



# Stratospheric Observing Primer

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SOFIA/USRA





# Space vs. Ground



Image credit: NASA astronaut Gregory Reid Wiseman



SOFIA Observers Workshop 20+21 May 2015





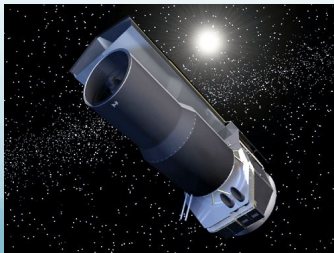


# Space vs. Ground



## Space

- **PRO**
  - Fantastic sensitivity/transmission
  - No atmospheric distortion (seeing)
- **CON**
  - Expensive (smaller telescopes)
  - Limited lifetime (cryo)\* Small Time Domain
  - No easy repairs



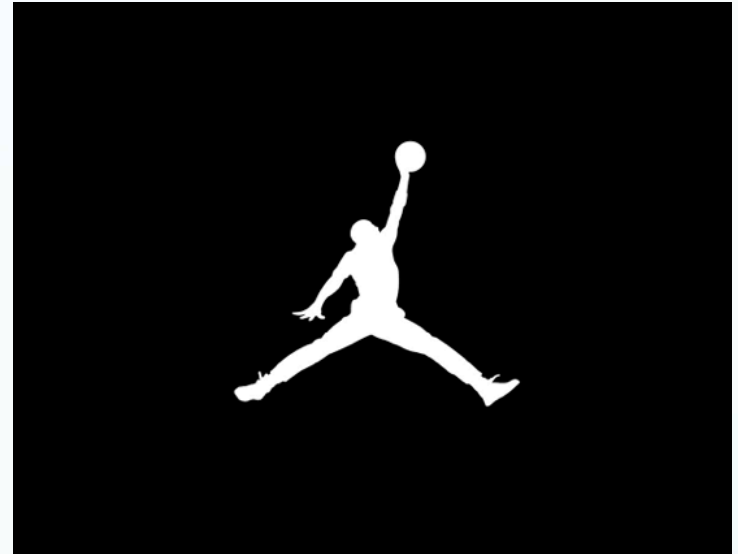
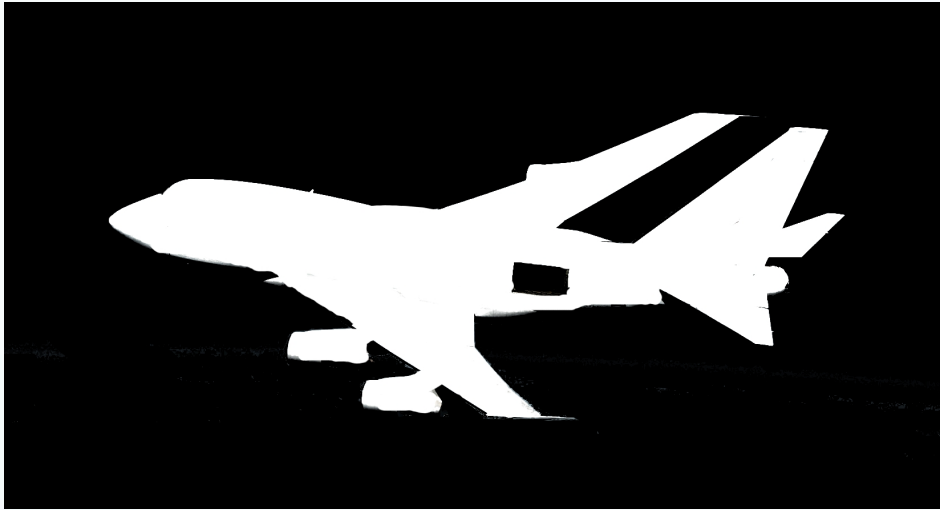
## Ground

- **PRO**
  - Cheaper (larger telescopes)
  - No cryo issue\* Large Time Domain
  - Repairable (duct tape?)
- **CON**
  - Worse transmission/sensitivity
  - Atmospheric distortion (seeing)





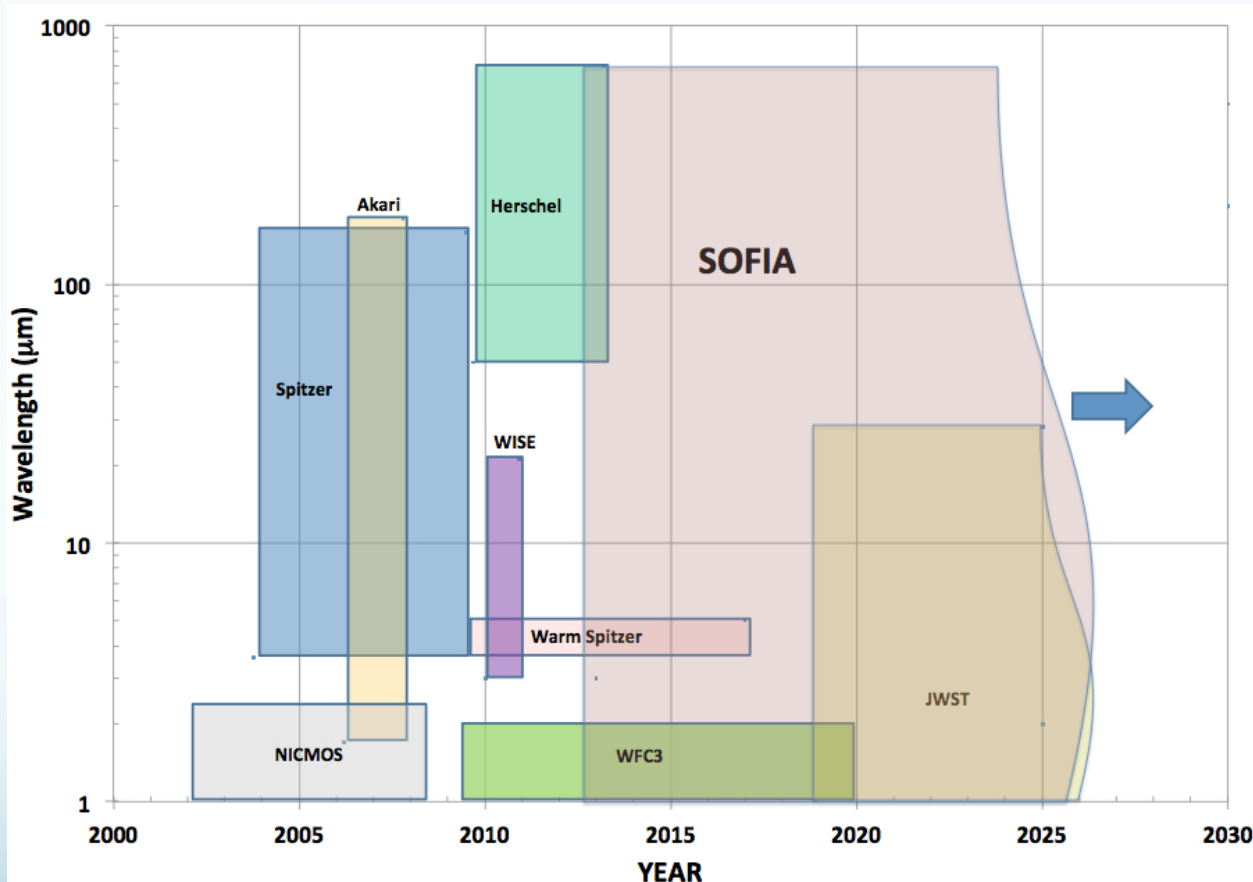
# AIR



- Advantage of both space and ground
- Unique issues with observing in plane
- Overview, highlighting a few instruments and modes



# Air



- Large Time Domain\*
- Repairable
- Cost savings over comparable space mission
- Instrument technology upgrades (upGreat, HAWC +)
- Better transmission than ground

