





<p><b>Expanded Very Large Array (VLA)</b> (cm-wavelength interferometer)</p>	<p><b>Atacama Large Millimeter Array (ALMA)</b> (mm/sub-mm wavelength interferometer)</p>
	
<p><b>Star formation, Magnetic fields, High energy astrophysics</b></p>	<p><b>Galaxy formation and evolution, Star and Planet Formation, Solar system</b></p>
<p><b>NRAO and the EVLA</b></p>	
<p><b>Green Bank Telescope (GBT)</b> (filled aperture)</p>	<p><b>Very Long Baseline Array (VLBA)</b> (trans-continental array)</p>
	
<p><b>Interstellar medium, Astrochemical basis of life, Extreme physics</b></p>	<p><b>Active galactic nuclei (AGN), Black hole and jet physics, Precision astrometry</b></p>



## Two Key Questions

- What should NRAO facilities be doing now in collaboration with NASA's Great Observatories?
- What should NASA's Great Observatories be doing now to support science with NRAO's future facilities (EVLA and ALMA)?



## The Era of the Great Observatories

- May 17<sup>th</sup>, NRAO “Legacy Projects Workshop”
  - 75 participants. Discussed science *and* policy
  - <http://www.aoc.nrao.edu/events/legacy/>
- NRAO must increase fraction of time for Large Proposals on all its facilities (**Next call: August 2006**)
  - Large Projects defined to be  $\geq 200$  hr of observing (i.e. 720 ksec, or  $\sim 133$  HST orbits)
  - Previously allocated 10%-20% to Large Proposals
    - VLA and VLBA observe 6000 and 4500 hr/yr, respectively
  - **New:** limit to 25-50% depending on proposal pressure
    - No more than 50% of observing time at one Local Sidereal Time in one VLA configuration
    - Big change. Social engineering to remove perceived bias.

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## Past Large Proposals

AS801	<a href="#">VLA COSMOS</a>	VLA A,C	2004-2005	264 hr	E. Schinnerer
AK563	<a href="#">Virgo: A Laboratory for Studying Galaxy Evolution</a>	VLA C	2004-2005	240 hr	J. Kenney
BL123	<a href="#">MOJAVE: Monitoring of Jets in Active galaxies with VLBA Experiments</a>	VLBA	2004-2005	14 x 24 hr	M. Lister
AW605	<a href="#">THINGS: The HI Nearby Galaxy Survey</a>	VLA B,C,D	2003-2005	293 hr	F. Walter
AH810	<a href="#">Coordinated Radio and Infrared Survey for High-Mass Star Formation</a>	VLA B	2005	40 hr pilot	M. Hoare
BL111	<a href="#">MOJAVE: Monitoring of Jets in Active galaxies with VLBA Experiments</a>	VLBA	2002-2004	17 x 24 hr	M. Lister
AK509	<a href="#">Cosmic Explosions</a>	VLA	2000-2003	30 hr/month	S. Kulkarni
BC120	<a href="#">Pulsar Astrometry with the VLBA</a>	VLBA	2002-2004	300 hr	S. Chatterjee
AS687	<a href="#">A Deep Radio Survey of the SIRTIF First-look Survey</a>	VLA B	2001-2002	240 hr	T. Soifer
AB628, AB879, AB950	<a href="#">FIRST Survey</a>	VLA B	1993-2002	3209 hr	R. Becker
AG592	<a href="#">HI Survey of Clusters in the Local Universe</a>	VLA C	2001-2002	360 hr	J. van Gorkom
AP397	<a href="#">A 4-meter All-sky Survey</a>	VLA BnA,B	2001	70 hr pilot	R. Perley
AT245	<a href="#">A Global, High Resolution HI Survey of the Milky Way</a>	VLA D	2000	260 hr	R. Taylor
C308	<a href="#">NRAO VLA Sky Survey</a>	VLA D, DnC	1993-1996	2939 hr	J. Condon



## Current Large Proposals

(see [www.vla.nrao.edu/astro/](http://www.vla.nrao.edu/astro/))

