [OI] channel maps



**essentially NO emission between 9 km/s and 11 km/s!**



# [OI] sample spectrum

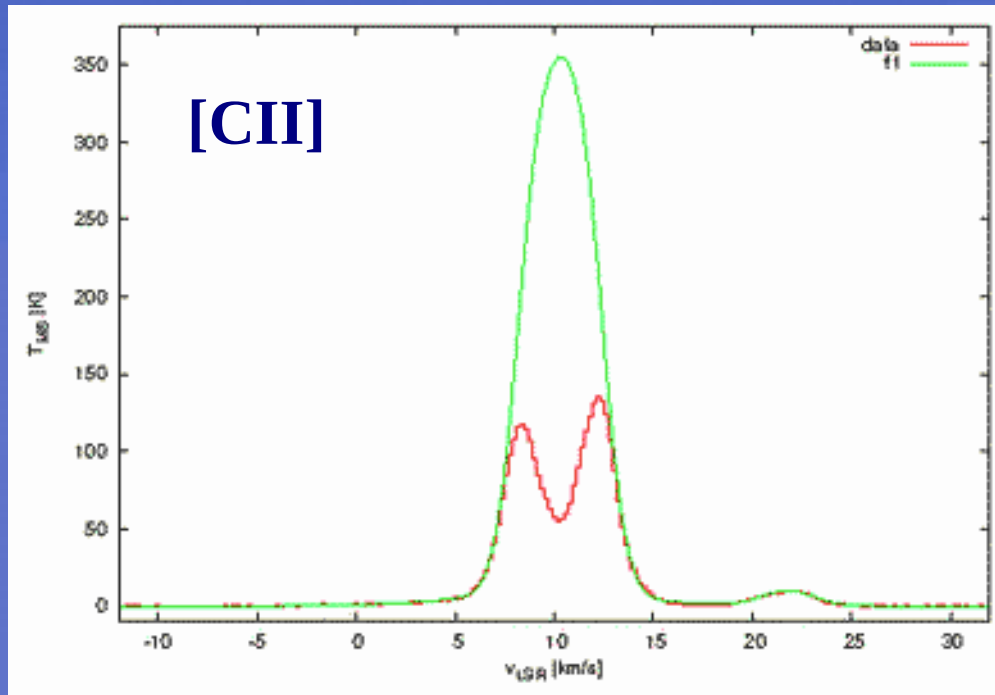
Fixed values  
in RED

Massive  
foreground  
absorption





# Foreground absorption artificially removed



**[OI]**

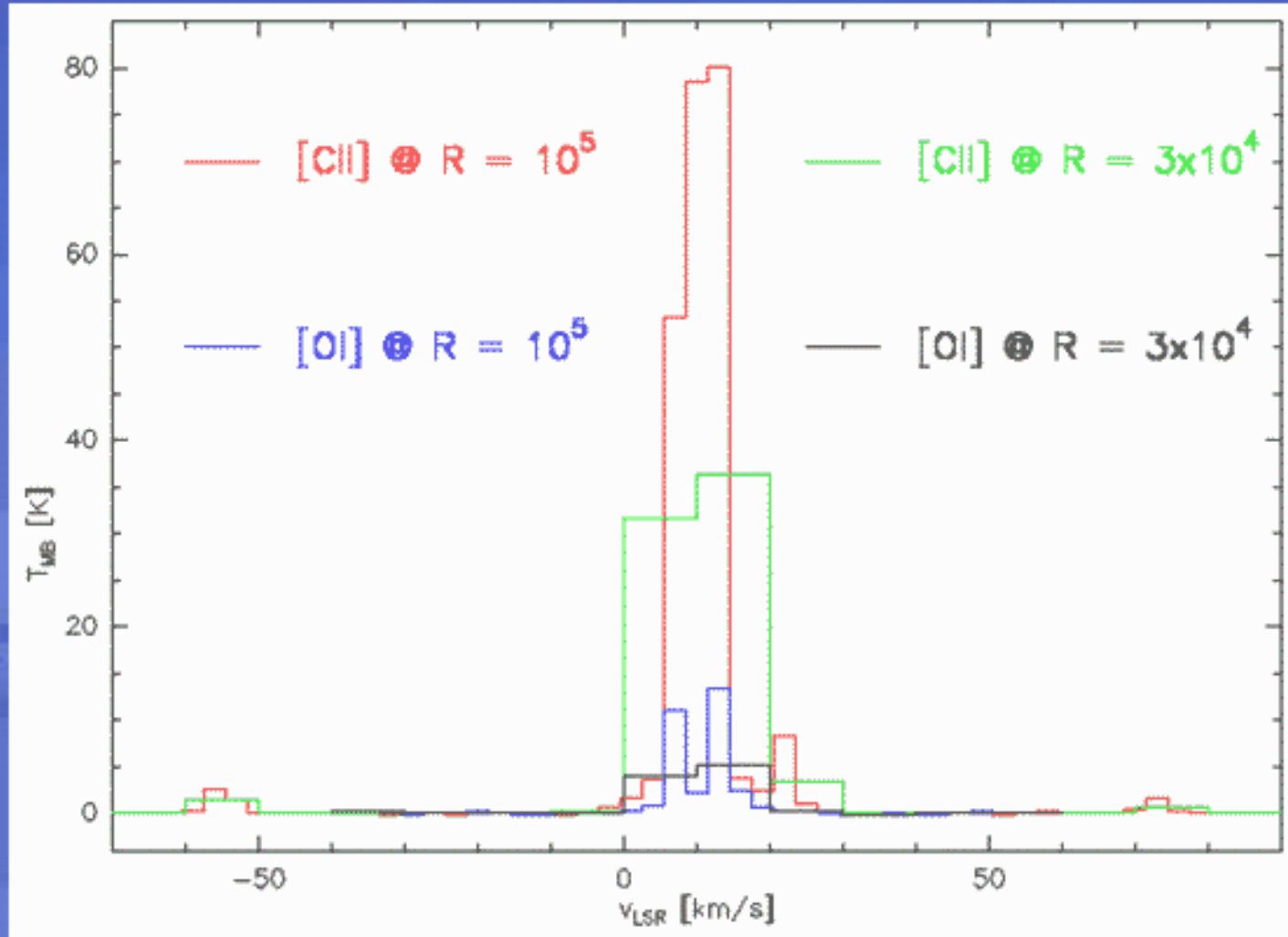


>60% of [CII] emission is  
obscured by foreground

~85% of [OI] emission is  
obscured by foreground



# Resolving power matters!



[CII] self-absorption disappears at  $R < 100000$   
[OI] self-absorption disappears at  $R < 30000$





$^{13}\text{CO } 13 \rightarrow 12$  on  $[^{13}\text{CII}]$





# Summary

- High resolution spectroscopy reveals strong foreground absorption in [OI] and [CII]:
  - ~85% of background [OI] emission obscured
  - >60% of background [CII] emission obscured
- Consistent with FG hydrogen column density of  $10^{22} \text{ cm}^{-2}$
- Background [OI] / [CII] intensity ratio ~ 5 indicates substantial density and radiation field
- Fine structure emission peaks just north (or wraps around) dense condensation (seen in  $^{13}\text{CO } 13 \rightarrow 12$ )

**Finally: High spectral resolution is crucial for accurate line intensities and to disentangle line of sight source structure**