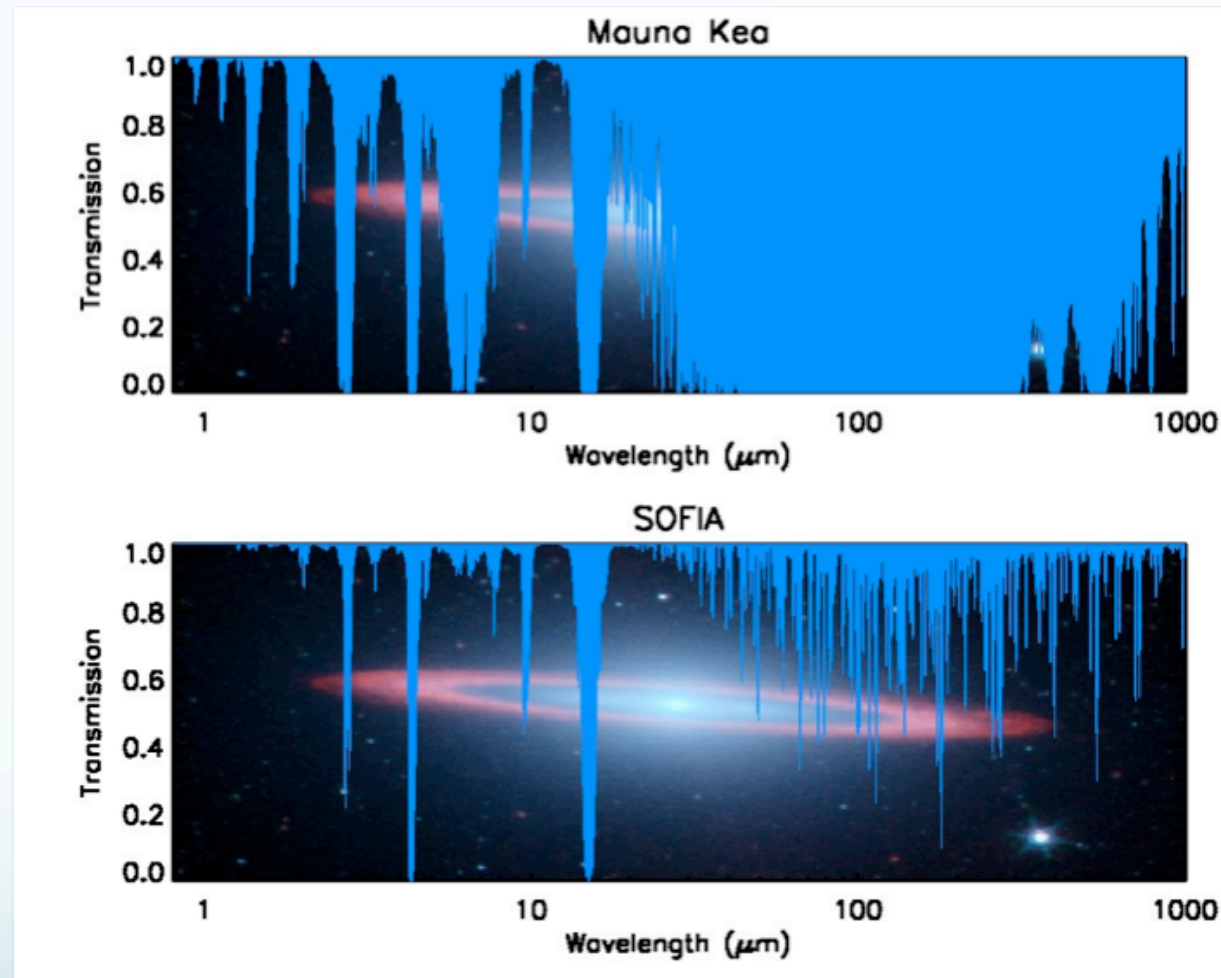




# Air

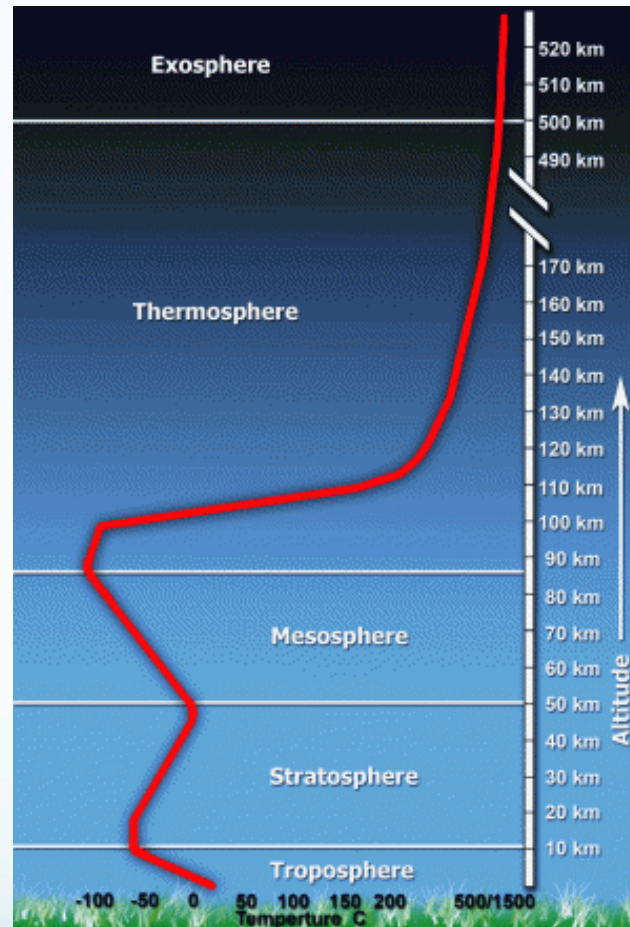


- Even at the very best ground-based sites, the transmission in the mid to far-infrared is poor or nonexistent
- Cool dust, light molecule rotation lines, atomic fine-structure lines etc., in this range provide unique tracers
- Operational elevation: 38,000 – 45,000 ft. (12 – 14 km) Above > 99.8% of atmospheric water vapor