AH884	<a href="#">The Coordinated Radio and Infrared Survey for High-Mass Star Formation (The CORNISH Survey)</a>	VLA B, BnA	2006-2007	360 hr	M. Hoare
BL137	<a href="#">MOJAVE II: Monitoring of Jets in Active galaxies with VLBA Experiments II. Entering the GLAST Era</a>	VLBA	2006-2007	384 hr	M. Lister
BT085	<a href="#">The VLBA Imaging and Polarimetry Survey (VIPS)</a>	VLBA	2006	195 hr	G. Taylor
BR100	<a href="#">The Spiral Structure and Kinematics of the Milky Way</a>	VLBA	2005-2007	270 hr	M. Reid
AK583	<a href="#">Cosmic Explosions</a>	VLA	2005-2006	20 hr/month	S. Kulkarni
AP452	<a href="#">VLA Low-frequency Sky Survey</a>	VLA BnA, B	2003-2006	690 hr	R. Perley



## VLA "Blank Field" Proposals

- Best telescope for deep radio integrations of extragalactic fields
- Crucial for localizing the IR/submillimeter galaxy population

AM857	A Deep & Unbiased Probe of Star Formation in the GOODS Northern Field	VLA A	2006	77 hr	G. Morrison
AO201	The SWIRE Deep Field at 90cm: A Steep Spectrum MicroJy Radio Population?	VLA A,C	2006	66 hr	F. Owen
AS859	Follow-up of the COSMOS 1.4 GHz Imaging Survey: Identification of Dusty Massive Starforming Systems	VLA A	2006	60 hr	E. Schinnerer
AY164	An In-depth Investigation of the Nature of the Faint 24 Micron Spitzer Sources and 1100 Micron AzTEC Sources in the FLS Verification Strip	VLA A,B	2006	96 hr	M. Yun



## Joint NRAO/GO Projects Now

Science Question	Project	Telescope(s)
Galaxy Evolution	Red-shifted molecular lines	GBT
Galaxy Evolution	More continuum surveys	VLA
Galaxy/BH connection	Imaging jets at $z > 2$	VLA, VLBA
Early Universe	High $z$ GRBs	VLA
Galactic Transients	SGRs, SNe, AXPs, BDs, XRBs	VLA, GBT, VLBA
Our Galaxy	Atomic & molecular gas, Ionized gas, B-fields	GBT & VLA

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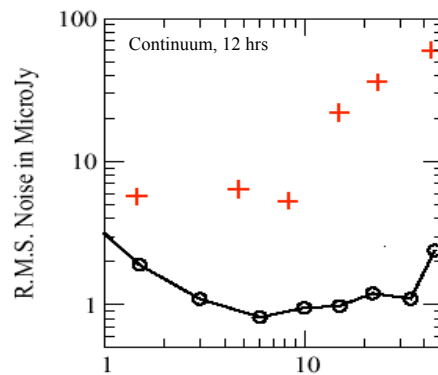
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## The EVLA Project

- 10x increase in continuum sensitivity
- Continuous frequency coverage from 1 to 50 GHz
- Designed for easy access by all astronomers
- NSF/AST (\$57M) + North American partners
  - Finished in 2012
  - First science 2009-2010



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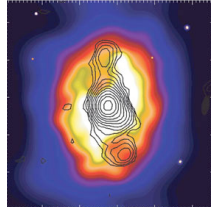
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## EVLA Design Driven By Four Themes

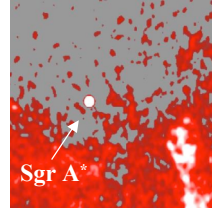
### Magnetic Universe

Measure the strength and topology of the cosmic magnetic field.



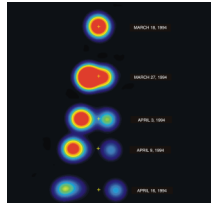
### Obscured Universe

Image young stars and massive black holes in dust enshrouded environments.



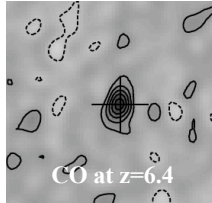
### Transient Universe

Follow the rapid evolution of energetic phenomena.



### Evolving Universe

Study the formation and evolution of stars, galaxies and AGN.



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## What should the GO's be doing now?

Science Question	Project	Telescope(s)
Our Galaxy	Complete census of nearby SF regions	Chandra, Spitzer
Our Galaxy	More galactic surveys	Spitzer
Globular clusters	Characterize compact object population	All
Galactic center	Stellar census and time domain	Spitzer, Chandra
Nearby galaxies	ISM, stellar and compact objects (ULXS and SNe)	All
Galaxy/BH connection	Compton thick AGN	Chandra
Extragalactic jets	Survey of flux-limited sample (3C) of radio jets	Chandra
Galaxy Clusters	Imaging all large clusters	Chandra



Fin pour nu