A Spitzer warm mission Ultra-Wide Survey as a target finder for the James Webb Space Telescope



James Webb Space Telescope



- 6.6m Telescope
- Launch in 2013 to L2 on Ariane 5.
- Successor to Hubble & Spitzer.
- Imaging and spectroscopy over 0.6 to 28.5 microns
 - NIRCam (U AZ)
 - NIRSpec (ESA)
 - MIRI (JPL+ESA)
 - TFI (CSA)
- All technology at TRL-6
- Lead: Goddard Space Flight Center
- Prime: Northrop Grumman Space Technology
- Operations: STScl
- Senior Scientist: Nobel Laureate John Mather



see: Space Science Reviews, 2006, 123/4, 485 or astro-ph/0606175

End of the dark ages: first light and reionization

- What are the first galaxies?
- When did reionization occur?
 Once or twice?
- What sources caused reionization?





Hubble Ultra Deep Field Hubble Space Telescope • Advanced Camera for Surveys

- Ultra-Deep NIR survey (1.8 nJy), spectroscopic & Mid-IR confirmation.
- QSO spectra: Ly-α forest
- Galaxy spectra: Balmer lines (2x10⁻¹⁹ ergs/cm²/sec)

Reionization



End of reionization is seen at z~6 in SDSS quasars.

Spectra of high-z quasars showing Gunn-Peterson trough (Fan et al. 2007)

~

8800 7000 7200	7400	7600	7800	8000	∧ (A 8200	8400	8600	8800	9000	9200	9400	9600	
J1148+5251 z	=6.42	الملاب						·		سالسد			-
J1030+0524 z	=6.28						+	N			····	····	+
J1623+3112 z	=6.22	,	<u> </u>	مى ب ەرمەر				-	····	،	•t	···· 1 · · ·	+
J1048+4637 z	=6.20	-	<u></u>			 		N	····	, .	~~~~	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	7
J1250+3130 z	=6.13			• • •								****	+
J1602+4228 z	=6.07		-	• • • •		+ +	- Am		·	~			-
J1630+4012 z	=6.05			• • •			~~	~~~	v~~~		20.22		÷
J1137+3549 z	=6.01					N	/					+ + +	-
J0818+1722 z	=6.00		man			- and			····		m	~~~	~
J1306+0356 z	=5.99						m		~~~~	····	····		+
J1335+3533 z	=5.95	~~~~	~~~~	• • •	1	, m	- the	-	Angeles		-	yman	Ŵ
J1411+1217 z	=5.93	()				m	····	+	····	+	····	····	+
J0840+5624 z	=5.85		Anna A			mm	مرك	-	hour	www.	nte Mile	viden	Ī,
J0005-0006 z	=5.85					A	·l	+··· • • • · ·	,		b	*	-
J1436+5007 z	=5.83					Ann			Alama	+	Jane A	1 march	4
J0836+0054 z	=5.82			• • •	No.		+	+ *	+	• • • • • • •		+ +	+
J0002+2550 z	=5.80			• • • •	~	~~~~~	~~~	+			····		+-
J0927+2001 z	=5.79	(alah da	have a	M	-					wheth	-	÷
J1044-0125 z	=5.74				m	~~~	****	·····		how	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	v	_
6800 7000 7200	7400	7600	7800	8000	8200	8400	8600	8800	9000	9200	9400	9600	



Universe is partially ionized with fraction x_e^0 at redshift z_{reion} , and fully reionized at z=7. (Spergel et al. 2007)

• WMAP polarization shows reionization started earlier.

• Complex history?

High Redshift Quasars





Evolution of quasar number density.

LF of quasars at $z \sim 6$. A break is expected at low luminosities, but not yet observed.



How well would a Spitzer survey do?

Compare two surveys:

- 3 µJy, 500 sq. degrees
- 10x exposure over 50 sq. degrees

Area, depth	z > 6	z > 8	z > 10
500 sq. deg., 3 µJy	130	22	0.8
50 sq. deg., 1 μJy	120	20	0.7

- Quasars in the shallow survey would typically be brighter, better for spectroscopic studies of reionization with JWST.
- Identification is non-trivial, need deep optical and/or NIR.