# Earth's Atmosphere



<http://www.ces.fau.edu/nasa/module-2/atmosphere/earth.php>

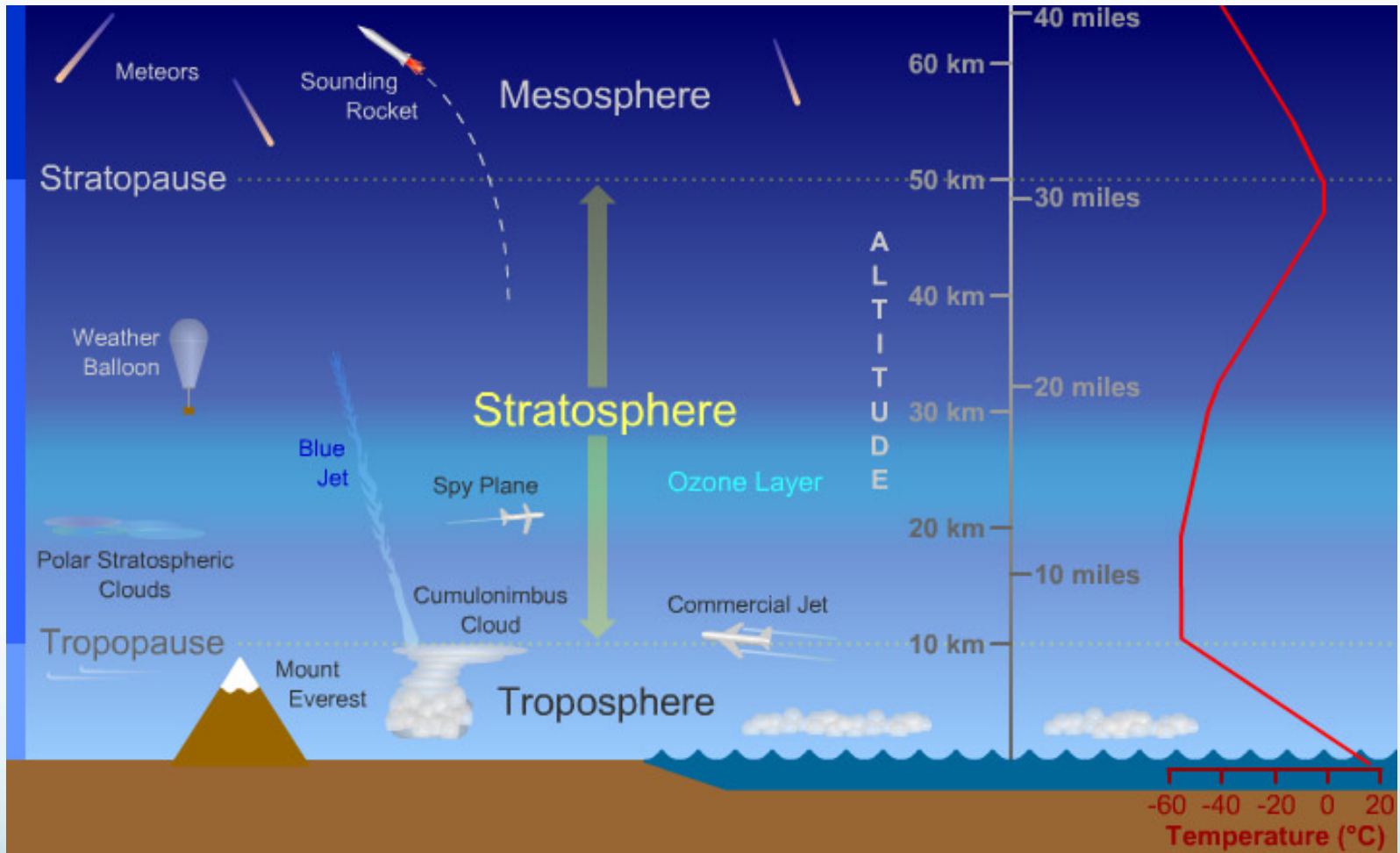


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# Earth's Atmosphere



<http://scied.ucar.edu/shortcontent/stratosphere-overview>



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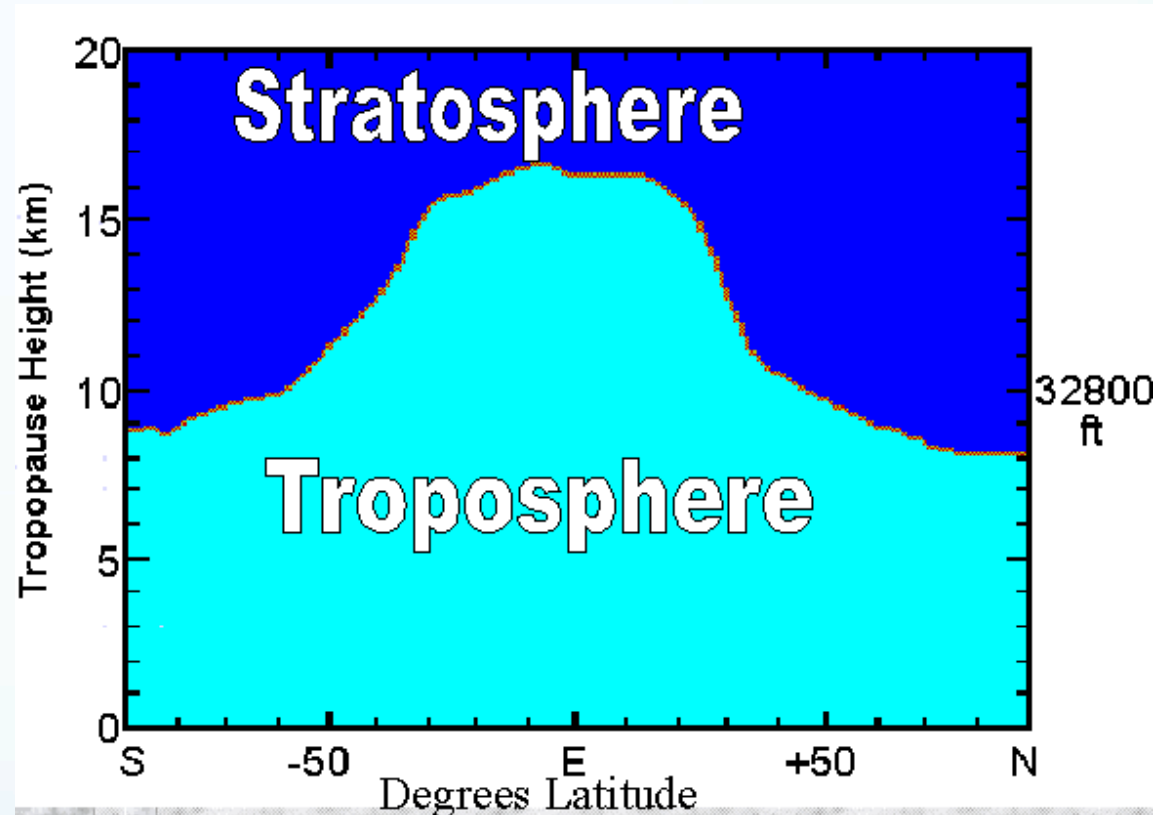




# Stratosphere



- Height of the Tropopause depends on the location
- Seasonal
- Colder regions have a lower Tropopause



<http://www.das.uwo.edu/~geerts/cwx/notes/chap01/tropo.html>

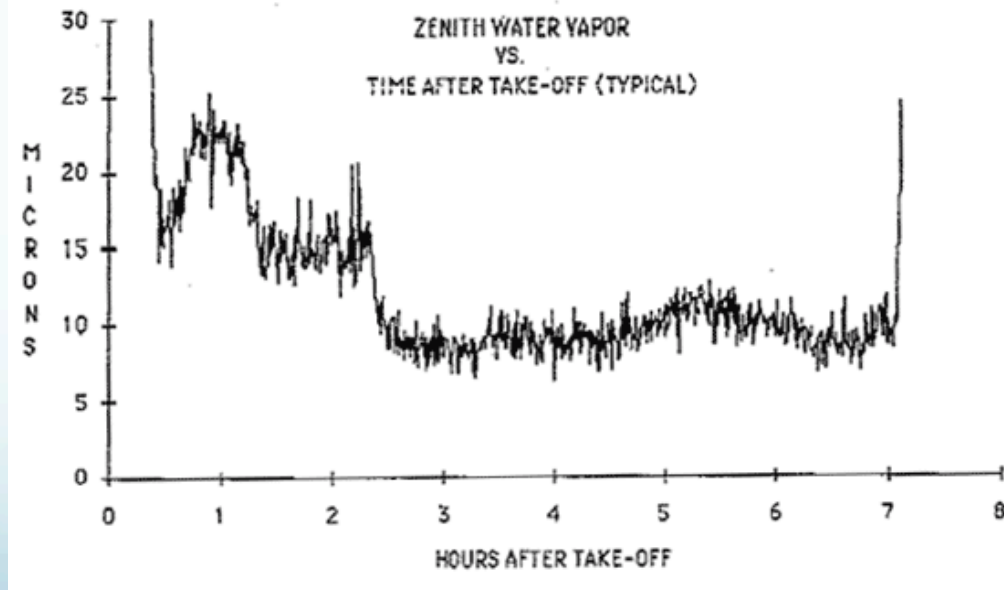




# Water Vapor Monitor



- Heterodyne mixer configured for the measurement of the 183 GHz rotational line of water.
- Mounted at a fixed elevation of  $40^\circ$  in the upper deck of the aircraft
- Stratospheric water vapor can be quite variable: as much as a factor of three on time scales as short as 15 minutes



KAO Example

<https://www.sofia.usra.edu/Science/instruments/waterVaporMonitor.html>

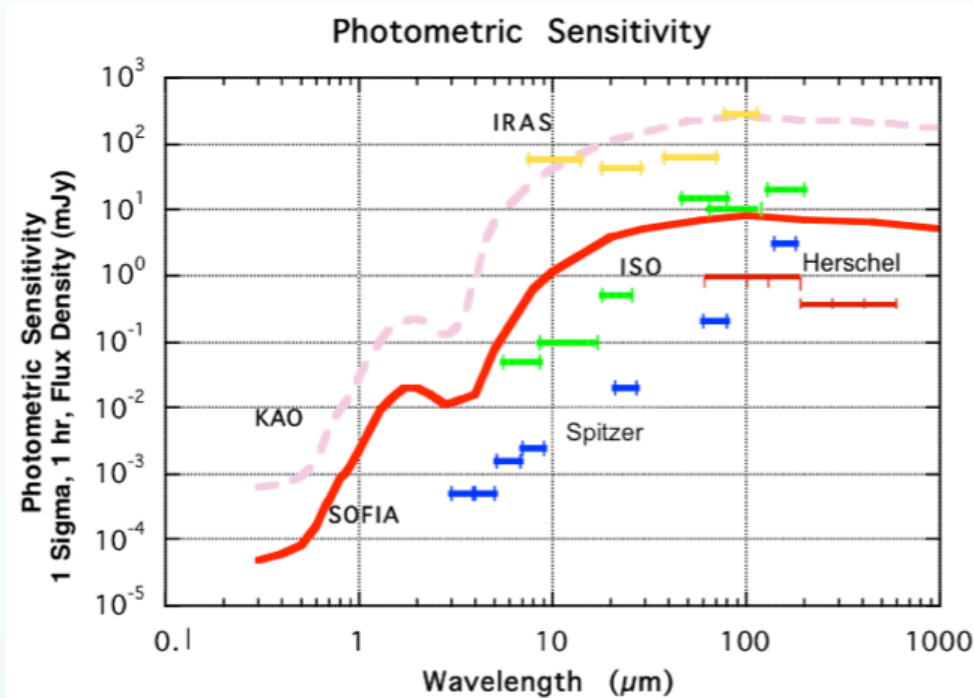


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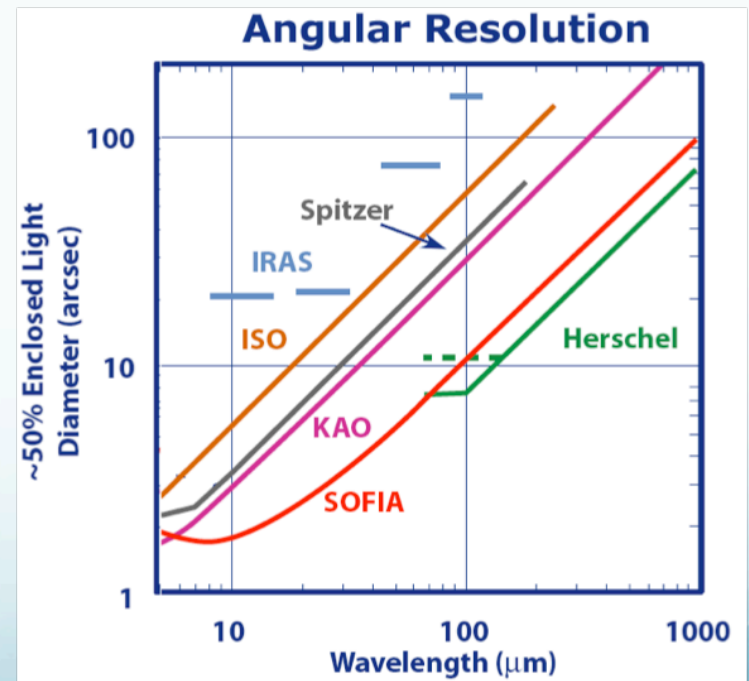




# Sensitivity/Resolution



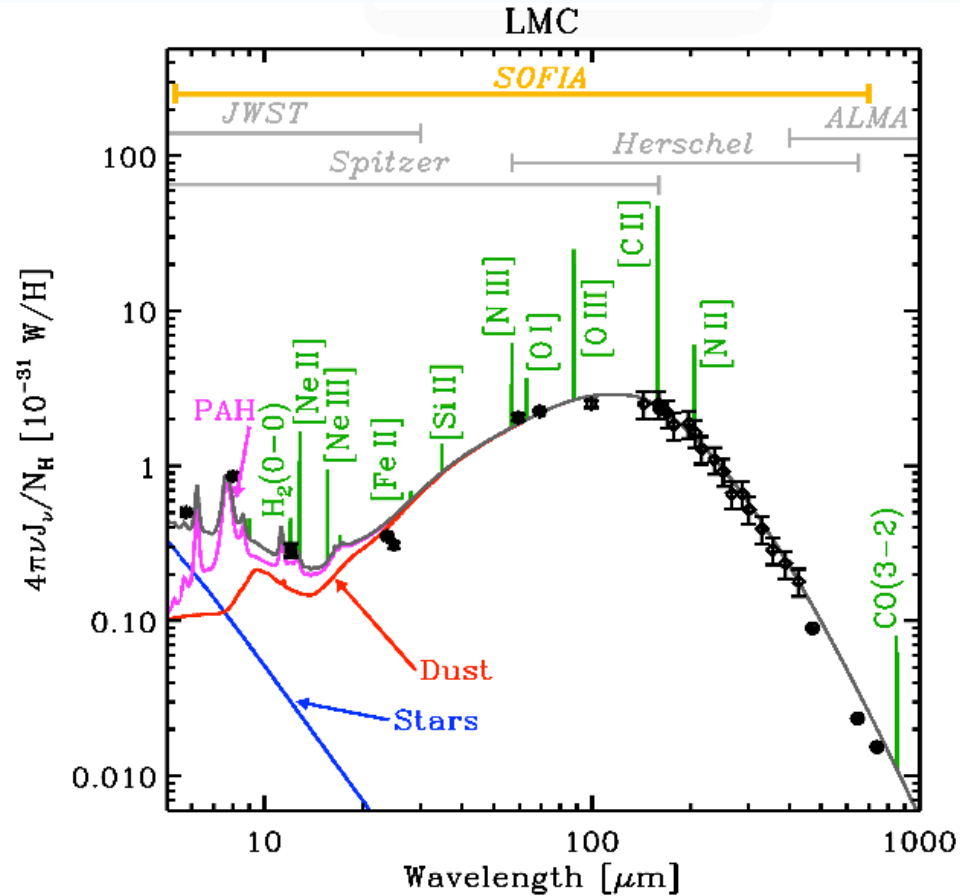
SOFIA is diffraction limited beyond  $\sim 25 \mu\text{m}$  and can produce sharper images than Spitzer



**Updated Sensitivities (SITE):**  
<https://dcs.sofia.usra.edu/proposalDevelopment/SITE/>







**Figure 3-2.** The spectral energy distribution of the entire LMC, based on data from Spitzer, IRAS and FIRAS (Bernard et al. 2008). SEDs are fitted with the dusty PDR model of Galliano et al. (2008). Spitzer has and Herschel will provide good photometric coverage of a galaxy's spectral energy distribution (SED) over a portion of the wavelengths. SOFIA will provide excellent wavelength coverage and spectroscopic capability across the entire SED. In the future, JWST and ALMA will provide complementary wavelength coverage and work on nearby galaxies and the most distant Universe. Figure courtesy of Galliano.



# Ground Similarities



- Observing techniques (still atmosphere)
  - Chop Nod
  - Scan Mapping
- Variability (weather)
- Calibration

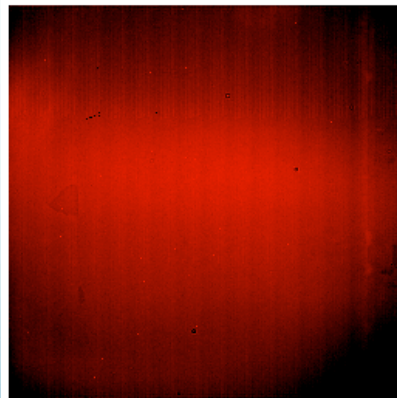




# Chopping and Nodding



- Done due to overwhelming photons from sky and telescope ( $\sim 100,000$  to 1)
- Chopping  $\sim$  a few HZ
- Nodding  $\sim$  10's of seconds
- Affects overhead to exposure time
- Distorts optical path (so limited chop amplitude)



FORCAST image (guess?)

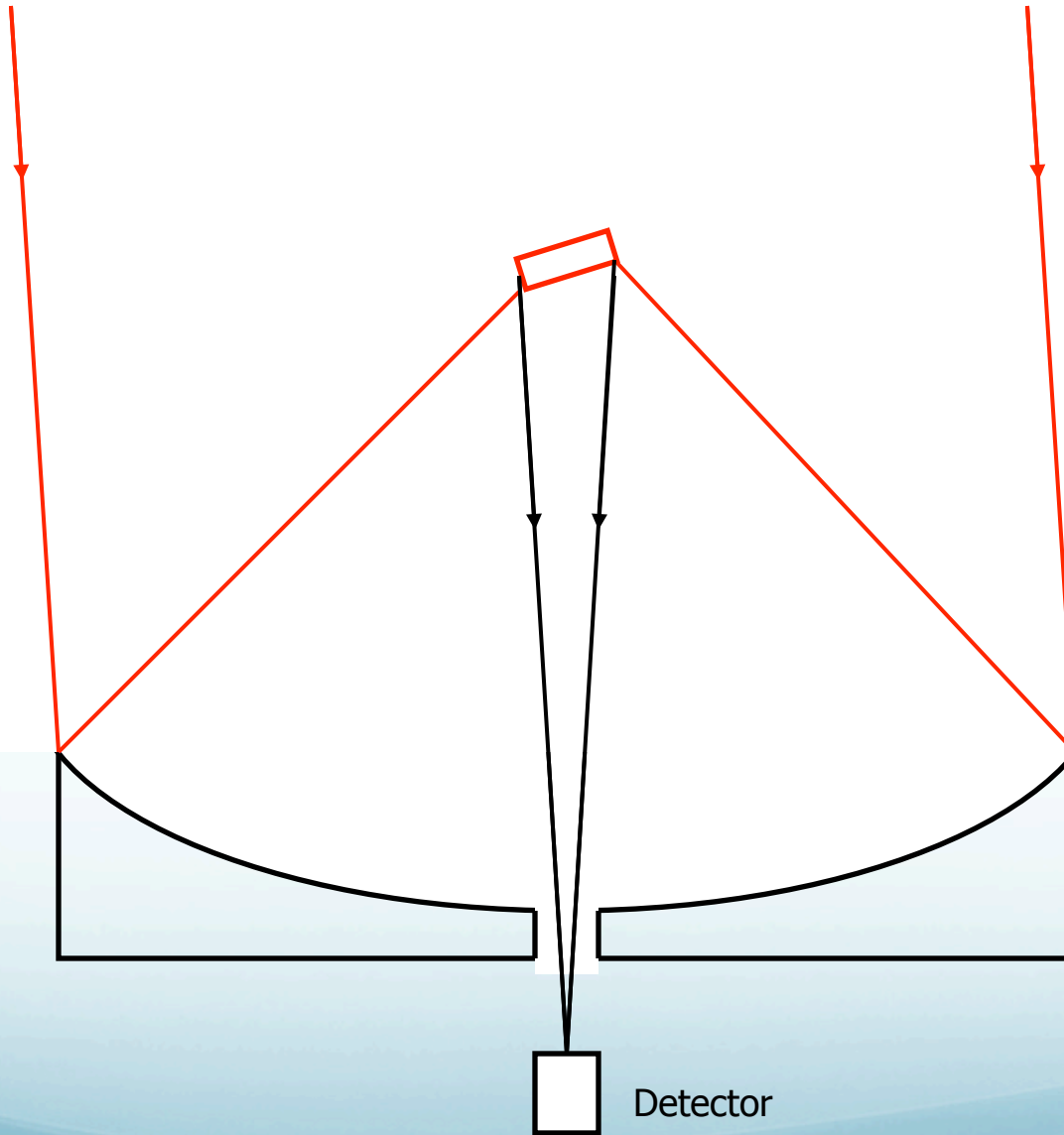




# Chopping Secondary

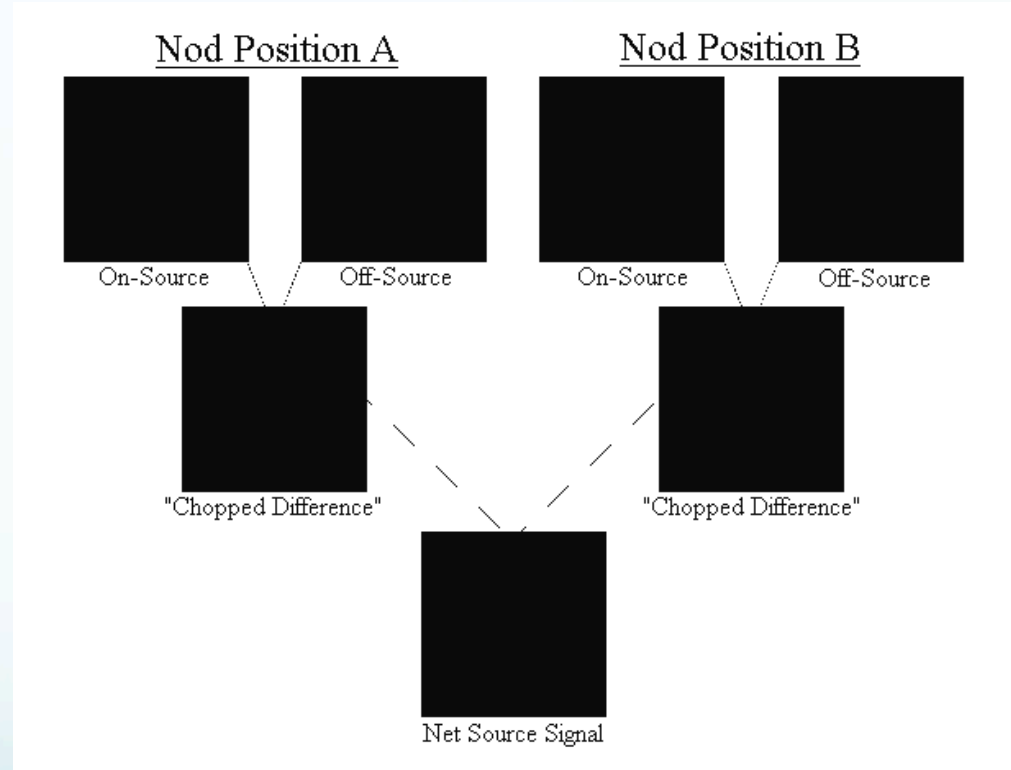
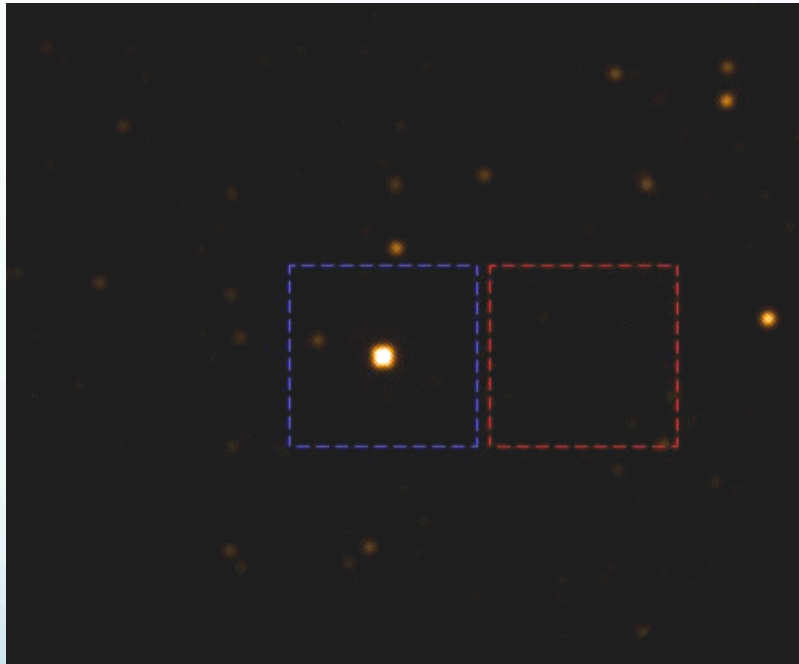


- Chop 1
- Chop 2



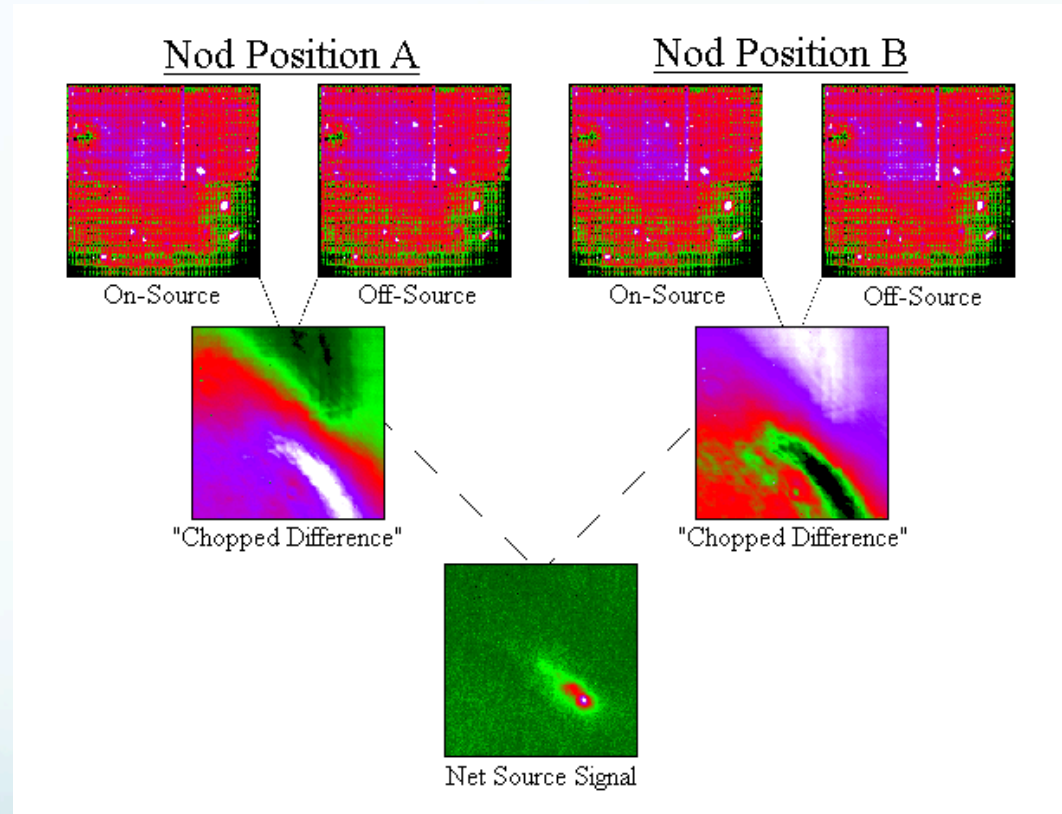
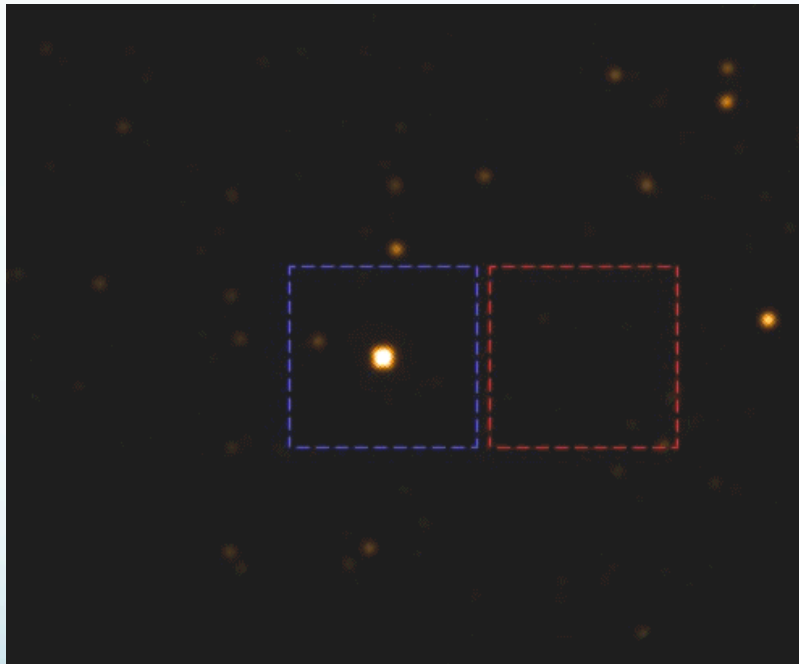


# Chop Nod/ Scan Mapping





# Chop Nod/ Scan Mapping

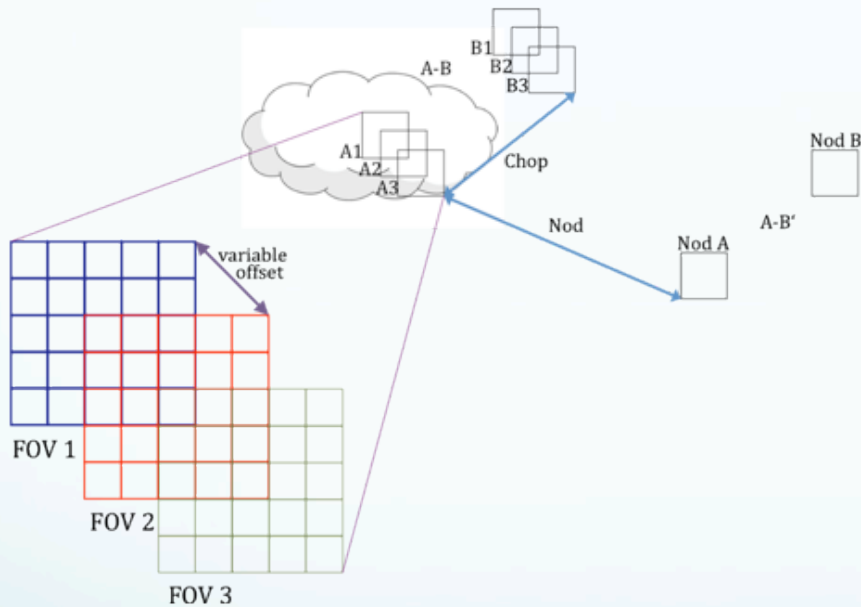




# Chop Nod/ Scan Mapping

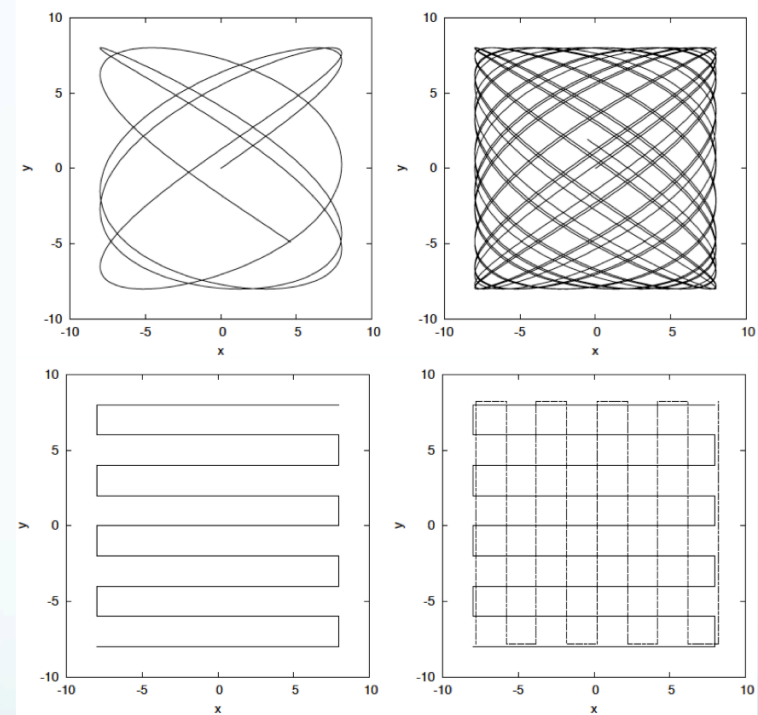


## Asymmetric Chop Nod



FIFI-LS

## Scan Mapping



HAWC+





# Unique Attributes

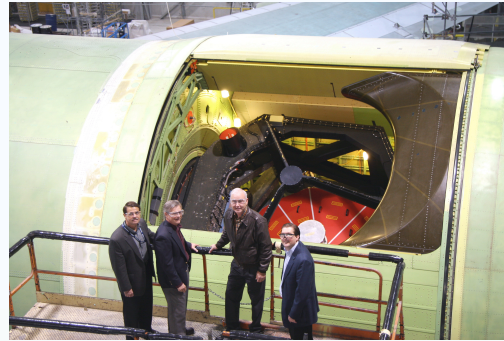


- Alt Az Telescope restrictions
  - LOS (Line of Sight) Rewinds
- Overhead uncertainty
- Flight Plans
- Calibration Differences

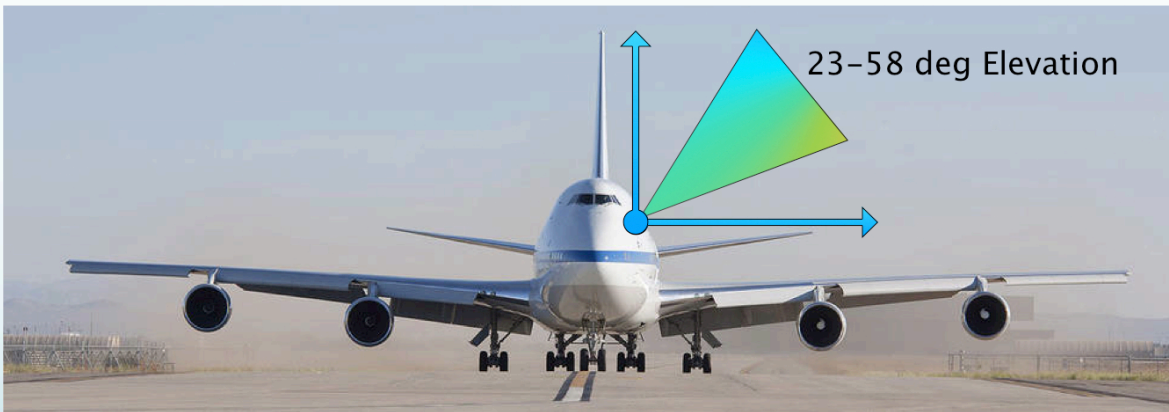




# Telescope Altitude



- About ~2.5m (~Hubble sized)
- Telescope points out one side of craft
- Limited altitude range (~35 degrees)







# Azimuth (LOS Rewinds)

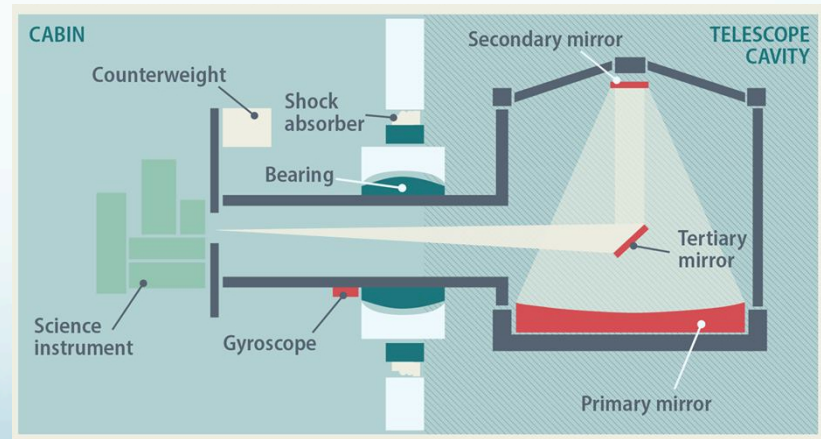
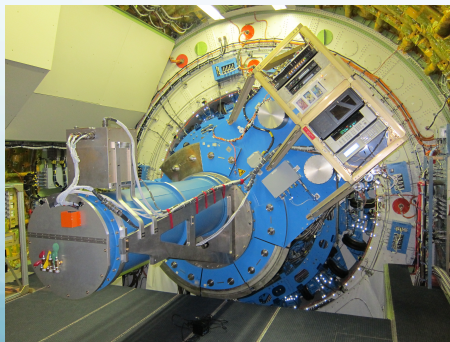
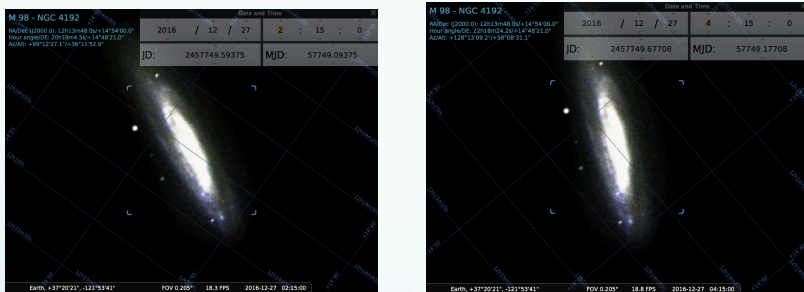


## 2hr observation March 2016



- No Rotator
- Limited to ~6 degrees
- Rotation rate based on location of object in sky
- Correction needed to prevent image smearing

## 2hr observation Dec 2016

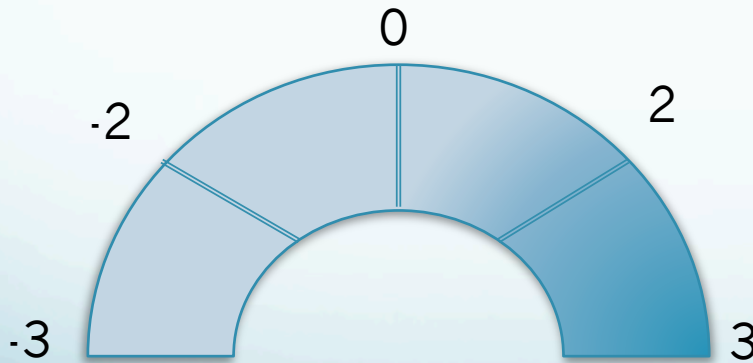
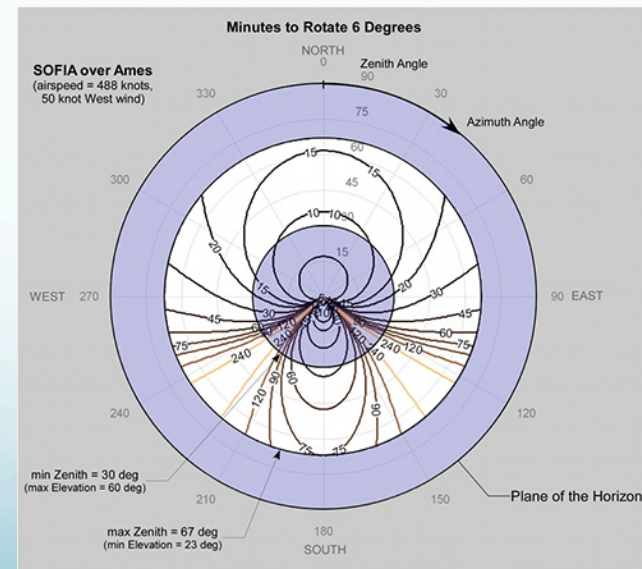
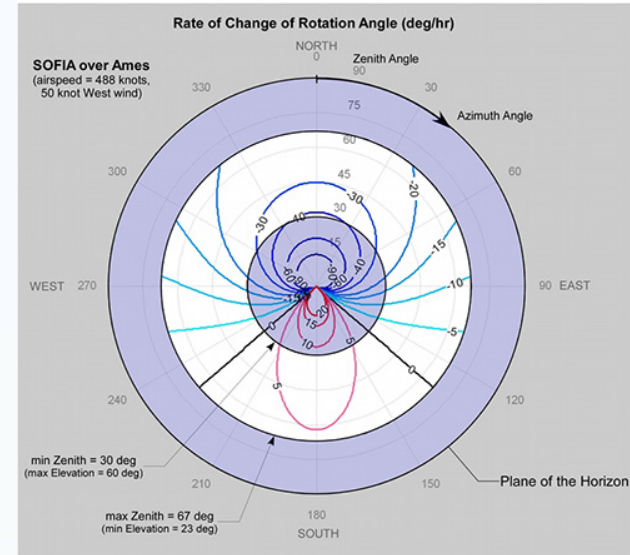




# LOS Rewinds



- Require overhead based on speed of rotation
- Not known till flight plan
- Full range 6 degrees, but keep with 3-4 degrees
- SI responsibility (see also Visibility tool (<https://dcs.sofia.usra.edu/observationPlanning/installVT/>))



<http://www.sofia.usra.edu/Science/ObserversHandbook/Cy2.html#LOS>

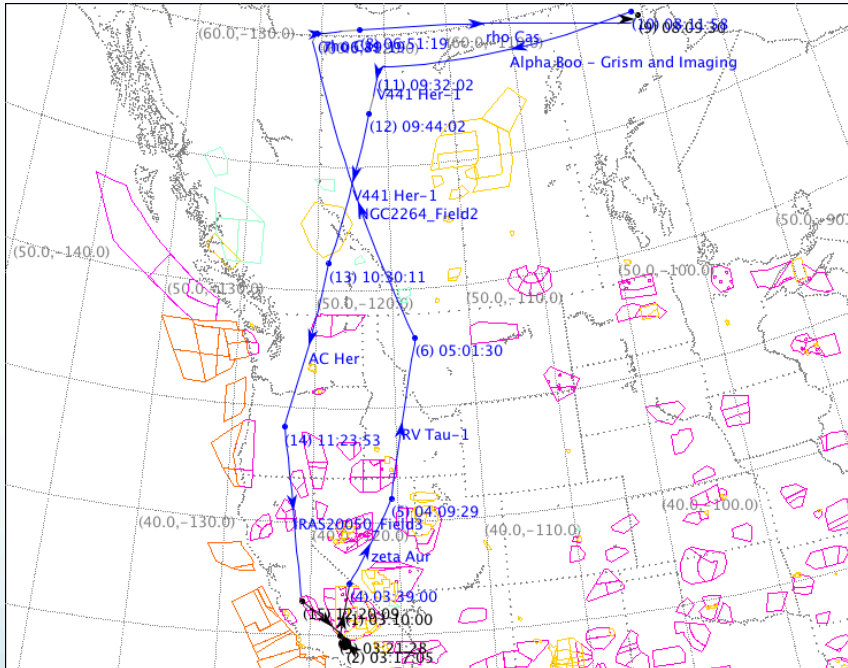




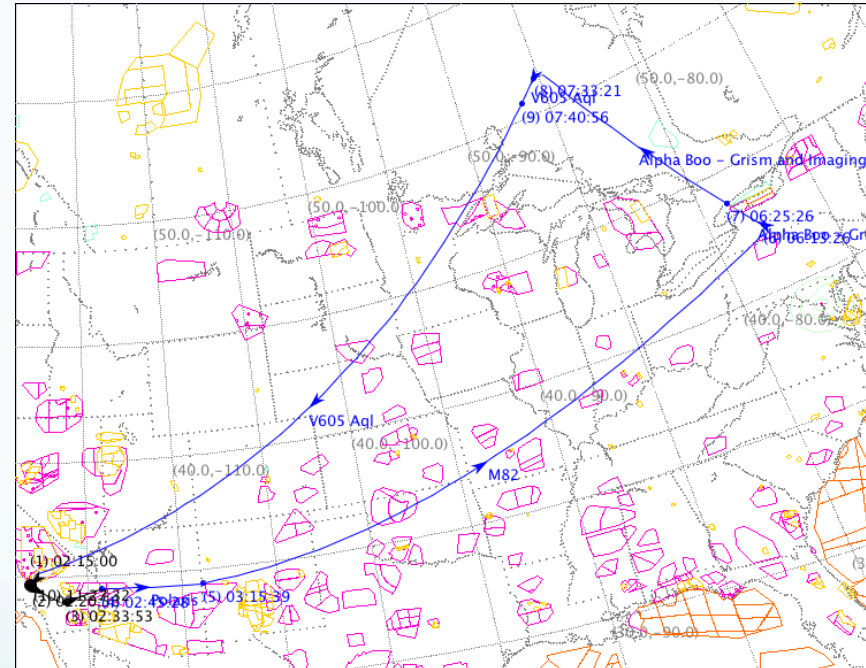
# Flight Plans



- Quantized Observing (limits to how short and how long)
- No real flexibility
- Complicated to plan



Flight Plan Name: File: 201403\_FO\_04\_post-MOpsV2.fp  
 Flight ID: 2014/03/27  
 Est. Takeoff Time: 2014-Mar-27 03:10 UTC  
 Est. Landing Time: 2014-Mar-27 12:48 UTC  
 Flight Duration: 09:38  
 Weather Forecast : 1200 Thu Feb 20 2014 - 0000 Sun Feb 23 2014 UTC  
 Saved: 2014-Feb-20 19:03 UTC User: kbower



Flight Plan Name: File: 201404\_FO\_05a\_post-Science.fp  
 Flight ID: 2014/05/07\_1  
 Est. Takeoff Time: 2014-May-07 02:15 UTC  
 Est. Landing Time: 2014-May-07 11:54 UTC  
 Flight Duration: 09:39  
 Weather Forecast : 1800 Tue Apr 01 2014 - 0600 Fri Apr 04 2014 UTC  
 Saved: 2014-Apr-05 23:31 UTC User: kbower





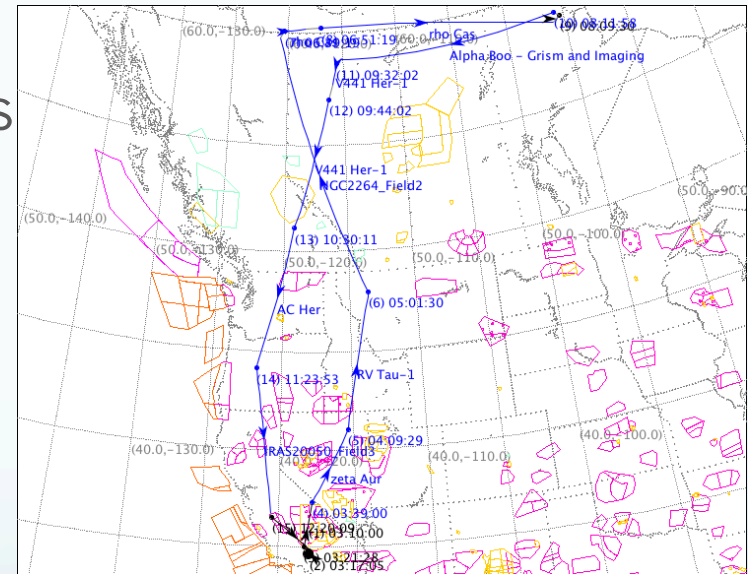




# Calibration



- Different by instrument (e.g. EXES vs FORCAST)
- FORCAST
  - Calibration accuracy  $\pm 20\%$
  - PSF calibration difficult
  - Limited number of bright stars
  - Averages done for flight series

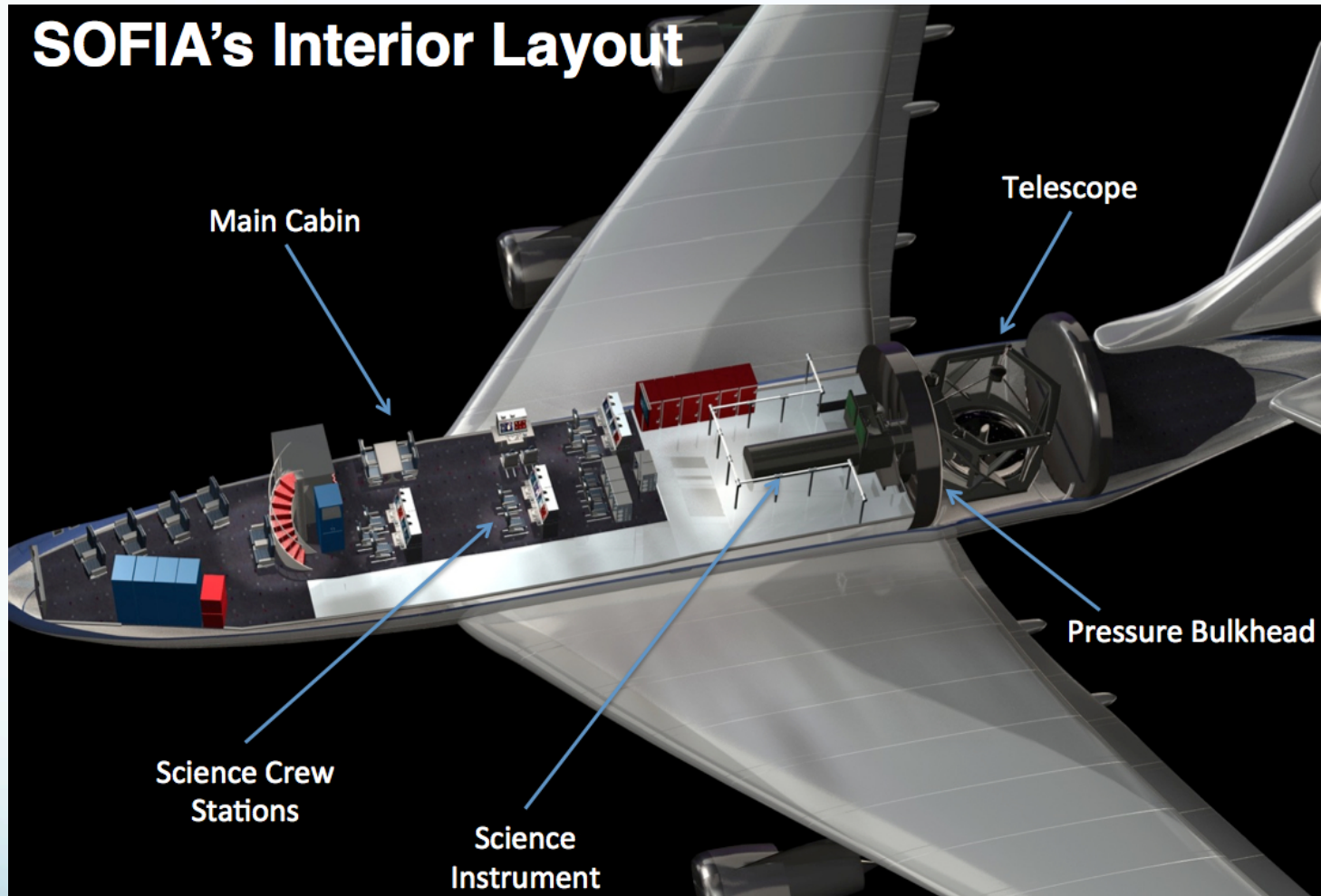


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# SOFIA interior







# SOFIA interior



Image Credit: Sky and Telescope



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# 15s SOFIA Flight



Video credit: Dr. Daniel Angerhausen, Rensselaer Polytechnic Institute  
<https://plus.google.com/+NASA/posts/dpVqr3DDBrq>



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